

Growth of Spread in Convection-allowing and Convection-parameterizing Ensembles

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Abstract

During the 2007 NOAA/Hazardous Weather Testbed Spring Experiment, a 10-member 4-km grid-spacing Storm-Scale Ensemble Forecast (SSEF) system was run in real-time to provide experimental severe weather forecasting guidance. Five SSEF system members used perturbed initial and lateral boundary conditions (ICs/LBCs) and mixed-physics (ENS4), and five members used only mixed-physics (ENS4^{phys}). This ensemble configuration facilitates a comparison of ensemble spread generated by a combination of perturbed ICs/LBCs and mixed-physics to that generated by only mixed-physics, which is examined herein. In addition, spread growth and spread-error metrics for the two SSEF system configurations are compared to similarly configured 20-km grid-spacing convection-parameterizing ensembles (ENS20 and ENS20^{phys}). 12 forecast fields are examined for 20 cases.

For most fields, ENS4 mean spread growth rates are higher than ENS20 for ensemble configurations with both sets of perturbations, which is expected as smaller scales of motion are resolved at higher resolution. However, when ensembles with only mixed-physics are compared, mass-related fields (i.e., geopotential height and mean-sea-level pressure) in ENS20^{phys} have slightly higher spread growth rates than ENS4^{phys}, likely resulting from the additional physics uncertainty in ENS20^{phys} from varied cumulus parameterizations which were not used at 4-km grid-spacing. For 4- and 20-km configurations, the proportion of spread generated by mixed-physics in ENS4 and ENS20 increased with increasing forecast lead time. In addition, low-level fields (e.g., 2-m temperature) had a higher proportion of spread generated by mixed-physics than mass-related fields. Finally, spread-error analyses revealed that ensemble variance from the current uncalibrated ensemble systems was not a reliable indicator of forecast uncertainty. Furthermore, ENS4 had better statistical consistency than ENS20 for some mass-related fields, wind-related fields, precipitation, and most unstable convective available potential energy (MUCAPE) with no noticeable differences for low-level temperature and dewpoint fields. The variety of results obtained for the different types of fields examined suggests that

future ensemble design should give careful consideration to the specific types of forecasts desired by the user.

1. Introduction

To sufficiently account for model and observational errors so that all possible states of the future atmosphere are simulated, perturbation strategies for recent short-range ensemble forecast (SREF) systems include: 1) perturbing the initial conditions (ICs; e.g., Toth and Kalnay 1997; Palmer et al. 1992; Molteni et al. 1996), 2) using different combinations of physical parameterizations (mixed-physics; e.g., Houtekamer et al. 1996; Stensrud et al. 2000; Du et al. 2004; Jones et al. 2007), and 3) using different numerical models (e.g., Hou et al. 2001; Wandishin et al. 2001; Du et al. 2004; Eckel and Mass 2005; Jones et al. 2007). In current SREF systems, sensible parameters influenced by small-scale processes that must be parameterized [e.g., planetary boundary layer (PBL) temperature/moisture and convective precipitation] are associated with notably underdispersive forecasts (Fritsch and Carbone 2004; Eckel and Mass 2005). The error growth for these sensible parameters typically contains a much larger contribution from model uncertainty relative to IC uncertainty than for synoptic-scale parameters (e.g., 500-hPa geopotential heights and winds, mean-sea-level pressure; Stensrud et al. 2000; Eckel and Mass 2005). The underdispersion may be a result of several deficiencies including 1) inadequate methods to account for model error, 2) inability to capture small-scale variability because of insufficient resolution (Eckel and Mass 2005), 3) coarsely resolved and temporally interpolated lateral boundary conditions (LBCs; Nutter et al. 2003), and 4) inadequate sampling of the most important growth directions by the limited-size ensemble.

One method commonly used to gain information on ensemble spread is to isolate the error sources by using different perturbation strategies for a set of forecasts (e.g., Houtekamer et al. 1996; Stensrud et al. 2000; Clark et al. 2008). For example, to isolate model errors, the “perfect analysis” assumption can be used, in which identical sets of ICs/LBCs are used to initialize various ensemble members with mixed-physics. Similarly, to isolate IC errors, the “perfect model” assumption can be used in which identically configured ensemble members are initialized with different sets of perturbed

ICs. During the 2007 NOAA/Hazardous Weather Testbed (HWT) Spring Experiment (SE07; Xue et al. 2007; Kong et al. 2007), a 10-member, 4-km grid-spacing Storm-Scale Ensemble Forecast (SSEF) system was run in real-time to provide severe weather forecasting guidance to the SE07 participants. Five of the SSEF members used perturbed ICs/LBCs and mixed-physics (ENS4; four perturbed members and one control member), while five members used only mixed-physics (ENS4^{phys}) so that the impacts of the different physical parameterization schemes could be isolated. This configuration of the 2007 SSEF system also facilitates an isolation of physics-related model errors because five members use the “perfect analysis” assumption. Unfortunately, because there were not any ensemble subsets with only IC perturbations, the “perfect model” assumption could not be assessed.

The goal of this paper will be to use the 2007 SSEF system to compare ensemble spread from a mixed-physics only ensemble to an ensemble with both mixed-physics and perturbed IC/LBCs for various fields in a convection-allowing ensemble. In addition, ensemble spread growth and spread-error relationships associated with the two 5-member subsets of the SSEF system will be compared to two similarly configured subsets of a 20-km grid-spacing convection-parameterizing ensemble to examine the impacts of horizontal resolution for various forecast fields. This paper is organized as follows: in section 2, a description of the data and methodology is provided, and in section 3 the results are examined. A summary and discussion are provided in section 4.

2. Data and Methodology

The 2007 SSEF system was run during April-June 2007 and used the WRF-ARW (Version 2.2.0; Skamarock et al. 2005) model. The 10 SSEF members were run by the Center for Analysis and Prediction of Storms (CAPS) of the University of Oklahoma, initialized daily at 2100 UTC, and integrated 33 hours over an approximately 3000 x 2500 km domain covering about two thirds of the continental United States (Fig. 1). SSEF system ensemble member specifications for ENS4 and

ENS4^{phys} are listed in Tables 1 and 2, respectively. Note, because of the 2100 UTC initialization, forecast hours 3 and 27 correspond to 0000 UTC. For the SSEF control member, the 2100 UTC analyses from NCEP's operational North American Mesoscale (NAM; Janjic 2003) model (at 12-km grid-spacing) were used for ICs and the 1800 UTC NAM 12-km forecasts were used for LBCs. For the members with perturbed ICs, perturbations were extracted from the 2100 UTC NCEP SREF system ICs (Du et al. 2004) and added to the 2100 UTC NAM analyses. Corresponding NCEP SREF system forecasts were used for LBCs (3-hr updates). Xue et al. (2007) and Kong et al. (2007) provide more details on the configurations.

For a comparison of the 5-member SSEF ensemble subsets to a similarly configured convection-parameterizing ensemble, a 30-member 20-km grid-spacing ensemble was generated at Iowa State University, which was also composed of WRF-ARW (Version 2.2.0) members. Fifteen of the 20-km members have mixed-physics and perturbed ICs/LBCs (ENS20), while another 15 members have only mixed-physics (ENS20^{phys}). Ensemble member specifications for ENS20 and ENS20^{phys} are provided in Tables 3 and 4, respectively. It should be noted that the ENS20 and ENS20^{phys} ensembles have one more set of varied physics schemes relative to ENS4 and ENS4^{phys}, because in addition to the different planetary boundary layer, microphysics, and surface layer schemes, both 20-km ensembles use different cumulus parameterizations (CPs; the 4-km ensembles do not use CP). However, 5-member subsets of ENS20 and ENS20^{phys} that use the same CP do have the same number of varied physics schemes as ENS4 and ENS4^{phys}. These five member subsets will be referred to as ENS20_{cp} and ENS20^{phys}_{cp} where CP refers to one of the three different CPs used: 1) Kain-Fritsch (KF; Kain and Fritsch 1993), 2) Betts-Miller-Janjic (BMJ; Betts 1986; Betts and Miller 1986; Janjic 1994), and 3) Grell-Devenyi (GD; Grell and Devenyi 2002). For the 20-km ensemble members, different sets of ICs and corresponding LBCs for each member are obtained directly from NCEP SREF members initialized at 2100 UTC.

Both ensembles use the RRTM short-wave radiation scheme (Mlawer et al. 1997) and Goddard

long-wave radiation scheme (Chou and Suarez 1994), along with the Noah land surface model (Ek et al. 2003). Varied PBL schemes include Mellor-Yamada-Janjic (MYJ; Mellor and Yamada 1982; Janjic 2002) and Yonsei University (YSU; Noh et al. 2003) schemes. Varied microphysics schemes include Thompson et al. (2004), WRF single-moment 6-class (WSM-6; Hong and Lim 2006), and Ferrier et al. (2002), and surface layer schemes include Monin-Obukhov (Monin and Obukhov 1954; Paulson 1970; Dyer and Hicks 1970; Webb 1970) and Janjic Eta (Janjic 1996, 2002). Note that none of the ensemble members use positive-definite advection of moisture, which may have contributed to high biases in the precipitation forecasts (e.g., Skamarock and Weisman 2008).

The forecasts were examined for 20 cases during April-June 2007 (Fig. 2). Note, these 20 cases are included in the 23 cases in which ENS4 and ENS20 precipitation forecasts were compared in Clark et al. (2009). The other three cases examined in Clark et al. (2009) are excluded from the current study because some of the ENS4^{phys} members were not available. As noted by Clark et al. (2009), the period examined was relatively active with a variety of convective precipitation events.

This study examines growth of spread (i.e., ensemble variance) and statistical consistency [i.e., correspondence between ensemble variance and mean-square-error of the ensemble mean (MSE)] for 12 fields: 500-, 700-, and 850-hPa geopotential height (500Z, 700Z, and 850Z, respectively), mean-sea-level pressure (MSLP), 2-meter temperature (T2), 2-meter dewpoint (Td2), 850-hPa wind magnitude (850WMAG), 850-hPa temperature (850T) and dewpoint (850Td), 3-hrly accumulated precipitation (PREC), most unstable convective available potential energy (MUCAPE), and magnitude of the 10-meter to 500-hPa shear vector (WSHR). The 12 fields examined are separated into those that are “mass-related”, or heavily dependent on the properties of the atmosphere within a vertical column (500Z, 700Z, 850Z and MSLP), and “low-level” fields that have more dependence on boundary layer processes and, thus, have a noticeable diurnal signal (T2, Td2, 850WMAG, 850T, 850Td, PREC, MUCAPE, and WSHR). In the subsequent analyses of ensemble variance and MSE, comparisons are

made between ensemble subsets that have different numbers of members. Thus, it should be kept in mind that the forecast probability distribution function (PDF) sampled by the smaller ENS4 membership should be less representative of the forecast PDF than the larger ENS20 membership, which would imply a less “certain” estimate of the ensemble variance in ENS4.

For the computation of MSE and application of a bias-correction procedure in Section 3d, operational 20-km grid-spacing Rapid Update Cycle (RUC) model analyses provided by NCEP and available at 1-hrly temporal resolution are used as “truth” for non-precipitation fields. The RUC analyses are generated using hourly intermittent 3DVAR assimilation cycles in which recent observations from various sources (e.g., wind profiler, radar, aircraft, surface METARS, satellite, etc.) are assimilated using the previous 1-hr RUC model forecasts as the background field. Additional information on the RUC model is found in Benjamin et al. (2004a, and b). For precipitation fields, stage IV multi-sensor rainfall estimates (Baldwin and Mitchell 1997) are used. Ensemble variance and MSE are computed according to Eqs. B7 and B6, respectively, in Eckel and Mass (2005), which are designed to account for an ensemble with a finite number of members. Finally, to obtain a more equitable comparison between the 4-km and 20-km forecast fields, the 4-km fields were remapped to a 20-km grid covering the central US, which is just a subdomain of the ENS20 members, using a neighbor-budget interpolation (e.g., Accadia et al. 2003).

3. Results

a. Spread growth

1) ENSEMBLE VARIANCE TIME SERIES

To illustrate the temporal evolution of spread growth during the 33-hour forecast period, time series of average ensemble variance for all 12 fields at 3-hrly intervals with box-plots overlaid to show variability are displayed in Fig. 3. In order to compare 4-km and 20-km ensembles with the same types

of perturbations, each panel in Fig. 3 displays ensemble variance for ENS4 and ENS20, or ENS4^{phys} and ENS20^{phys}. Note that the different y-axis scales in Fig. 3 do not allow an easy comparison of growth rates between ensembles with both IC/LBC perturbations and mixed-physics and those with only mixed-physics (Phys) for each field; these comparisons are made in the next section. A number of distinct features can be seen in these time series. First, for the mass-related fields (Figs. 3a-h), ENS4 and ENS20 (Figs. 3a, c, e, and g) have a generally linear increase in mean spread and it appears that ENS4 spread is increasing at a faster rate than ENS20, which is a generally expected behavior because the smaller scales being resolved in ENS4 should be associated with faster perturbation growth that feeds back to the larger scales (e.g., Lorenz 1969). Also, the biggest differences between ENS4 and ENS20 at each time occur at the higher ends of the variance distributions, i.e. differences in the upper part of the box-plot ranges are greater than differences in the lower part, indicating that the distributions are more right-skewed in ENS4 relative to ENS20. For the mass-related fields in ENS4^{phys} and ENS20^{phys} (Figs. 3b, d, f, and h), the ENS20^{phys} mean spread increases at a faster rate than in ENS4^{phys}, with the exception of mean MSLP spread (Fig. 3h) which appear to be similar. In addition, the spread increase in ENS4^{phys} and ENS20^{phys} is not linear as it was for ENS4 and ENS20, but instead has a ~ 6 -hr period during forecast hours 21-27 (1800-0000 UTC) during which spread increases at a noticeably faster rate than the other times. This 6-hr period corresponds to when peak insolation occurs and likely corresponds to when the different physics parameterizations are most active and thus result in the most spread increase. For example, in the central US, the boundary layer typically reaches its maximum depth by early afternoon so that turbulent processes that must be parameterized are occurring over a relatively deep layer. In addition, peak heating and resulting well mixed boundary layers also lead to shallow and deep convective clouds requiring microphysics and cumulus (only for ENS20 and ENS20^{phys}) parameterizations to be more active relative to other times.

For the low-level fields (Figs. 3i-x), ENS4 and ENS20 mean variances (Figs. 3i, k, m, o, q, s, u,

and w) have clear diurnal signals superimposed on increasing trends. The differences in mean variances between ENS4 and ENS20 vary among the different fields analyzed. For example, 850WMAG variances in ENS4 and ENS20 are very similar over the entire forecast period (Fig. 3m); Td2 variances and variance growth rates are higher in ENS4 relative to ENS20 for most of the forecast period (Fig. 3k); and WSHR variances are similar until forecast hour 21, when there is a marked increase in ENS4 variances relative to ENS20 (Fig. 3w). The amplitude and phase of the diurnal signal also vary among the different fields. For example, 850T and 850Td ENS4 and ENS20 variances (Figs. 3o and q, respectively) have smaller amplitudes relative to the other low-level fields, and peak variances occur for 850WMAG around 0900-1500 UTC (Fig. 3m), for PREC at 0600-0900 UTC (Fig. 3s), and for MUCAPE at 2100-0000 UTC (Fig. 3u). The peak variances tend to match the time at which the forecasts of the variable considered are maximized. To illustrate this behavior, time series of mean domain averaged T2 from all cases for ENS4 and ENS20 members and ensemble means, as well as RUC analyses are displayed in Fig. 4. The peak T2 values occur around forecast hour 25 when there also appears to be the most spread in domain averaged T2 among the ensemble members. Also worth noting in Fig. 4 is that all ENS4 members are cooler than the RUC analyses when T2 peaks, however, ENS20 has about equal numbers of members with warm and cool biases resulting in a mean that is very close to the RUC analysis. Coniglio et al. (2009) found similar cool biases for mean 2-m temperatures in convection-allowing WRF model simulations run over the central US during spring 2008

2) VARIANCE GROWTH RATES

It was possible to subjectively infer differences in mean ensemble variance growth rates from the analysis conducted in Fig. 3; however, to compare ENS4 and ENS20 to ENS4^{phys} and ENS20^{phys} and better quantify mean variance growth rates, a simple objective method was developed using a standard formula for growth rate:

$$(\text{Var}_f - \text{Var}_i) / \text{Var}_i * 100\%, \quad (1)$$

where Var_i and Var_f are initial and final mean variance, respectively. To reduce the impact of the diurnal cycle signal on the variance growth rates, mean variances at forecast hours 9 and 33 are used as the initial and final values, respectively, because these forecast hours are separated by 24 hours or one complete diurnal cycle. Also, to smooth out some of the higher frequency variability in the mean variance time series, the `lowess()` function in the R statistical software package (R Development Core Team 2007), which uses locally-weighted polynomial regression, was used as a low-pass filter. For mass-related and low-level fields, the proportion of points influencing the filtered value at each forecast hour, or the “smoother span”, was set to 0.25 and 0.1667, respectively (i.e., ~ 8 and 6 points). A slightly larger smoother span was used for the mass-related fields than for low-level fields to adequately filter particularly high frequency variability during the first 12 hours of the forecasts for the mass-related fields. Finally, the mean variances at all forecast hours from $\text{ENS4}^{\text{phys}}$, ENS20 , $\text{ENS20}^{\text{phys}}$, and $\text{ENS20}^{\text{phys}}_{\text{cp}}$ subsets were adjusted by the difference between their variance at forecast hour 9 and that from ENS4 (i.e., the entire time series was shifted by a constant). Thus, the variance growth for all ensemble subsets was computed relative to the same initial mean variance (i.e. the ENS4 mean variance) to allow for comparison between ensemble subsets. The adjusted variance can be expressed as:

$$\sigma_{\text{ENS}}^2 * = \sigma_{\text{ENS}}^2 + [\sigma_{\text{ENS}(\text{fhr}9)}^2 - \sigma_{\text{ENS4}(\text{fhr}9)}^2] \quad (2)$$

where, $\sigma_{\text{ENS}}^2 *$ is the adjusted variance at any forecast hour for a specified ensemble subset; σ_{ENS}^2 and $\sigma_{\text{ENS}(\text{fhr}9)}^2$ are the unadjusted variance for the specified ensemble subset at any forecast hour and at

forecast hour 9, respectively; and $\sigma^2_{\text{ENS4}(\text{thr9})}$ is the unadjusted variance at forecast hour 9 for the ENS4 ensemble. This adjustment step was important because the growth rates are sensitive to the initial variance which varies according to the field and ensemble subset. Using this procedure to compute spread growth rate, a 100% growth rate for any field can be interpreted as a 24-hr doubling of spread relative to the initial spread in ENS4. Statistical significance tests were performed for growth rate comparisons of ENS4 vs. ENS20 and ENS4^{phys} vs. ENS20^{phys} for each field using Welch's t-test ($\alpha = 0.05$). This test determined whether the differences between the average filtered and adjusted variances of ENS4 and ENS20 (or ENS4^{phys} and ENS20^{phys}) at forecast hour 33¹ were significant.

The mean variance growth rates obtained from this methodology are shown in Fig. 5 (statistically significant differences indicated by asterisks). For mass-related fields, ENS4 growth rates are ~30% higher than those of ENS20 with differences that are statistically significant. These differences are consistent with faster perturbation growth expected as smaller scales are resolved that feed back to the larger scales (Lorenz 1969; Smagorinsky 1969). The ENS20_{cp} subsets (i.e., 5-member subsets with same CP) tend to have slightly lower growth rates than ENS20, which is consistent with ENS20 having one additional source of model uncertainty relative to ENS20_{cp} subsets from varied CPs. ENS20 also has one more source of model uncertainty than ENS4, but higher resolution in ENS4 apparently has a greater impact than the additional model uncertainty in ENS20 because ENS4 has much higher growth rates.

For the mass-related fields, the “mixed-physics only” ensemble subsets (denoted “Phys” in Fig. 5) have growth rates ~90% lower than the subsets with both sets of perturbations (denoted “IC/LBC+Phys” in Fig. 5). The much lower growth rates for mass-related fields when using only mixed-physics is similar to results found by Kong et al. (2007) using a similar dataset. The differences in growth rates occur because the different physics schemes that parameterize surface and boundary

¹ Note, after variances are adjusted to the same initial value at forecast hour 9, the variance at forecast hour 33 is the value that determines the growth rate; hence, variances at forecast hour 33 are used for the significance tests.

layer processes mainly influence the PBL, so that mass-related fields dependent on an entire vertical column of the atmosphere exhibit little impact. In addition, the different microphysics and cumulus parameterizations, which can possibly have a more direct influence on layers of the atmosphere above the PBL, are only impacted where the schemes are active, which is usually only over a small fraction of the domain. On the other hand, IC perturbations can directly affect all atmospheric layers and are present over the entire model domain.

All of the growth rates for mass-related fields in ENS20^{phys} are larger than those from ENS4^{phys} with differences that are statistically significant. Thus, unlike the ENS4 vs. ENS20 comparison, the impact of one additional source of model uncertainty in ENS20^{phys} (from varied CPs) is greater than the impact of higher resolution in ENS4^{phys}. Also, similar to the ENS20_{cp} subsets, the ENS20^{phys}_{cp} subsets have smaller growth rates relative to ENS20^{phys} resulting from having one less source of model uncertainty (no varied CPs).

For low-level fields, the mean spread growth rates are much more variable than in the mass-related fields, and statistically significant differences occur between ENS4 and ENS20 (or ENS4^{phys} and ENS20^{phys}) for T2, Td2, PREC, and WSHR, with ENS4 (or ENS4^{phys}) having the higher growth rates for all the significant differences. This may indicate that, for some variables, higher resolution in ENS4 and ENS4^{phys} results in larger spread growth rates despite the extra source of model uncertainty in ENS20 and ENS20^{phys}, but for other variables, the extra source of model uncertainty in ENS20 and ENS20^{phys} “balances out” the impacts of higher resolution.

It is suspected that, for some of the low-level fields, systematic model biases associated with certain parameterization schemes and combinations of parameterization schemes are impacting the growth rates. These biases are important to consider in the context of an ensemble because systematic biases that increase forecast uncertainty do so “artificially” since the associated errors are not uncertain (Eckel and Mass 2005). Thus, as discussed in recent studies (e.g., Eckel and Mass 2005; Yuan et al.

2007; Hamill and Whitaker 2007; Yussouf and Stensrud 2008), calibration should be performed on raw ensemble output to achieve maximum ensemble utility. Because the small number of cases examined in this study makes it difficult to obtain a useful “training period”, a calibration that could be applied in a real-time forecasting environment is not attempted. The possible influence of bias associated with particular CPs on MUCAPE mean spread growth rates is illustrated by a time series of domain averaged MUCAPE for all ENS4 and ENS20 ensemble members (Fig. 6). At forecast hour 33, which is the time used as Var_f in the growth rate calculation (Eq. 1), ENS20_{GD} members (11-15) tend to have the largest biases in MUCAPE, followed by ENS20_{KF} (6-10) and ENS20_{BMJ} (1-5) members. These biases are consistent with the MUCAPE mean spread growth rates in Fig. 5 (i.e., larger biases in MUCAPE at forecast hour 33 inflate the spread and corresponding growth rates). The impact of bias on the variance growth rates will be explored in Section 3d.

Some other interesting features are revealed from the MUCAPE time series. First, for both ENS4 and ENS20 ensembles, all members that use the MYJ PBL scheme tend to have much higher MUCAPE than members using YSU, especially for the second diurnal peak within the forecast period. In addition, the ENS4 members that use Thompson microphysics with MYJ have higher MUCAPE than members using WSM-6 or Ferrier with MYJ. However, in the ENS20 ensemble, members that use WSM-6 with MYJ have slightly higher MUCAPE than members using Thompson. Also, the ENS20 MYJ members tend to have the peak MUCAPE occur about 1-3 hours earlier than in the RUC analysis, but all YSU members have the peak occurring at the same time as in the RUC analysis. For the ENS20 ensemble, both MYJ and YSU members have the peak MUCAPE occurring about 3-4 hours earlier than in the RUC analysis. The sensitivity of MUCAPE to the different PBL schemes very likely results from systematic temperature and moisture biases associated with each scheme. For example, it has been well documented that the YSU tends to form boundary layers that are too deep, warm, and dry; while the MYJ has a tendency for relatively shallow, cool, and moist boundary layers (e.g., Kain et al. 2005;

3) MIXED-PHYSICS ENSEMBLE VARIANCE CONTRIBUTION

To estimate the percent contribution of mixed-physics to spread in the IC/LBC+Phys ensembles, the ratio of the mean ensemble variance in the Phys ensembles to that of the corresponding IC/LBC+Phys ensembles [i.e. $\{\text{Var}(\text{Phys})/\text{Var}(\text{IC/LBC+Phys})\} * 100\%$] is computed for all 12 fields at forecast hours 09 and 33 (Fig. 7). Similar to the spread growth rate comparisons, statistical significance tests were performed using Welch's t-test ($\alpha = 0.05$) for differences in mixed-physics variance contributions between ENS4 and ENS20. Note the actual contributions to ensemble spread in the IC/LBC+Phys ensembles not only result from separate contributions from IC/LBC perturbations and mixed-physics, but also from an interaction term (which could be positive or negative) between the two error sources. Because ensembles using only IC/LBC perturbations were not used in this experiment, it is not possible to diagnose this interaction term, and the estimate of spread contribution from mixed-physics assumes the interaction term is negligible.

For mass-related fields, the mixed-physics variance contributions in ENS20 are larger than ENS4 with differences that are statistically significant, consistent with the additional model uncertainty in ENS20 from varied CPs. The ENS20_{cp} subsets generally have contributions similar to those in ENS4. The mixed-physics contributions to ensemble spread for mass-related fields decrease as higher atmospheric levels are examined which likely occurs because the higher levels are impacted less by the boundary layer where the boundary layer physics have the greatest impact (Fig. 7). Relative to the low-level fields, mixed-physics contributions to ensemble spread are generally much smaller for the mass-related fields, which is consistent with earlier discussed results (e.g., Figs. 3 and 5). Perhaps the most noticeable feature for the mass-related fields in Fig. 7 is that the mixed-physics variance contributions for all ensemble subsets are higher at forecast hour 33 than 09, implying that the influence of model

uncertainty on ensemble spread increases within the forecast period analyzed.

For the low-level fields, similar to the variance growth rates (Fig. 5), there is much more variability in variance contributions among the different fields examined. Mixed-physics variance contributions range from around 10% for 850Td in ENS4 at forecast hour 9 to around 85% for PREC in ENS4 and ENS20 at forecast hour 33. Furthermore, similar to the mass-related fields, the ensemble subsets for most of the low-level fields have increasing mixed-physics variance contributions with increasing forecast lead time. By far, the highest mixed-physics variance contributions occur with PREC, which is not surprising because, overall, the physics parameterizations are particularly active in association with precipitation and two of the parameterizations (cumulus and microphysics schemes) are directly associated with precipitation production.

4) IMPACT OF MODEL BIASES ON MEAN VARIANCE GROWTH RATES

To explore the impact biases in different fields have on the mean spread growth rates, systematic and non-systematic biases are removed from each ensemble member for each field at all forecast hours using a procedure based on probability matching (Ebert 2001) described by Clark et al. (2009). The procedure uses probability matching to reassign the distribution of a forecast field with that of the observed field (RUC or stage IV analysis) so that the modified forecast fields have the same spatial pattern as the original forecasts, but have values adjusted so the distribution of their amplitudes exactly matches that of the analyses resulting in zero bias. Thus, the adjusted ensemble member forecasts all have the same distribution of values in the analyses, but with different spatial patterns. Ensemble variances computed from the adjusted forecasts can be interpreted as the variance resulting solely from the placement of “features”.

The mean variance growth rates for “bias-corrected” fields along with the differences between the bias-corrected growth rates and raw growth rates are displayed in Fig. 8. For the mass-related

fields, the growth rates for the IC/LBC+Phys ensembles are noticeably less relative to the raw growth rates, as shown by the growth rate differences (Fig. 8b). In addition, the differences between ENS4 and ENS20 bias-corrected growth rates are much less than for the raw growth rates and no longer significant, which implies that the higher spread in raw ENS4 mass-related fields relative to ENS20 (Fig. 5) can be attributed to larger differences in forecasts of the amplitude of features as opposed to placement.

In general, the variance growth rates for the low-level fields change less than for the mass-related fields after the bias-correction procedure is applied (Fig. 8b), which may be related to differences in the evolution of amplitude errors with increasing forecast lead time. For the mass-related fields, IC/LBC perturbations mainly contribute to amplitude errors which become larger as forecast lead time increases. Thus, applying the bias-correction procedure, which eliminates all amplitude errors, will decrease the ensemble variance more at the later lead times resulting in slower spread growth rates. However, for the low-level fields, the physics perturbations can quickly create large differences in the amplitude of features. However, these differences can quickly saturate (e.g., Stensrud et al. 2000) so that at later forecast lead times the amplitude errors are similar to those at earlier times. Thus, eliminating all amplitude errors will affect ensemble variance similarly at all forecast lead times and variance growth rates will not be strongly impacted. The variance growth rates for low-level fields that are strongly impacted by bias-correction are likely those that have amplitude errors (bias) that *change* with forecast lead time. For example, the domain averaged time series of T2 (Fig. 4) implies that T2 biases are similar for corresponding periods within the diurnal cycle for ENS4 and ENS20 ensemble members. Consistent with these similar biases, the growth rates for bias-corrected T2 are not very different from the raw T2 growth rates. However, domain averaged time series of MUCAPE (Fig. 6) show that MUCAPE biases during the latter part of the forecast (hours 24-33) are very different than those from the first part of the forecast (hours 0-9), and, consistent with the different biases, some of the

growth rates for bias-corrected MUCAPE change dramatically relative to the raw MUCAPE growth rates (Fig. 8b).

To more clearly show how the variance growth rates are affected by the bias-correction procedure, time series of mean differences between raw and bias-corrected ensemble variance for MSLP, T2, and MUCAPE are shown in Fig. 9. As discussed above, for MSLP, the bias-correction reduces ensemble variance more at later forecast lead times (Figs. 9a-b). Trends are not as noticeable for the other low-level fields. However, it is clear from Fig. 9f that the impact of the bias-correction is dependent on the CP used; for ENS20_{GD} (ENS20_{BMJ}), the bias-correction procedure results in more (less) positive ensemble variances at later forecast lead times. Furthermore, ENS20_{BMJ} has larger mean ensemble variances than ENS4 and ENS20 as well as the other ENS20_{cp} sub-ensembles, which should be further investigated in future work.

Other notable features from the bias-corrected growth rates in Fig. 8 are that ENS4 almost always has higher growth rates than ENS20 which is expected because smaller scales of motion (which have faster growth rates) are resolved in ENS4. However, for many of the variables (e.g., PREC, MUCAPE, and T2), ENS20_{cp} sub-ensembles actually have larger mean spread growth rates than ENS20. This is unexpected behavior because one less source of model uncertainty in ENS20_{cp} should result in slower spread growth. Further analyses, which are beyond the scope of this paper, are needed to explain this unexpected behavior.

b. Spread-skill relationship

Ideally, in a skillful ensemble that accurately accounts for all sources of forecast uncertainty, the ensemble variance should be a reliable predictor of the forecast skill (e.g., Grimit and Mass 2007). To quantify the variance-MSE relationship, past works have used linear correlations (e.g., Jones et al. 2007). This study also employs variance-MSE linear correlations, but care should be taken interpreting

the correlation coefficients because, as shown by Grit and Mass (2007), error statistics tend to exhibit increasing variance with increasing ensemble spread so that the variance-MSE relationship cannot be assumed to be linear. Thus, as noted in a similar analysis conducted by Jones et al. (2007), the linear correlation coefficients only provide an estimate of the predictability of ensemble skill.

Statistical consistency describes how well the ensemble variance matches the MSE when averaged over many cases (Talagrand et al. 1999; Eckel and Mass 2005). Thus, unlike the variance-MSE relationship, the amount of correlation is not considered with statistical consistency. A statistical consistency analysis can also provide information on whether an ensemble system is over- or under-dispersive. In an under-dispersive (over-dispersive) ensemble the average MSE is larger (smaller) than the ensemble variance. In this study, the variance-MSE and statistical consistency analyses are used as simple methods for inferring the impacts of the different spread growth rates on the quality of the ensemble forecasts. Note that we only examine spread-skill metrics for the ensembles that have both IC/LBC perturbations and mixed-physics because the lack of IC/LBC perturbations in ENS4^{phys} and ENS20^{phys} degrades the statistical consistency for all fields (not shown).

To illustrate the variance-MSE correlation in the ENS4 and ENS20 ensembles for the different fields examined, scatter-plots of ensemble variance vs. MSE are displayed in Fig. 10. Each panel in Fig. 10 contains variance-MSE points for each case and for each forecast hour (20 cases x 33 times = 660 points for each panel) and correlation coefficients indicate the degree of correspondence between ensemble variance and MSE (i.e., the reliability of ensemble variance as a predictor of forecast skill). Rank histograms (e.g., Hamill 2001) provided in Fig. 10 valid at forecast lead times of 09 and 33 also provide information regarding representation of forecast uncertainty: flat rank histograms imply an accurate depiction of forecast uncertainty, U-shaped (n-shaped) rank histograms imply under-dispersion (over-dispersion), and right (left) skewness indicates a tendency for over-prediction (under-prediction). To allow for a more convenient comparison between ENS4 and ENS20, the 16 bins composing the

ENS20 rank histograms were regrouped into 6 bins which each contain an equal portion of the original 16 bins². This “regrouping” technique has also been used in Clark et al. (2009).

To illustrate statistical consistency, the average MSE and ensemble variance at forecast hours 9 and 33 for fields in ENS4 and ENS20 are shown in Fig. 11. These forecast hours are chosen because they were the times used to compute spread growth rates. To estimate whether the variance/MSE differences between ENS4 and ENS20 were statistical significance, a resampling procedure using a 95% confidence interval was applied (Wilks 1995 pp. 145-150). Differences between average randomly resampled variance and MSE for two artificial datasets were computed 10000 times, and the distribution of the differences between the two artificial datasets was used to determine whether the “true” statistic fell outside of the 95% confidence interval which would suggest statistical significance.

For mass-related fields in Figs. 10a-h, the variance-MSE correlations in ENS4 and ENS20 are very low suggesting that ensemble variance is not a reliable indicator of forecast skill for these fields. The highest correlations occur for the MSLP forecasts from ENS4 ($R^2 = 0.19$). Considering previous work that has also found small spread-error correlations for fields like mid-tropospheric geopotential height (e.g., Buizza 1997), these results are not surprising. However, there are noticeable differences in the distribution of variance-MSE points for the mass-related fields: in ENS4 there are more points to the right of the diagonal than in ENS20 indicating that ensemble variance is greater than MSE more frequently in ENS4. Furthermore, the variance-MSE points in ENS20 appear to be positioned in a vertically oriented “plume”, while those in ENS4 veer towards the right (i.e. toward higher values of ensemble variance).

These results for mass-related fields are reflected by the statistical consistency analysis in Figs.

2 The formulas for regrouped rank histogram bins can be expressed as: $ENS20_{regroup}(1) = ENS20(1) + ENS20(2) + ENS20(3) \times 2/3$; $ENS20_{regroup}(2) = ENS20(3) \times 1/3 + ENS20(4) + ENS20(5) + ENS20(6) \times 1/3$; $ENS20_{regroup}(3) = ENS20(6) \times 2/3 + ENS20(7) + ENS20(8)$; $ENS20_{regroup}(4) = ENS20(9) + ENS20(10) + ENS20(11) \times 2/3$; $ENS20_{regroup}(5) = ENS20(11) \times 1/3 + ENS20(12) + ENS20(13) + ENS20(14) \times 1/3$; $ENS20_{regroup}(6) = ENS20(14) \times 2/3 + ENS20(15) + ENS20(16)$, where $ENS20_{regroup}(x)$ is the value for regrouped rank histogram bins $x = 1, 2, \dots, 6$, and $ENS20(x)$ is the value for raw rank histogram bins $x = 1, 2, \dots, 16$.

11a-d. At forecast hour 33, differences between ensemble variance and MSE in ENS4 are noticeably less than in ENS20 indicating that ENS4 is more statistically consistent than ENS20 (although, differences were only found to be significant for 850Z and MSLP). The better statistical consistency in ENS4 results not only from greater spread at forecast hour 33, but also lower MSEs relative to ENS20 (Figs. 11a-d). Furthermore, while ENS4 forecasts for mass-related fields do not exhibit much change in statistical consistency between forecast hours 9 and 33, ENS20 forecasts become increasingly under-dispersive. The increasing under-dispersion in ENS20 is also indicated by the change to more “u-shaped” rank histograms (Figs. 10b, d, f, and h; gray shaded) from forecast hours 9 to 33. The ENS4 rank histograms (Figs. 10b, d, f, h; black outline) do not exhibit a noticeable change in shape and are right-skewed at both times implying under-prediction of mass-related fields. This under-prediction could possibly result from a general cool bias in the lower part of the troposphere, but further examination is beyond the scope of this study. For mass-related fields, the statistical consistency results imply that the faster spread growth rates in ENS4 are contributing to more reliable forecasts than in ENS20.

The variance-MSE correlations for low-level fields in ENS4 and ENS20 (Figs. 10i-x) are quite variable among the fields examined, with 850WMAG (Figs. m-n), PREC (Figs. s-t), MUCAPE (Figs. u-v), and WSHR (Figs. w-x) having the highest values. Furthermore, the rank histograms indicate that ENS4 and ENS20 both suffer from systematic biases and/or under-dispersion for most of the low-level fields. For example, warm T2 biases and dry Td2 biases are revealed from the right and left skewed rank histograms, respectively, in Figs. 10i-l. Also, the U-shaped rank histograms for 850Td (Fig. 10q) imply under-dispersion. The statistical consistency analyses (Figs. 11e-l) show that by forecast hour 33 there is little difference between ENS4 and ENS20 statistical consistency for temperature and dewpoint fields (T2, Td2, 850T, and 850Td), while ENS4 has better statistical consistency for the other fields (850WMAG, PREC, MUCAPE, and WSHR all of which have differences that are statistically

significant). Similar to the mass-related fields, the more statistically consistent forecasts for low-level fields in ENS4 appear to result from a combination of lower MSEs and higher ensemble variances relative to ENS20.

4. Summary and discussion

This study compared ensemble spread growth and spread-error relationships for 12 different fields in a 4-km grid-spacing convection-allowing WRF model ensemble to that from a similarly configured but coarser 20-km grid-spacing convection-parameterizing WRF model ensemble. Ensemble subsets that used both IC/LBC perturbations and mixed physics were compared along with subsets that only contained mixed physics. In addition, the contribution of ensemble variance from the mixed-physics in the 4-km and 20-km ensembles was inferred by comparing the mixed-physics only ensemble subsets to those that contained both IC/LBC perturbations and mixed-physics. A total of 20 cases were examined for a domain centered over the central US. Main findings are summarized below. In addition, Table 5 highlights the main differences that were observed for the various ensemble comparisons.

Spread growth rates for mass-related fields were higher in ENS4 than in ENS20 by about 30%. The mixed-physics only ensemble subsets (Phys) had much smaller spread growth rates than ensemble subsets with both IC/LBC perturbations and mixed-physics (IC/LBC+Phys). For low-level fields, spread growth rates were quite variable among the different fields examined, however, for all of the statistically significant differences between ENS4 and ENS20 (or ENS4^{phys} and ENS20^{phys}), the 4-km ensembles had the higher spread growth rates. The differences between spread growth rates in 4- and 20-km ensemble configurations are summarized in the first two columns of Table 5.

The contributions to spread from mixed-physics in the mass-related fields were generally much smaller than for the low-level fields, consistent with the differences between the IC/LBC+Phys and

Phys spread growth rates. In addition, the contribution to spread from mixed-physics increased with increasing forecast lead time. Similar to the spread growth rates, there was much more variability in the spread contributions among the different low-level fields than for the mass-related fields. For most of the statistically significant differences between ENS4 and ENS20 mixed-physics spread contributions, ENS20 had larger contributions (third and fourth columns of Table 5) which was related to ENS20 having an additional source of model uncertainty in the form of varied cumulus parameterizations.

It was suspected from examination of time series for domain averaged low-level fields that systematic model biases may be having an impact on mean ensemble spread, as found by Clark et al. (2009). To explore these potential impacts, biases in all forecast fields were removed by replacing the distributions of values in the forecast fields with the distribution of values in corresponding RUC or stage IV analyses as described by Clark et al. (2009). It was found that the spread growth rates for the mass-related fields tended to be impacted most by the bias-correction procedure, which may be related to differences in how amplitude errors evolve with increasing forecast lead time. The MUCAPE fields in the ENS20_{BMJ} and ENS20_{GD} ensemble subsets were found to be the low-level fields most influenced by the bias-correction procedure because of different systematic biases at later forecast lead times relative to earlier ones.

Finally, to put the spread growth rates for the different fields examined in an appropriate forecasting context, an analysis of the variance-MSE relationship and statistical consistency was conducted. This analysis was important because increased ensemble dispersion does not necessarily imply a better spread-skill relationship. Variance-MSE correlation coefficients indicated that, in general, ensemble variance was not a reliable indicator of forecast uncertainty. Furthermore, at forecast hour 33, ENS4 had better statistical consistency than ENS20 for mass-related fields, 850WMAG, PREC, MUCAPE, and WSHR; however, for temperature and dewpoint fields (T2, Td2, 850T, and 850Td), there were no noticeable differences (fifth column of Table 5). It was found that a combination

of higher spread and lower MSEs contributed to the improved statistical consistency in ENS4.

Generally, the results from this study could be interpreted as encouraging for future convection-allowing ensemble systems simply because of the improved statistical consistency for many fields. However, additional work is needed to diagnose why the higher resolution of ENS4 did not seem to improve spread-error metrics for temperature and dewpoint fields. Perhaps the parameterization schemes that the temperature and dewpoint fields depend on are simply not very sensitive to grid-spacing, or perhaps a reduction of the large systematic errors in the temperature and dewpoint forecasts would result in better statistical consistency. Further work should also analyze larger sets of cases for different periods and regions. Because this study focuses on the spring season in the central US, a time period and region characterized by frequent convective systems, examination of other times/regions in which strong convection is not as prevalent may give different results. Furthermore, it would be very useful to assess how much the improved statistical consistency in some fields increases the skill of probabilistic forecasts. Finally, the behavior of ensemble spread observed in this study should be helpful for future ensemble design. In particular, the variety of results for different fields suggests that future ensemble design should give careful consideration to the specific types of forecast fields desired by the user. Also, the recognition of systematic model biases should provide motivation for improving the physics parameterizations used with convection-allowing grid-spacing.

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Figure 1 Domains for a) NCEP SREF ensemble members b) ENS4 and ENS20 ensemble members, and c) the analyses conducted in this study.

Figure 2 Gray shaded dates indicate when 10-member SSEF system simulations were conducted for SE2007 and dark gray shading indicates which cases are used in the analysis for this study.

Figure 3 Time series of mean ensemble variance from ENS4 (gray line) and ENS20 (black line) for the variables a) 500Z, c) 700Z, and e) 850Z, g) MSLP, i) T2, k) Td2, m) 850WMAG, o) 850T, q) 850Td, s) PREC, u) MUCAPE, w) WSHR. b), d), f), h), j), l), n), p), r), t), v), and x) same as a), c), e), g), i), k), m), o), q), s), u), and w) except for ENS4^{phys} (gray line) and ENS20^{phys} (black line). Boxplots overlay the mean at each time interval. For the boxplots, the interquartile-range (IQR) is indicated by the area enclosed by a box, outliers defined by values outside of 1.5*IQR are marked by dots, and horizontal lines mark the smallest and largest values that are not outliers.

Figure 4 Time series of mean domain averaged 2-m temperature from all cases for the ENS4 and ENS20 ensemble members and means, along with the RUC analysis.

Figure 5 Mean variance growth rates from the ENS4, ENS4^{phys}, ENS20, and ENS20^{phys} ensembles for fields shown in Figure 3. Growth rates for five member subsets of ENS20 and ENS20^{phys} that have the same cumulus parameterization are also shown (marked by the horizontal lines on the lightest gray histogram). The histograms to the left for each variable indicate growth rates for ensembles that have IC/LBC perturbations and mixed-physics (IC/LBC+Phys) and the ones to the right are for mixed-physics only ensembles (Phys). Asterisks below the histogram sets indicate statistically significant differences between ENS4 and ENS20 or ENS4^{phys} and ENS20^{phys} [black (gray) asterisks indicate ENS4 or ENS4^{phys} (ENS20 or ENS20^{phys}) had larger growth rates]. A legend is provided at the top of the figure.

Figure 6 Time series of mean domain averaged MUCAPE for the ENS20 ensemble members and RUC analyses. A legend is provided in the upper right portion of the plot with numbers indicating ENS20 ensemble members corresponding to those listed in Table 3.

Figure 7 Mean variance ratio [%; $\text{Var}(\text{ENS4}^{\text{phys}})/\text{Var}(\text{ENS4})$ and $\text{Var}(\text{ENS20}^{\text{phys}})/\text{Var}(\text{ENS20})$] from the 4-km and 20-km grid-spacing ensembles and from the 5-member 20-km ensemble subsets for the fields in Figure 3. The histograms to the left (right) for each field are for forecast hour 9 (33). Asterisks below the histogram sets indicate statistically significant differences between ENS4 and ENS20 [black (gray) asterisks indicate ENS4 (ENS20) had larger variance contributions. A legend is provided at the top of the figure.

Figure 8 a) Same as Fig. 5, except for bias-corrected forecasts. b) Differences between the bias-corrected and raw growth rates (bias-corrected minus raw).

Figure 9 Differences in mean ensemble variance between the raw and bias-corrected forecasts from ENS4 and ENS20 for the fields a) MSLP (hPa^2), c) T2 (K^2), and e) MUCAPE [$(\text{J/kg})^2$]. b), d), and f) same as a), c) and e) except for the ENS20_{cp} configurations. The triangles denote forecast hours 09 and 33, which are the times used in the computation of variance growth rates.

Figure 10 Scatter plots of ensemble variance vs. MSE of the ensemble mean from ENS4 (black dots) for the raw fields a) 500Z, c) 700Z, and e) 850Z, g) MSLP, i) T2, k) Td2, m) 850WMAG, o) 850T, q) 850Td, s) PREC, u) MUCAPE, w) WSHR. b), d), f), h), j), l), n), p), r), t), v), and x) same as a), c), e), g), i), k), m), o), q), s), u), and w) except for ENS20 (grey dots). Correlation coefficients are provided in the right-top corner of each panel and rank histograms for ENS4 (black outline) and ENS20 (grey shaded) for forecast hours 09 and 33 are displayed in the bottom-right corner of the ENS20 plots.

Figure 11 Average MSE of the ensemble mean and ensemble variance at forecast hours 9 and 33 for the ENS4 and ENS20 ensembles for the fields, a) 500Z (m^2), b) 700Z (m^2), c) 850Z (m^2), d) MSLP

(hPa²), e) T2 (K²), f) Td2 (K²), g) 850WMAG [(m/s)²], h) 850T (K²), i) 850Td (K²), j) PREC (mm²), k) MUCAPE [(J/kg)²], and l) WSHR [(m/s)²]. The times at which differences between the ensemble variance and MSE differences in ENS4 and ENS20 are statistically significant are highlighted the darker shade of gray

List of Tables

Table 1 ENS4 ensemble member specifications. NAMA and NAMf indicate NAM forecasts and analyses, respectively; em_pert and nmm_pert are perturbations from different SREF members; and em_n1, em_p1, nmm_n1, and nmm_p1 are different SREF members that are used for LBCs. The remaining table elements are described in the text.

Table 2 Same as Table 1 except for ENS4^{phys} ensemble member specifications.

Table 3 ENS20 ensemble member specifications. The members are grouped into 5-member subsets that have the same cumulus parameterizations. The ICs/LBCs table elements represent various SREF members and the remaining table elements are described in the text.

Table 4 Same as Table 3 except for ENS20^{phys} ensemble member specifications.

Table 5 Summary of results. For each variable, the different columns indicate whether faster spread growth, larger mixed-physics spread contributions, or better statistical consistency was observed for the 4- or 20-km grid-spacing ensemble configurations. Only statistically significant differences are indicated.

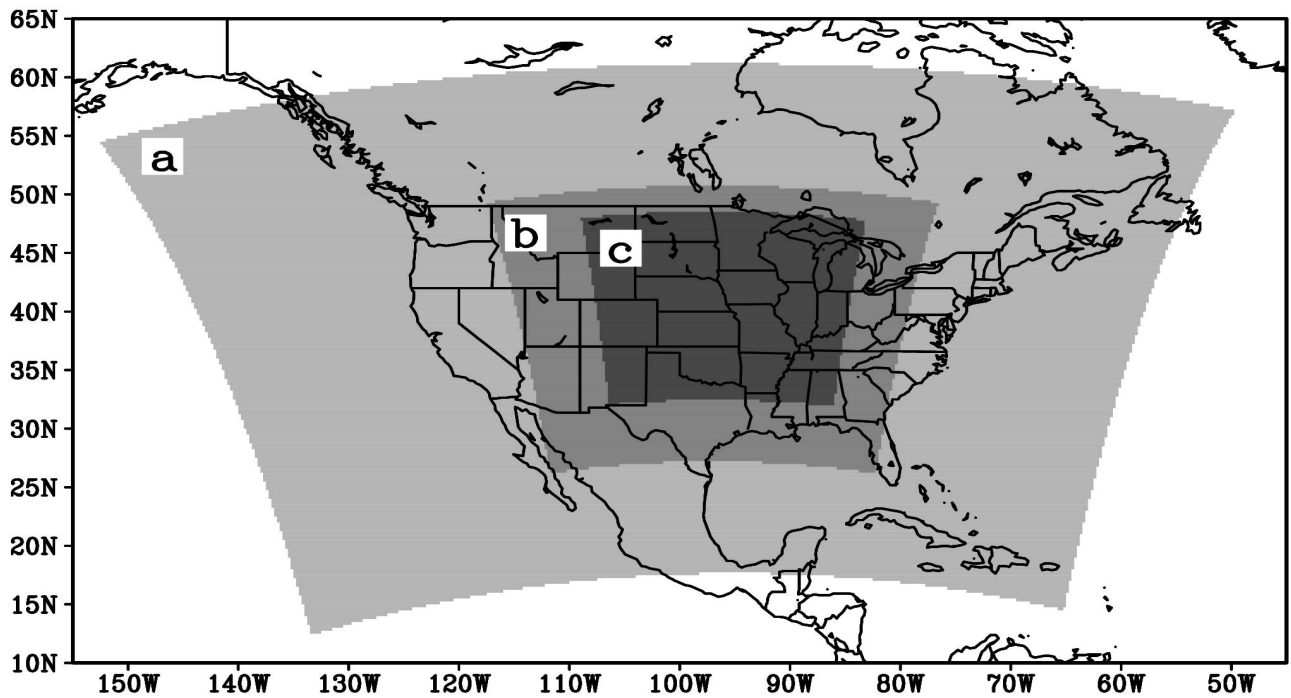


Figure 1 Domains for a) NCEP SREF ensemble members b) ENS4 and ENS20 ensemble members, and c) the analyses conducted in this study.

April 2007							May 2007							June 2007						
S	M	T	W	T	F	S	S	M	T	W	T	F	S	S	M	T	W	T	F	S
1	2	3	4	5	6	7			1	2	3	4	5						1	2
8	9	10	11	12	13	14	6	7	8	9	10	11	12	3	4	5	6	7	8	9
15	16	17	18	19	20	21	13	14	15	16	17	18	19	10	11	12	13	14	15	16
22	23	24	25	26	27	28	20	21	22	23	24	25	26	17	18	19	20	21	22	23
29	30						27	28	29	30	31			24	25	26	27	28	29	30

Figure 2 Gray shaded dates indicate when 10-member SSEF system simulations were conducted for SE2007 and dark gray shading indicates which cases are used in the analysis for this study.

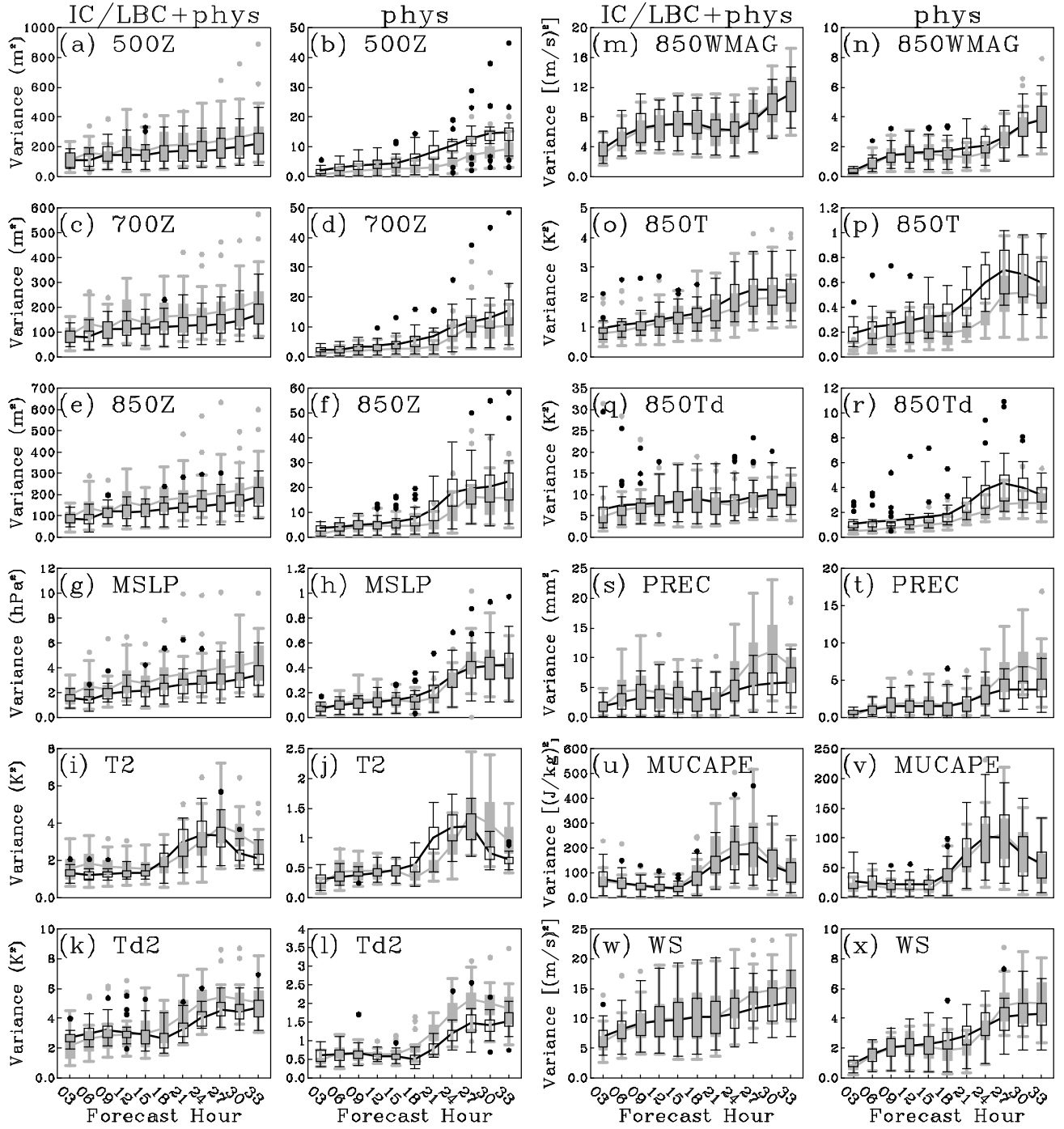


Figure 3 Time series of mean ensemble variance from ENS4 (gray line) and ENS20 (black line) for the variables a) 500Z, c) 700Z, and e) 850Z, g) MSLP, i) T2, k) Td2, m) 850WMAG, o) 850T, q) 850Td, s) PREC, u) MUCAPE, w) WS. b) 500Z, d) 700Z, f) 850Z, h) MSLP, j) T2, l) Td2, n) 850WMAG, p) 850T, r) 850Td, t) PREC, v) MUCAPE, x) WS. The same as a), c), e), g), i), k), m), o), q), s), u), and w) except for ENS4^{phys} (gray line) and ENS20^{phys} (black line). Boxplots overlay the mean at each time interval. For the boxplots, the interquartile-range (IQR) is indicated by the area enclosed by a box, outliers defined by values outside of 1.5*IQR are marked by dots, and horizontal lines mark the smallest and largest values that are not outliers.

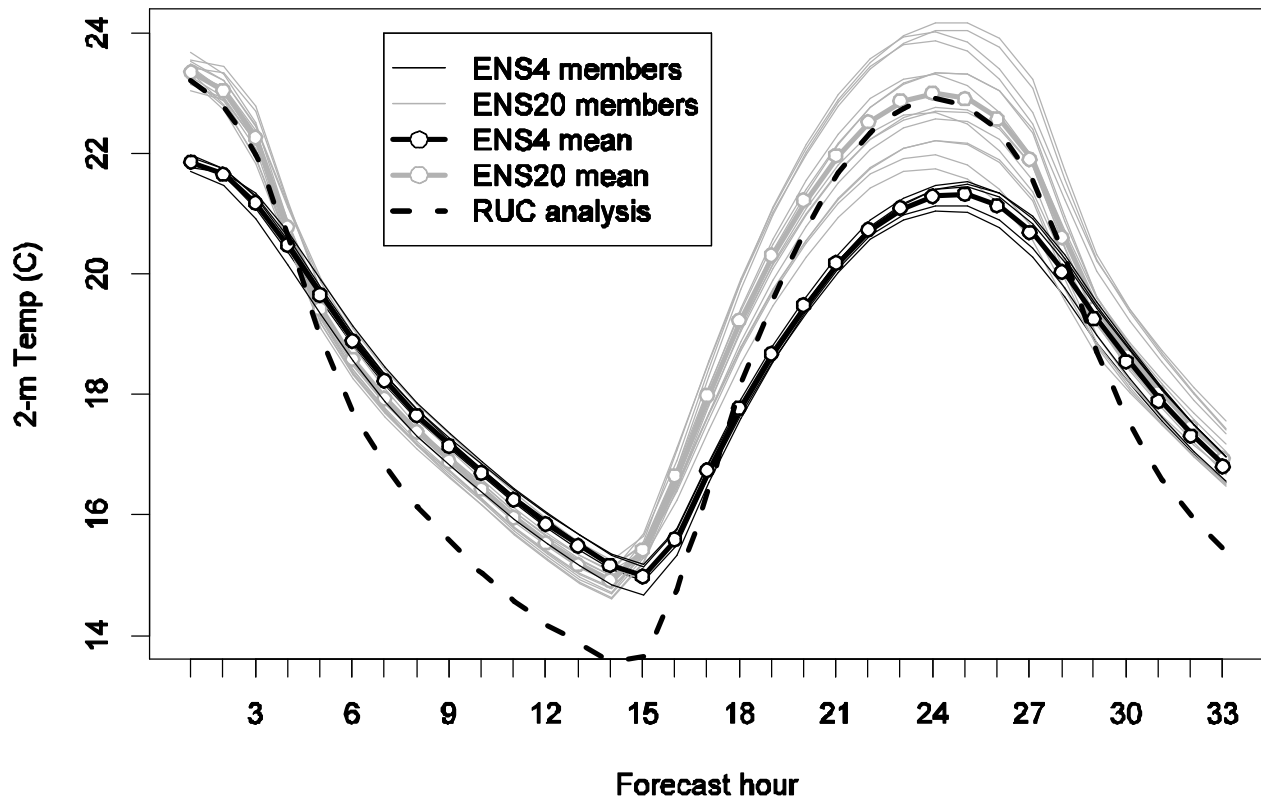


Figure 4 Time series of mean domain averaged 2-m temperature from all cases for the ENS4 and ENS20 ensemble members and means, along with the RUC analysis.

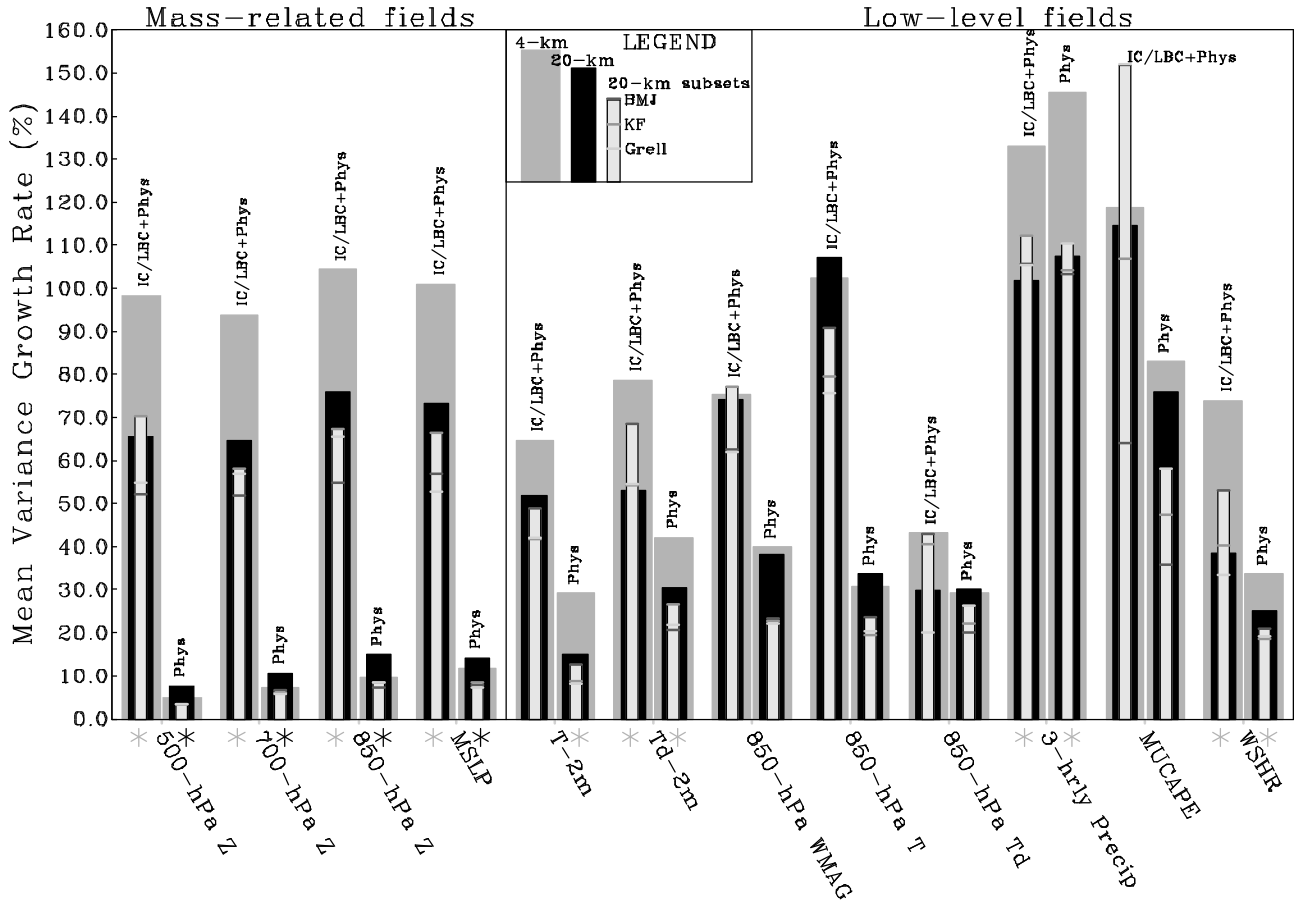


Figure 5 Mean variance growth rates from the ENS4, ENS4^{phys}, ENS20, and ENS20^{phys} ensembles for fields shown in Figure 3. Growth rates for five member subsets of ENS20 and ENS20^{phys} that have the same cumulus parameterization are also shown (marked by the horizontal lines on the lightest gray histogram). The histograms to the left for each variable indicate growth rates for ensembles that have IC/LBC perturbations and mixed-physics (IC/LBC+Phys) and the ones to the right are for mixed-physics only ensembles (Phys). Asterisks below the histogram sets indicate statistically significant differences between ENS4 and ENS20 or ENS4^{phys} and ENS20^{phys} [black (gray) asterisks indicate ENS4 or ENS4^{phys} (ENS20 or ENS20^{phys}) had larger growth rates]. A legend is provided at the top of the figure.

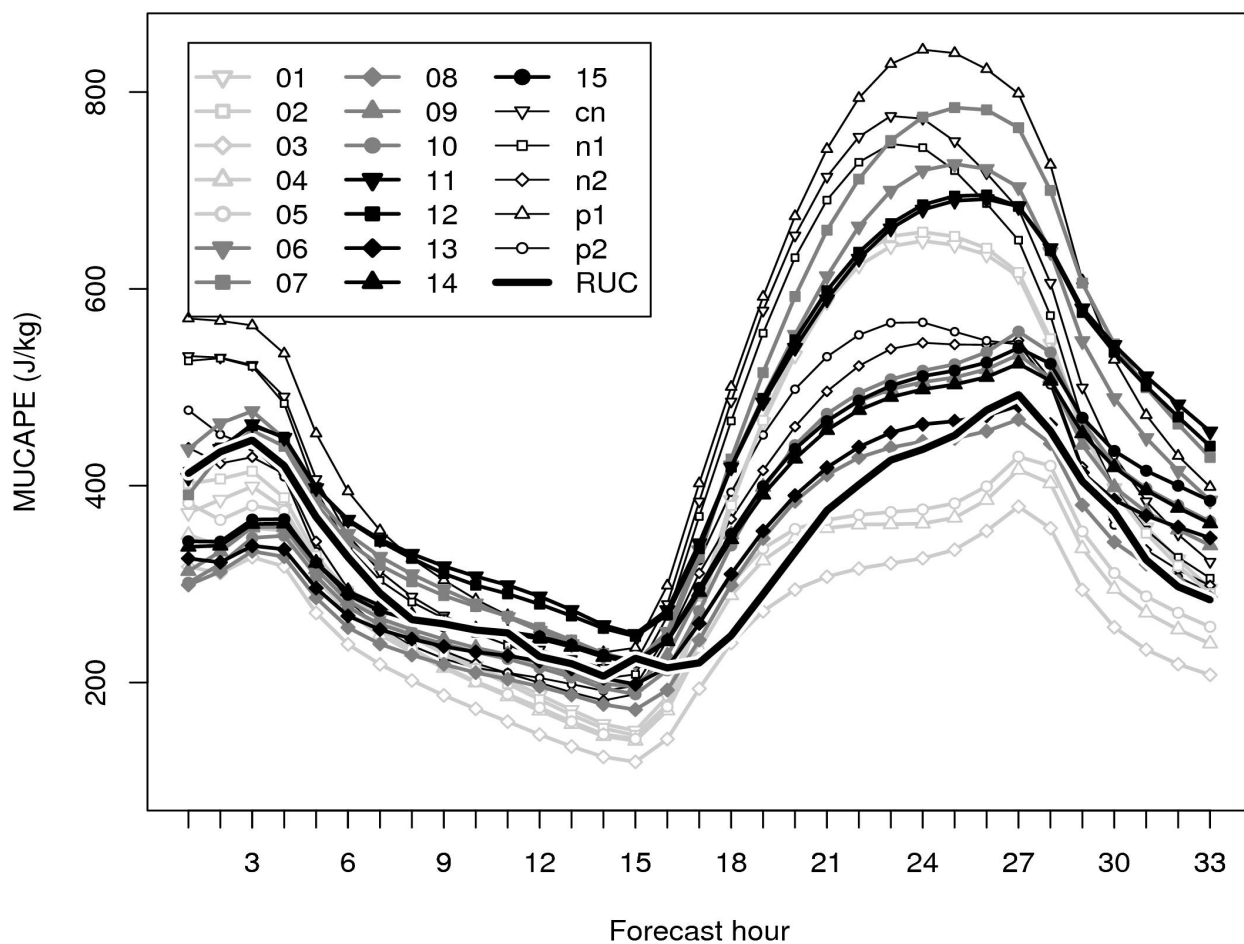


Figure 6 Time series of mean domain averaged MUCAPE for the ENS20 ensemble members and RUC analyses. A legend is provided in the upper right portion of the plot with numbers indicating ENS20 ensemble members corresponding to those listed in Table 3.

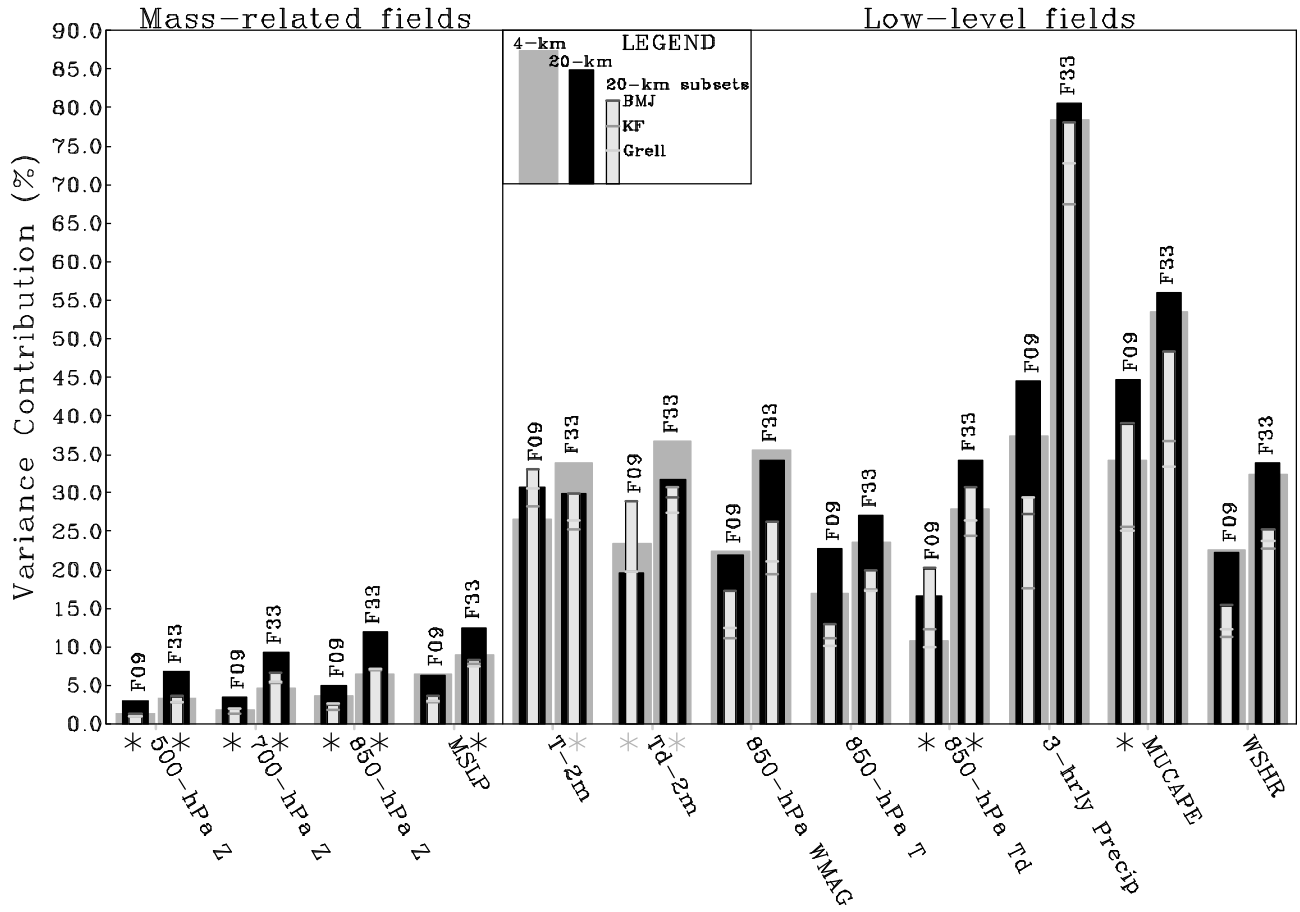


Figure 7 Mean variance ratio [%; $\text{Var}(\text{ENS4}^{\text{phys}})/\text{Var}(\text{ENS4})$ and $\text{Var}(\text{ENS20}^{\text{phys}})/\text{Var}(\text{ENS20})$] from the 4-km and 20-km grid-spacing ensembles and from the 5-member 20-km ensemble subsets for the fields in Figure 3. The histograms to the left (right) for each field are for forecast hour 9 (33). Asterisks below the histogram sets indicate statistically significant differences between ENS4 and ENS20 [black (gray) asterisks indicate ENS4 (ENS20) had larger variance contributions. A legend is provided at the top of the figure.

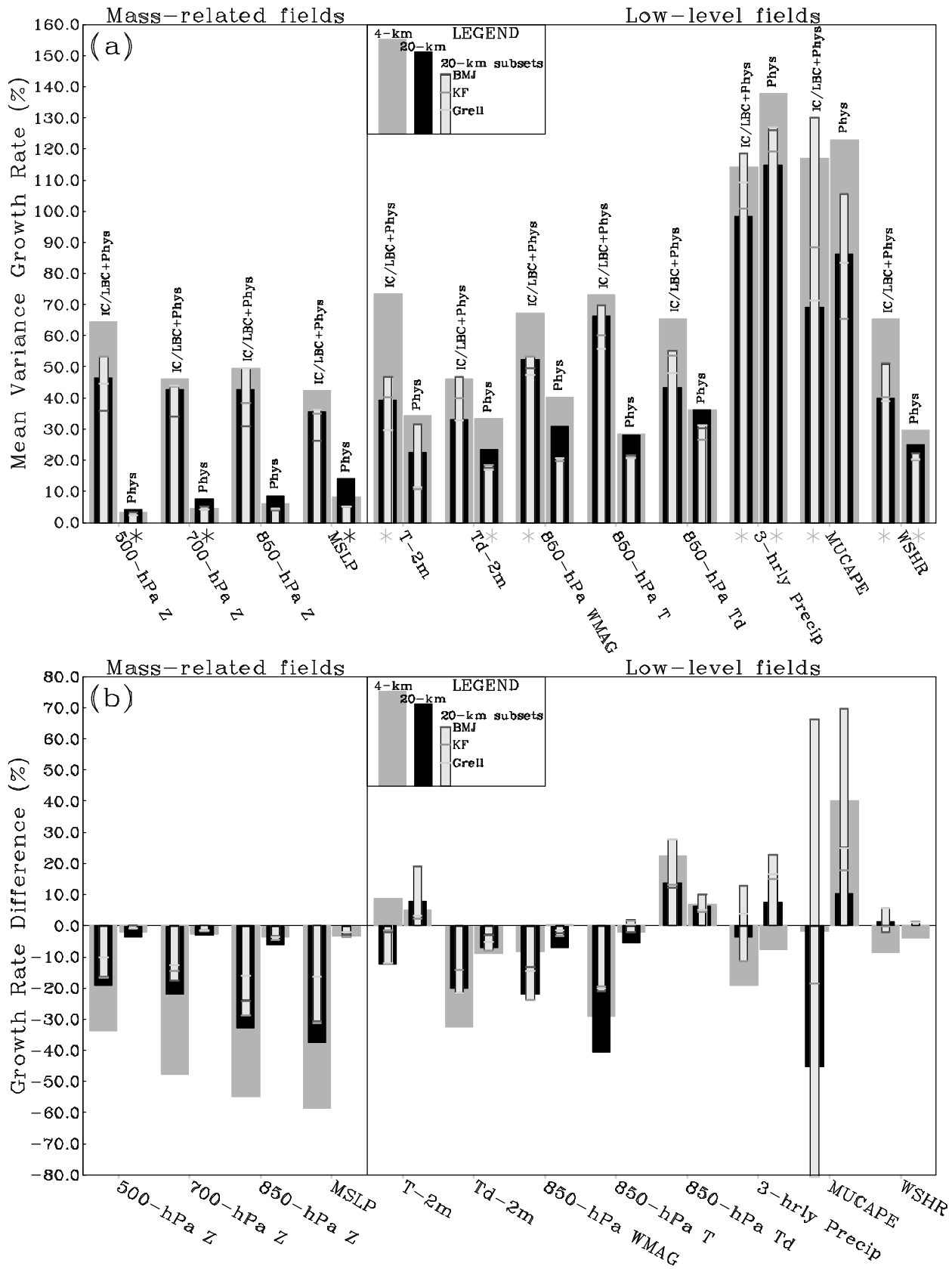


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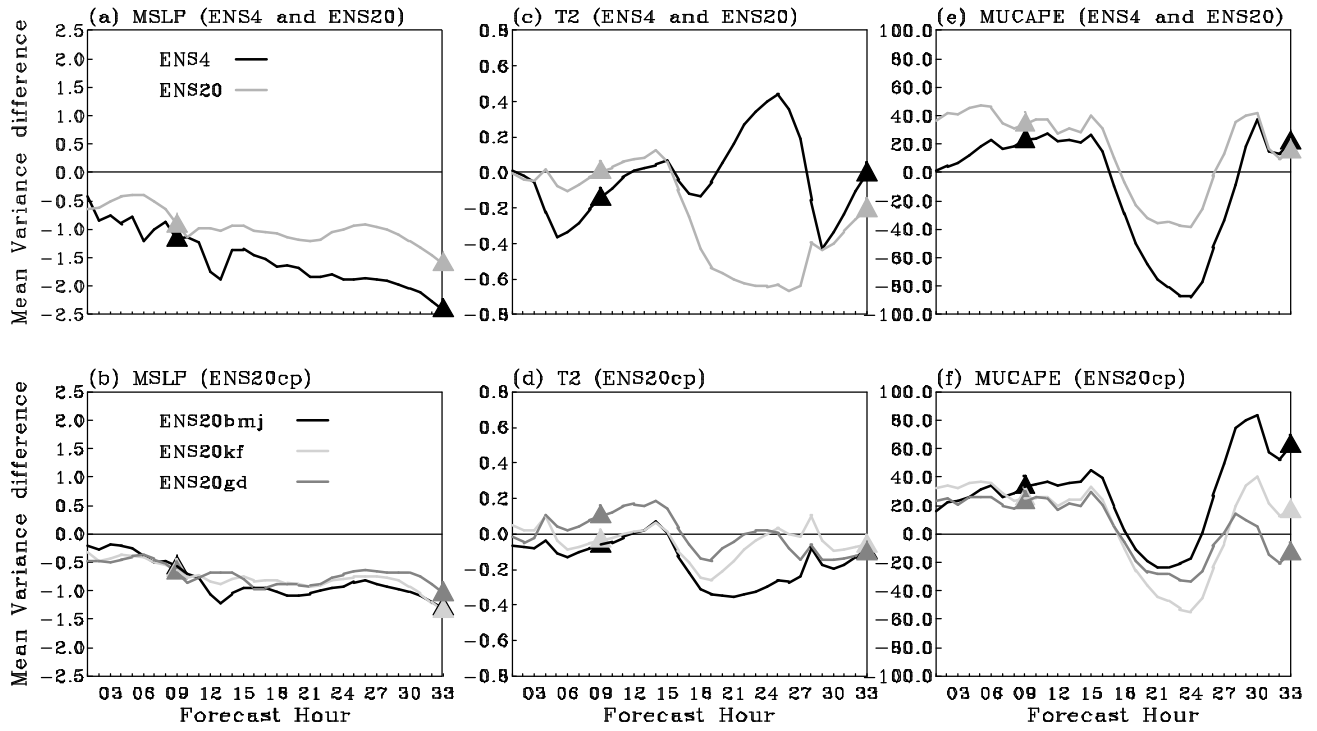


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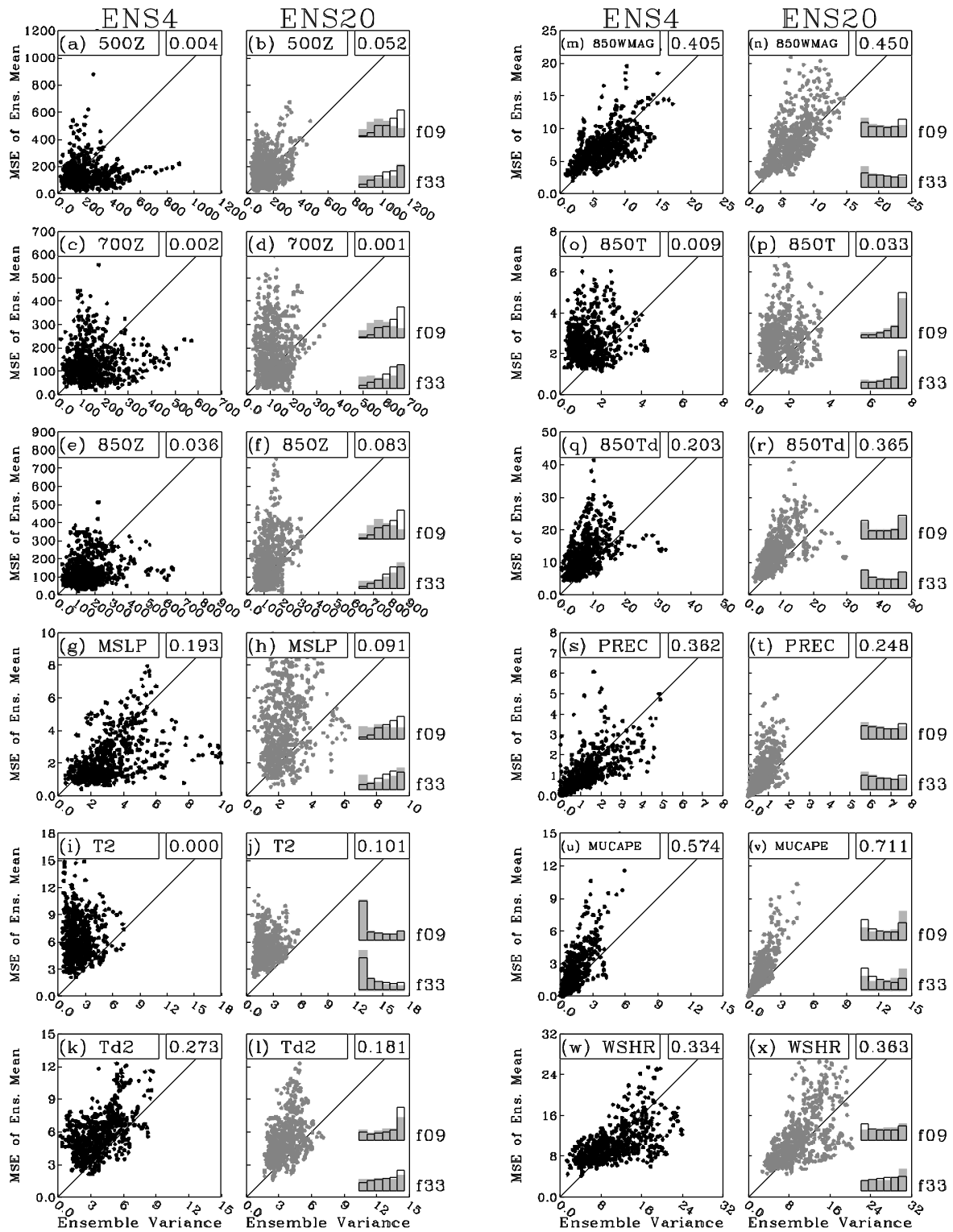


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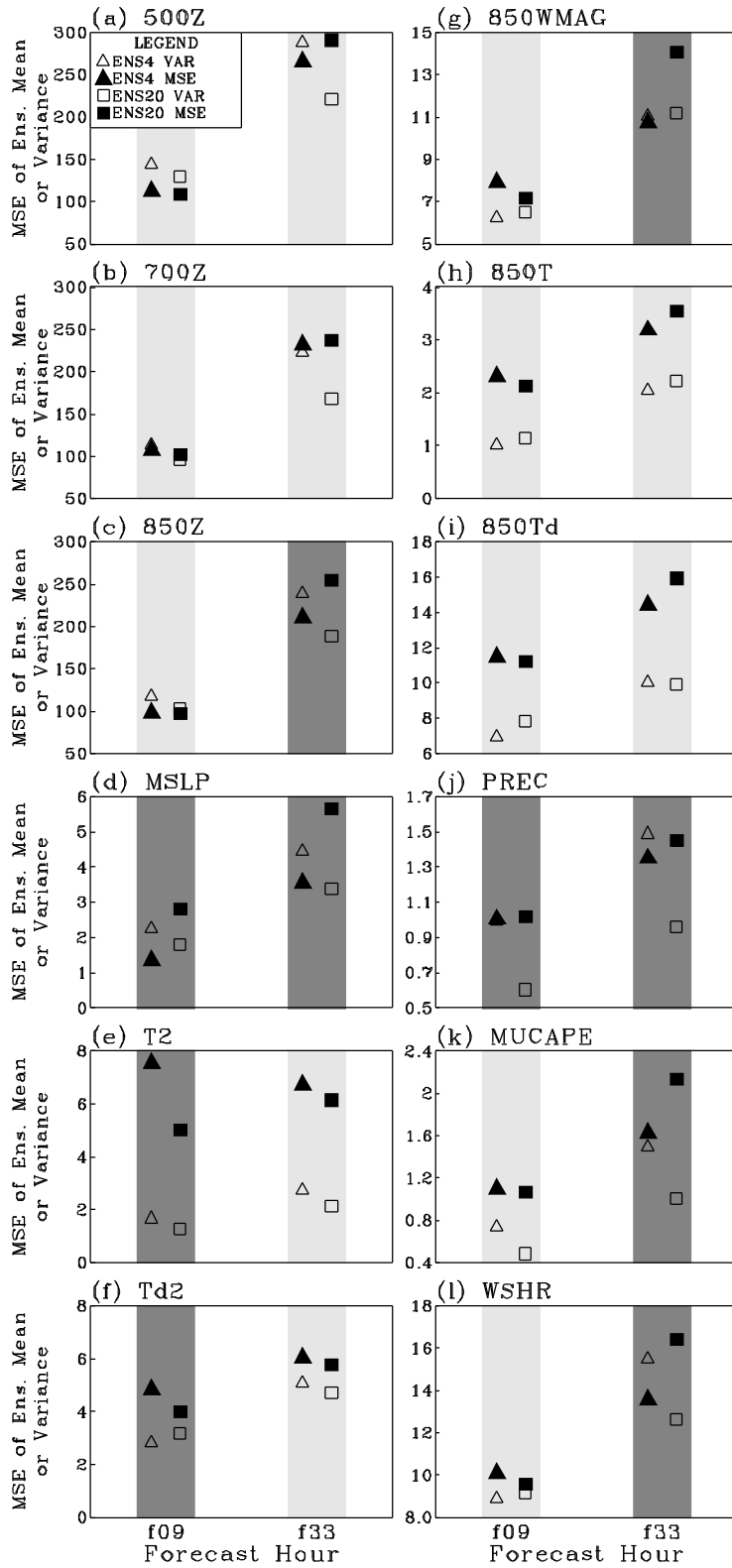


Figure 11 Average MSE of the ensemble mean and ensemble variance at forecast hours 9 and 33 for the ENS4 and ENS20 ensembles for the fields, a) 500Z (m^2), b) 700Z (m^2), c) 850Z (m^2), d) MSLP (hPa^2), e) T2 (K^2), f) Td2 (K^2), g) 850WMAG [$(m/s)^2$], h) 850T (K^2), i) 850Td (K^2), j) PREC (mm^2), k) MUCAPE [$(J/kg)^2$], and l) WSHR [$(m/s)^2$]. The times at which differences between the ensemble variance and MSE differences in ENS4 and ENS20 are statistically significant are highlighted the darker shade of gray.

Table 1 ENS4 ensemble member specifications. NAMa and NAMf indicate NAM forecasts and analyses, respectively; em_pert and nmm_pert are perturbations from different SREF members; and em_n1, em_p1, nmm_n1, and nmm_p1 are different SREF members that are used for LBCs. The remaining table elements are described in the text.

Ensemble Member	ICs	LBCs	Microphysics Scheme	Surface Layer Scheme	Boundary Layer Scheme
CN	21Z NAMa	18z NAMf	WSM-6	Janjic Eta	MYJ
N1	CN - em_pert	21z SREF em_n1	Ferrier	Janjic Eta	MYJ
P1	CN + em_pert	21z SREF em_p1	Thompson	Janjic Eta	MYJ
N2	CN - nmm_pert	21z SREF nmm_n1	Thompson	Monin-Obukhov	YSU
P2	CN + nmm_pert	21z SREF nmm_p1	WSM-6	Monin-Obukhov	YSU

Table 2 Same as Table 1 except for ENS4^{phys} ensemble member specifications.

Ensemble Member	ICs	LBCs	Microphysics Scheme	Surface Layer Scheme	Boundary Layer Scheme
PH1	21Z NAMa	18Z NAMf	Thompson	Janjic Eta	MYJ
PH2	21Z NAMa	18Z NAMf	Ferrier	Janjic Eta	MYJ
PH3	21Z NAMa	18Z NAMf	WSM-6	Monin-Obukhov	YSU
PH4	21Z NAMa	18Z NAMf	Thompson	Monin-Obukhov	YSU
PH5	21Z NAMa	18Z NAMf	Ferrier	Monin-Obukhov	YSU

Table 3 ENS20 ensemble member specifications. The members are grouped into 5-member subsets that have the same cumulus parameterizations. The ICs/LBCs table elements represent various SREF members and the remaining table elements are described in the text.

Ensemble Member	ICs/LBCs	Cumulus Scheme	Microphysics	Surface Layer	Boundary Layer
ENS20_{BMJ}					
1	em_ctl	BMJ	Thompson	Janjic Eta	MYJ
2	em_p1	BMJ	WSM-6	Janjic Eta	MYJ
3	em_n1	BMJ	WSM-6	Monin-Obukhov	YSU
4	nmm_ctl	BMJ	Thompson	Monin-Obukhov	YSU
5	nmm_p1	BMJ	Ferrier	Monin-Obukhov	YSU
ENS20_{KF}					
6	nmm_n1	KF	Thompson	Janjic Eta	MYJ
7	eta_ctl1	KF	WSM-6	Janjic Eta	MYJ
8	eta_n1	KF	WSM-6	Monin-Obukhov	YSU
9	eta_n2	KF	Thompson	Monin-Obukhov	YSU
10	eta_n3	KF	Ferrier	Monin-Obukhov	YSU
ENS20_{GD}					
11	eta_n4	Grell	Thompson	Janjic Eta	MYJ
12	eta_p1	Grell	WSM-6	Janjic Eta	MYJ
13	eta_p2	Grell	WSM-6	Monin-Obukhov	YSU
14	eta_p3	Grell	Thompson	Monin-Obukhov	YSU
15	eta_p4	Grell	Ferrier	Monin-Obukhov	YSU

Table 4 Same as Table 3 except for ENS20^{phys} ensemble member specifications.

Ensemble Member	ICs/LBCs	Cumulus Scheme	Microphysics	Surface Layer	Boundary Layer
ENS20^{phys}_{BMJ}					
16	eta_ctl2	BMJ	Thompson	Janjic Eta	MYJ
17	eta_ctl2	BMJ	WSM-6	Janjic Eta	MYJ
18	eta_ctl2	BMJ	WSM-6	Monin-Obukhov	YSU
19	eta_ctl2	BMJ	Thompson	Monin-Obukhov	YSU
20	eta_ctl2	BMJ	Ferrier	Monin-Obukhov	YSU
ENS20^{phys}_{KF}					
21	eta_ctl2	KF	Thompson	Janjic Eta	MYJ
22	eta_ctl2	KF	WSM-6	Janjic Eta	MYJ
23	eta_ctl2	KF	WSM-6	Monin-Obukhov	YSU
24	eta_ctl2	KF	Thompson	Monin-Obukhov	YSU
25	eta_ctl2	KF	Ferrier	Monin-Obukhov	YSU
ENS20^{phys}_{GD}					
26	eta_ctl2	Grell	Thompson	Janjic Eta	MYJ
27	eta_ctl2	Grell	WSM-6	Janjic Eta	MYJ
28	eta_ctl2	Grell	WSM-6	Monin-Obukhov	YSU
29	eta_ctl2	Grell	Thompson	Monin-Obukhov	YSU
30	eta_ctl2	Grell	Ferrier	Monin-Obukhov	YSU

Table 5 Summary of results. For each variable, the different columns indicate whether faster spread growth, larger mixed-physics spread contributions, or better statistical consistency was observed for the 4- or 20-km grid-spacing ensemble configurations. Only statistically significant differences are indicated.

	Faster spread growth		Larger mixed-physics spread contribution		Better statistical consistency at forecast hour 33
	IC/LBC+Phys	Phys	Forecast hour 9	Forecast hour 33	
500Z	4-km	20-km	20-km	20-km	-----
700Z	4-km	20-km	20-km	20-km	-----
850Z	4-km	20-km	20-km	20-km	4-km
MSLP	4-km	20-km	-----	20-km	4-km
T2	-----	4-km	-----	4-km	-----
Td2	4-km	4-km	4-km	4-km	-----
850WMAG	-----	-----	-----	-----	4-km
850T	-----	-----	-----	-----	-----
850Td	-----	-----	20-km	20-km	-----
PREC	4-km	4-km	-----	-----	4-km
MUCAPE	-----	-----	20-km	-----	4-km
WSHR	4-km	4-km	-----	-----	4-km