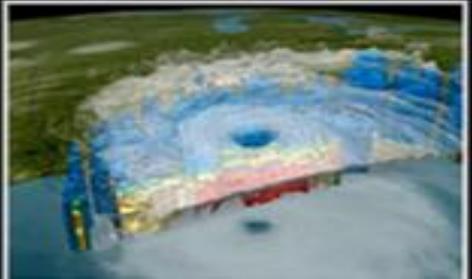
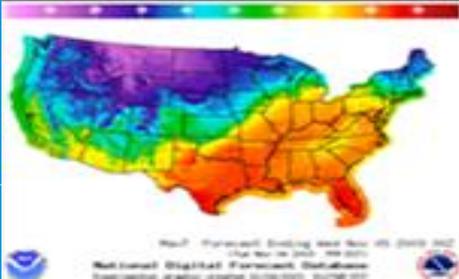
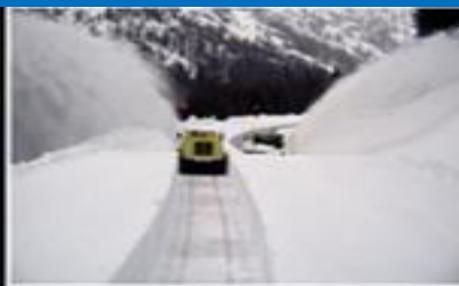
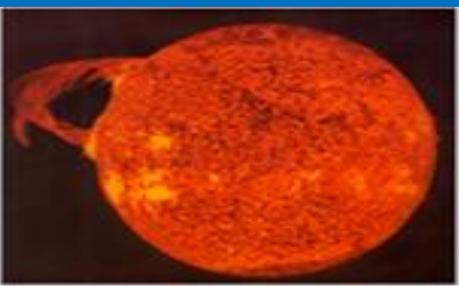
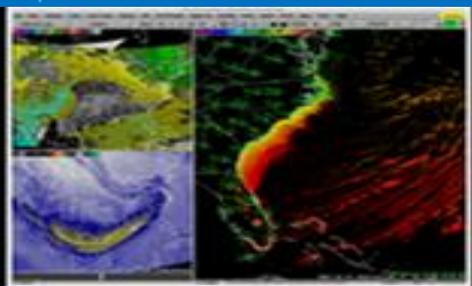




NOAA
National
Weather
Service

An overview on Mesoscale Convective Systems and their forward motion

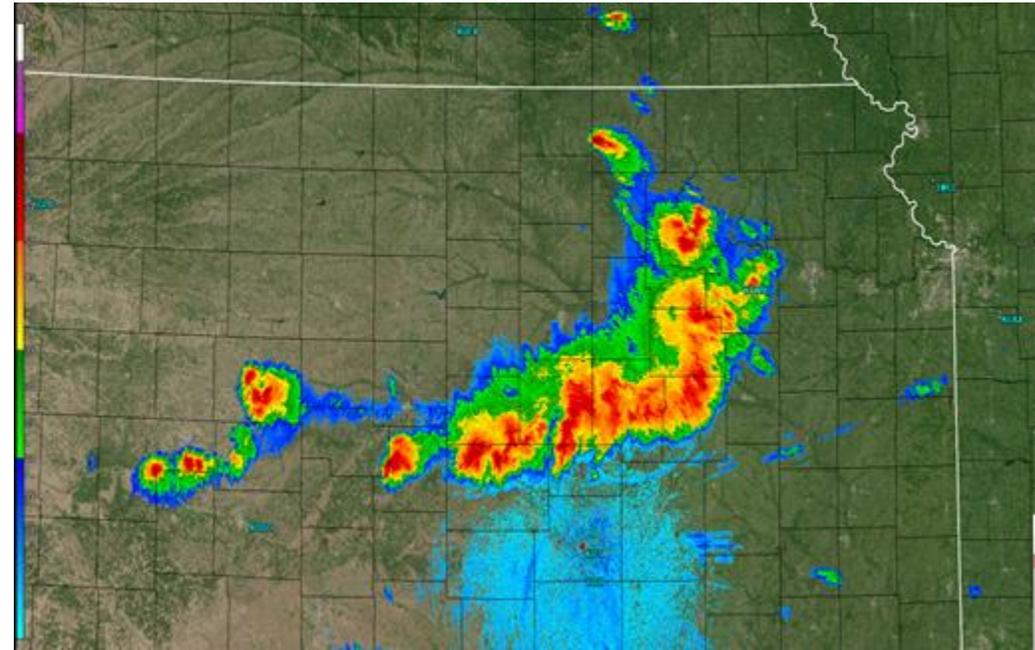
Brian Squitieri, Ph.D.
Storm Prediction Center



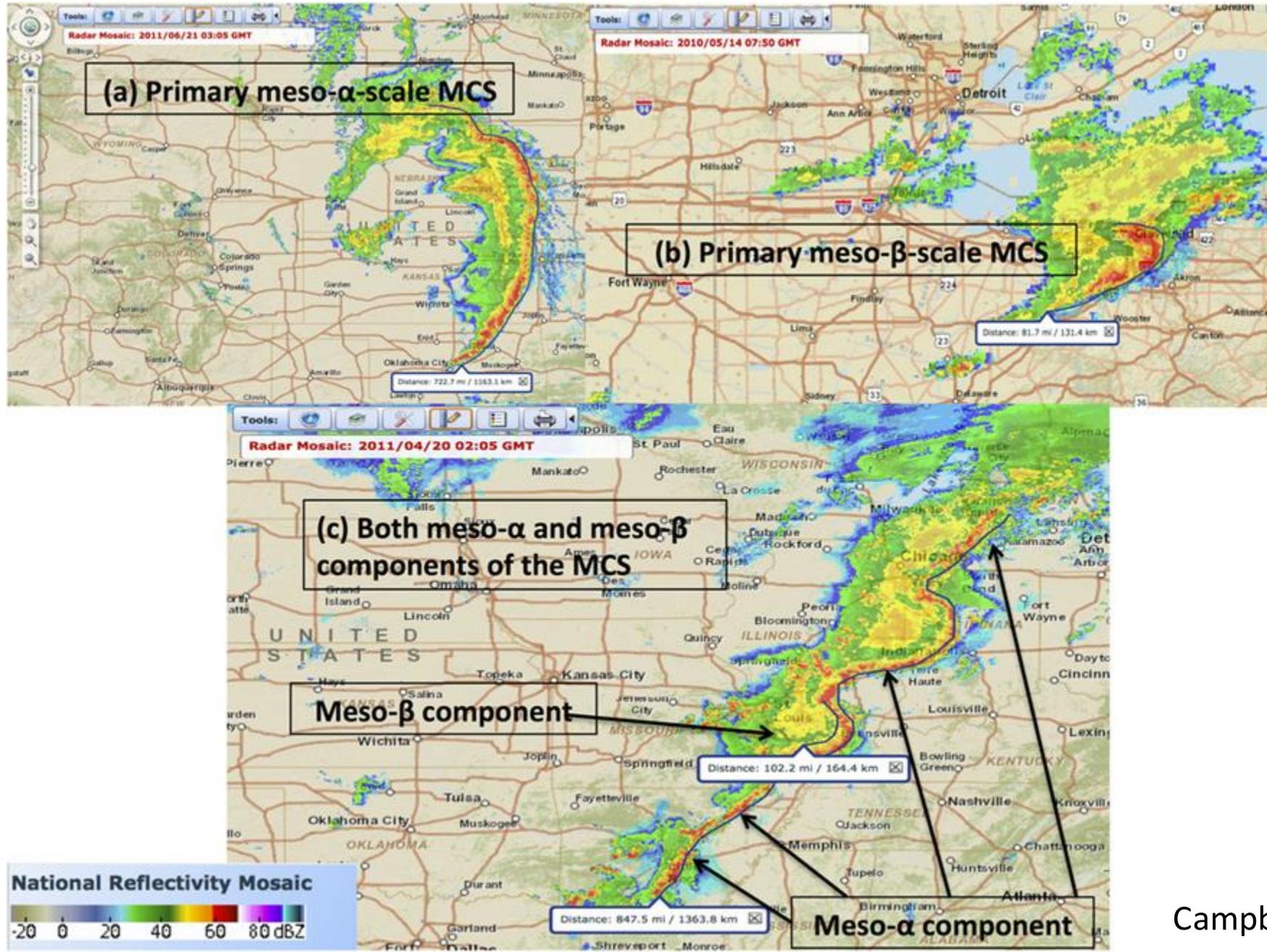
What is an MCS?

A **Mesoscale Convective System (MCS)** is “a cloud system that occurs in connection with an ensemble of thunderstorms and produces a contiguous precipitation area on the order of 100 km or more in horizontal scale in at least one direction” (AMS Glossary – 2025).

- This definition can be applied liberally to any organized thunderstorm clusters that:
 - Is at least 100 km long
 - Lasts for at least 3 hours
 - Shares a common feature, such as a trailing precipitation region or cold pool.
- MCSs may take on many forms, morphology and evolution
- MCSs may be accompanied by all thunderstorm hazards:
 - Heavy rain and potential flooding
 - Frequent Lightning
 - Strong, damaging winds/gusts
 - Tornadoes
 - Hail



MCS Types



MCSs on the spatial and temporal spectrum

MCSs on the spatial and temporal spectrum

- Parker and Johnson (2000): MCS defined as convective phenomenon where Coriolis acceleration is the same order of magnitude as all other terms in Navier-Stokes equations of motion.

MCSs on the spatial and temporal spectrum

- Parker and Johnson (2000): MCS defined as convective phenomenon where Coriolis acceleration is the same order of magnitude as all other terms in Navier-Stokes equations of motion.

$$\frac{du}{dt} = -\frac{1}{\rho} \frac{\partial p}{\partial x} + fv + F_u$$

MCSs on the spatial and temporal spectrum

- Parker and Johnson (2000): MCS defined as convective phenomenon where Coriolis acceleration is the same order of magnitude as all other terms in Navier-Stokes equations of motion.

$$\frac{du}{dt} = -\frac{1}{\rho} \frac{\partial p}{\partial x} + \boxed{fv} + F_u$$

MCSs on the spatial and temporal spectrum

- Parker and Johnson (2000): MCS defined as convective phenomenon where Coriolis acceleration is the same order of magnitude as all other terms in Navier-Stokes equations of motion.

$$\frac{du}{dt} = -\frac{1}{\rho} \frac{\partial p}{\partial x} + \boxed{fv} + F_u$$

- Markowski and Richardson (2010): Most Phenomena have space/time scale ratios on the same order of magnitude (10 m s^{-1})

MCSs on the spatial and temporal spectrum

- Parker and Johnson (2000): MCS defined as convective phenomenon where Coriolis acceleration is the same order of magnitude as all other terms in Navier-Stokes equations of motion.

$$\frac{du}{dt} = -\frac{1}{\rho} \frac{\partial p}{\partial x} + \boxed{fv} + F_u$$

- Markowski and Richardson (2010): Most Phenomena have space/time scale ratios on the same order of magnitude (10 m s^{-1})
- In mid-latitudes: $f = 2\Omega \sin\Phi \sim 10^{-4} \text{ s}^{-1}$, so $fv = (10^{-4} \text{ s}^{-1}) * (10 \text{ m s}^{-1}) = 10^{-3} \text{ m s}^{-2}$

MCSs on the spatial and temporal spectrum

- Parker and Johnson (2000): MCS defined as convective phenomenon where Coriolis acceleration is the same order of magnitude as all other terms in Navier-Stokes equations of motion.

$$\frac{du}{dt} = -\frac{1}{\rho} \frac{\partial p}{\partial x} + \boxed{fv} + F_u$$

- Markowski and Richardson (2010): Most Phenomena have space/time scale ratios on the same order of magnitude (10 m s^{-1})
- In mid-latitudes: $f = 2\Omega \sin\Phi \sim 10^{-4} \text{ s}^{-1}$, so $fv = (10^{-4} \text{ s}^{-1}) * (10 \text{ m s}^{-1}) = 10^{-3} \text{ m s}^{-2}$
- In mid-latitudes, $L < 1000 \text{ km}$ means that Coriolis becomes less relevant, $L > 1000 \text{ km}$ and ageostrophic motions become less significant, so:

MCSs on the spatial and temporal spectrum

- Parker and Johnson (2000): MCS defined as convective phenomenon where Coriolis acceleration is the same order of magnitude as all other terms in Navier-Stokes equations of motion.

$$\frac{du}{dt} = -\frac{1}{\rho} \frac{\partial p}{\partial x} + \boxed{fv} + F_u$$

- Markowski and Richardson (2010): Most Phenomena have space/time scale ratios on the same order of magnitude (10 m s^{-1})
- In mid-latitudes: $f = 2\Omega \sin\Phi \sim 10^{-4} \text{ s}^{-1}$, so $fv = (10^{-4} \text{ s}^{-1}) * (10 \text{ m s}^{-1}) = 10^{-3} \text{ m s}^{-2}$
- In mid-latitudes, $L < 1000 \text{ km}$ means that Coriolis becomes less relevant, $L > 1000 \text{ km}$ and ageostrophic motions become less significant, so:
- In mid-latitudes: $-\frac{1}{\rho} \frac{\partial p}{\partial x} = \frac{1}{1 \text{ kg m}^{-3}} \frac{10 \text{ mb}}{1000 \text{ km}} \sim 10^{-3} \text{ m s}^{-2}$, with F_u also 10^{-3} m s^{-2}

MCSs on the spatial and temporal spectrum

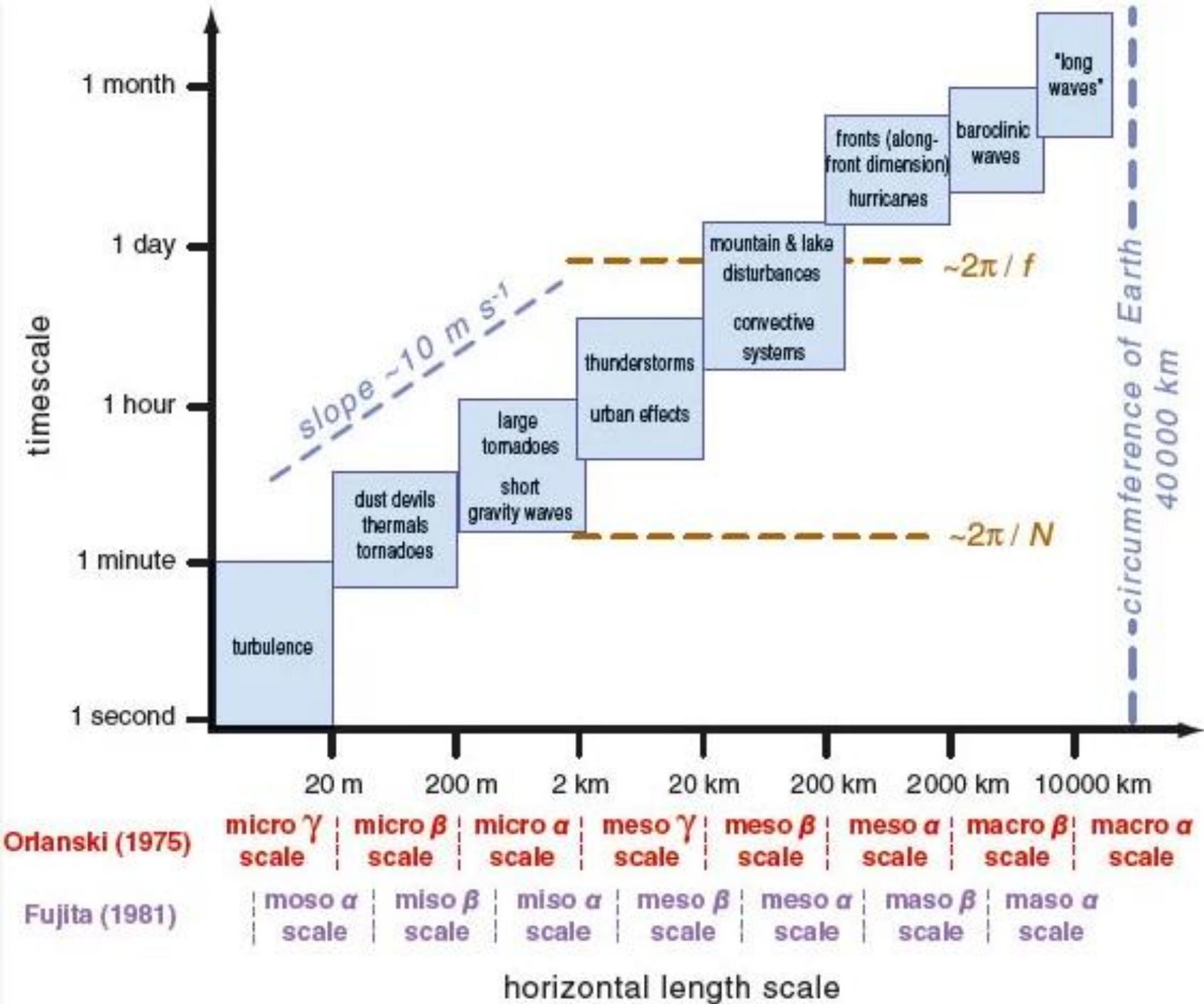
- Parker and Johnson (2000): If Coriolis is equal to other forces in Navier-Stokes, then the timescale should be f^{-1} .

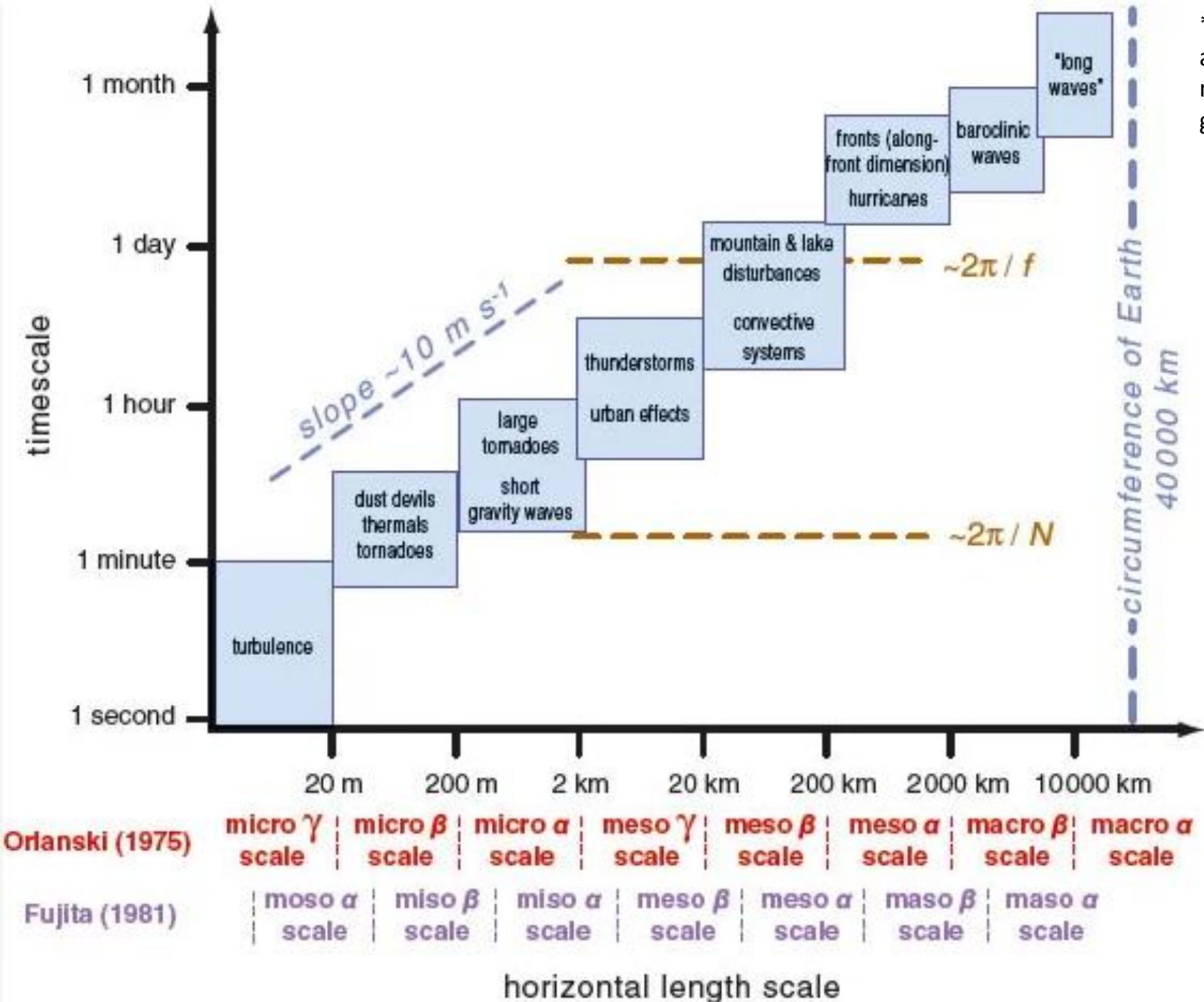
MCSs on the spatial and temporal spectrum

- Parker and Johnson (2000): If Coriolis is equal to other forces in Navier-Stokes, then the timescale should be f^{-1} .
- This works out to: $\frac{1}{f} = \frac{1}{2\Omega \sin\Phi} \sim \frac{1}{10^{-4} s^{-1}} \sim 10^4 s$, or 10,000 s (divided by 3600 s) is 2.77 or **3 h**.

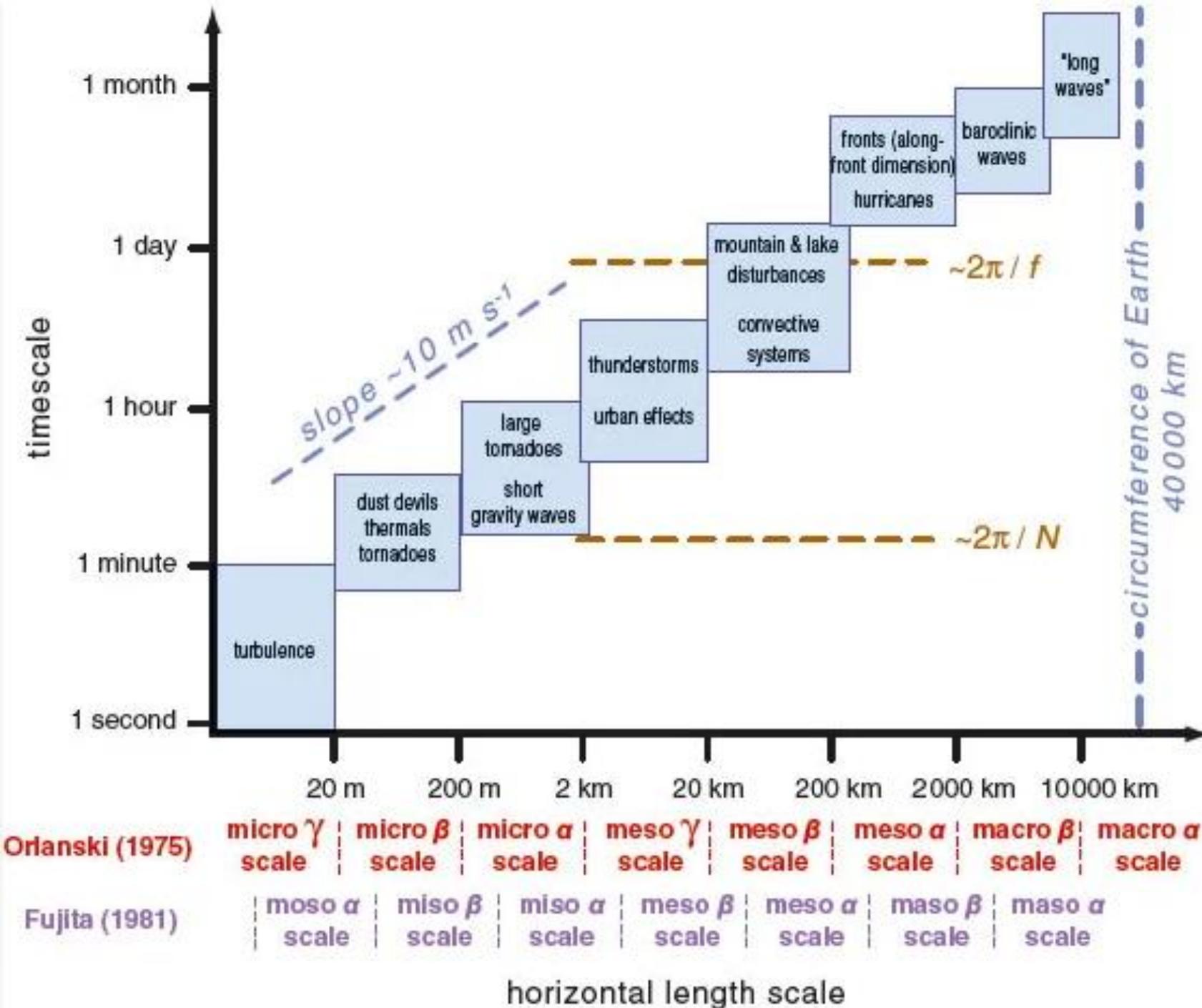
MCSs on the spatial and temporal spectrum

- Parker and Johnson (2000): If Coriolis is equal to other forces in Navier-Stokes, then the timescale should be f^{-1} .
- This works out to: $\frac{1}{f} = \frac{1}{2\Omega \sin\Phi} \sim \frac{1}{10^{-4} s^{-1}} \sim 10^4 s$, or 10,000 s (divided by 3600 s) is 2.77 or **3 h**.
- Parker and Johnson (2000): Making an advective assumption, length scale is $L = U\tau = 10 m s^{-1} * 10^4 s = 10^5 m$ or **100 km**.



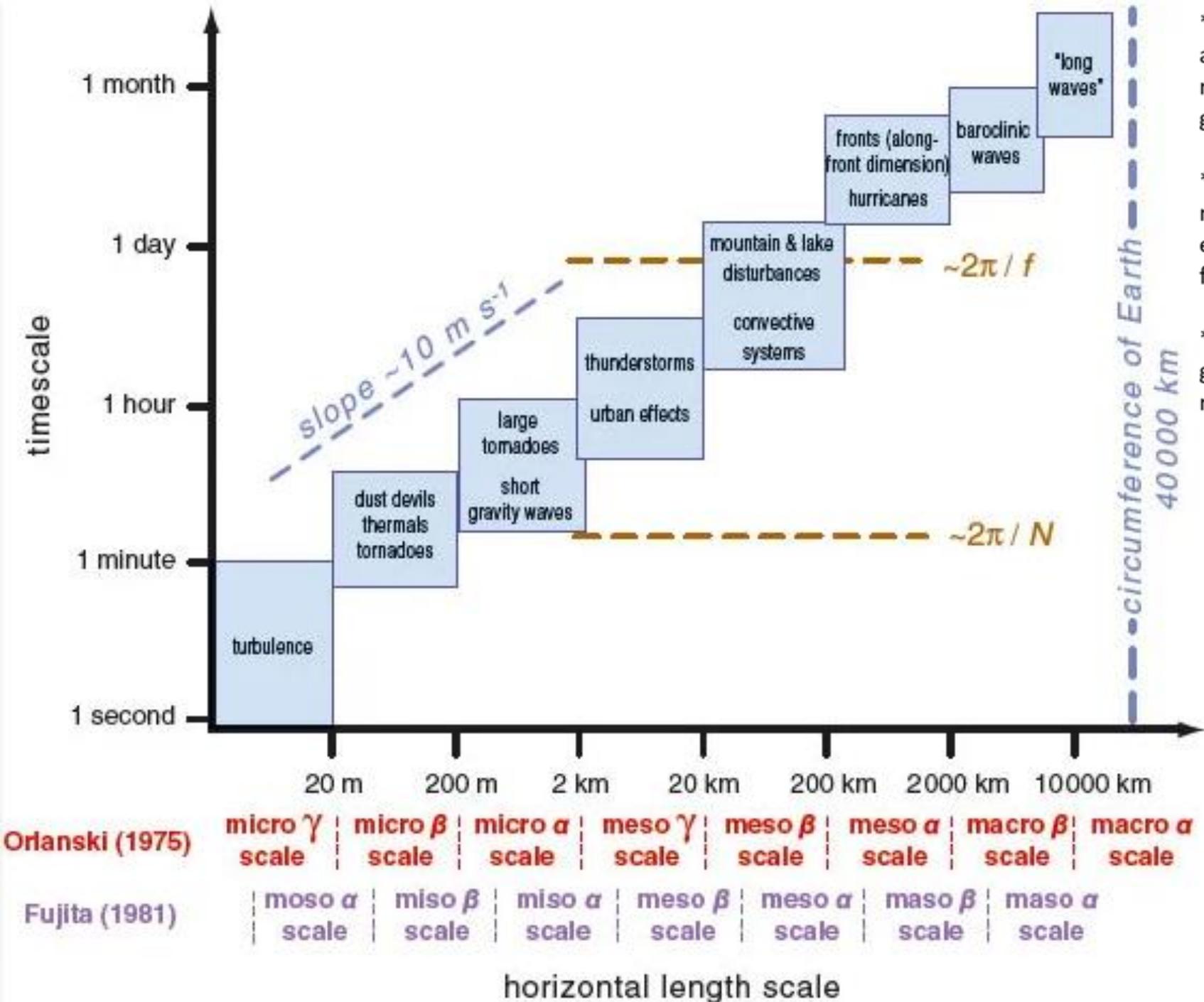


*Synoptic-scales allows for neglecting vertical accelerations and geostrophic wind advection, and micro-scale allows for Coriolis and horizontal pressure gradient forces to be neglected.



*Synoptic-scales allows for neglecting vertical accelerations and geostrophic wind advection, and micro-scale allows for Coriolis and horizontal pressure gradient forces to be neglected.

*Mesoscale doesn't allow for these forces to be neglected, and Markowski and Richardson (2010) explicitly cite long-lived MCSs as an example where all forces need to be accounted for.

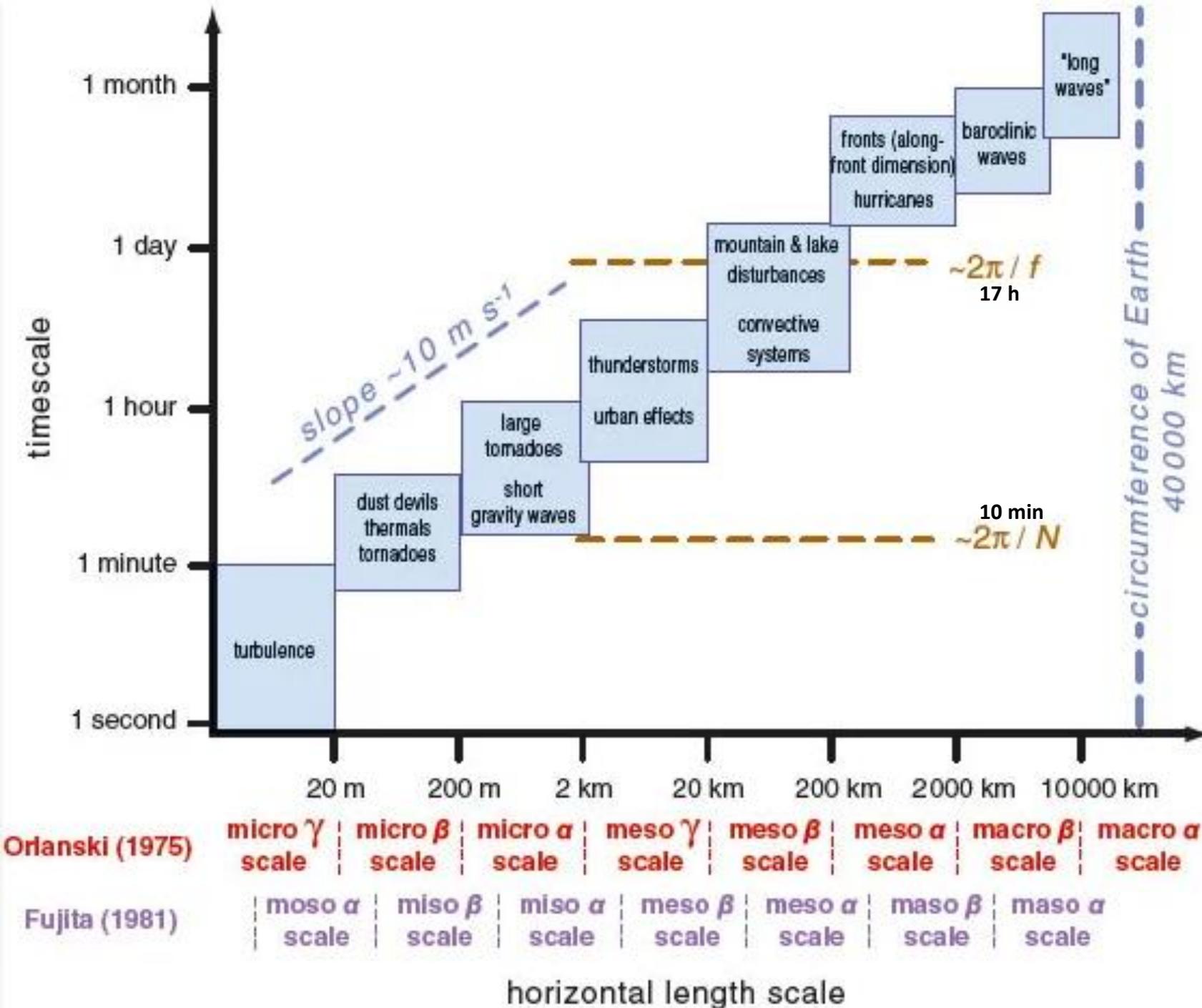


*Synoptic-scales allows for neglecting vertical accelerations and geostrophic wind advection, and micro-scale allows for Coriolis and horizontal pressure gradient forces to be neglected.

*Mesoscale doesn't allow for these forces to be neglected, and Markowski and Richardson (2010) explicitly cite long-lived MCSs as an example where all forces need to be accounted for.

*Synoptic phenomena in mid-latitudes are usually generated by baroclinic instability, which is usually realized by disturbances with 3X the wavelengths of:

$$L_R = \frac{NH}{f} \sim 1000 \text{ to } 1500 \text{ km}$$



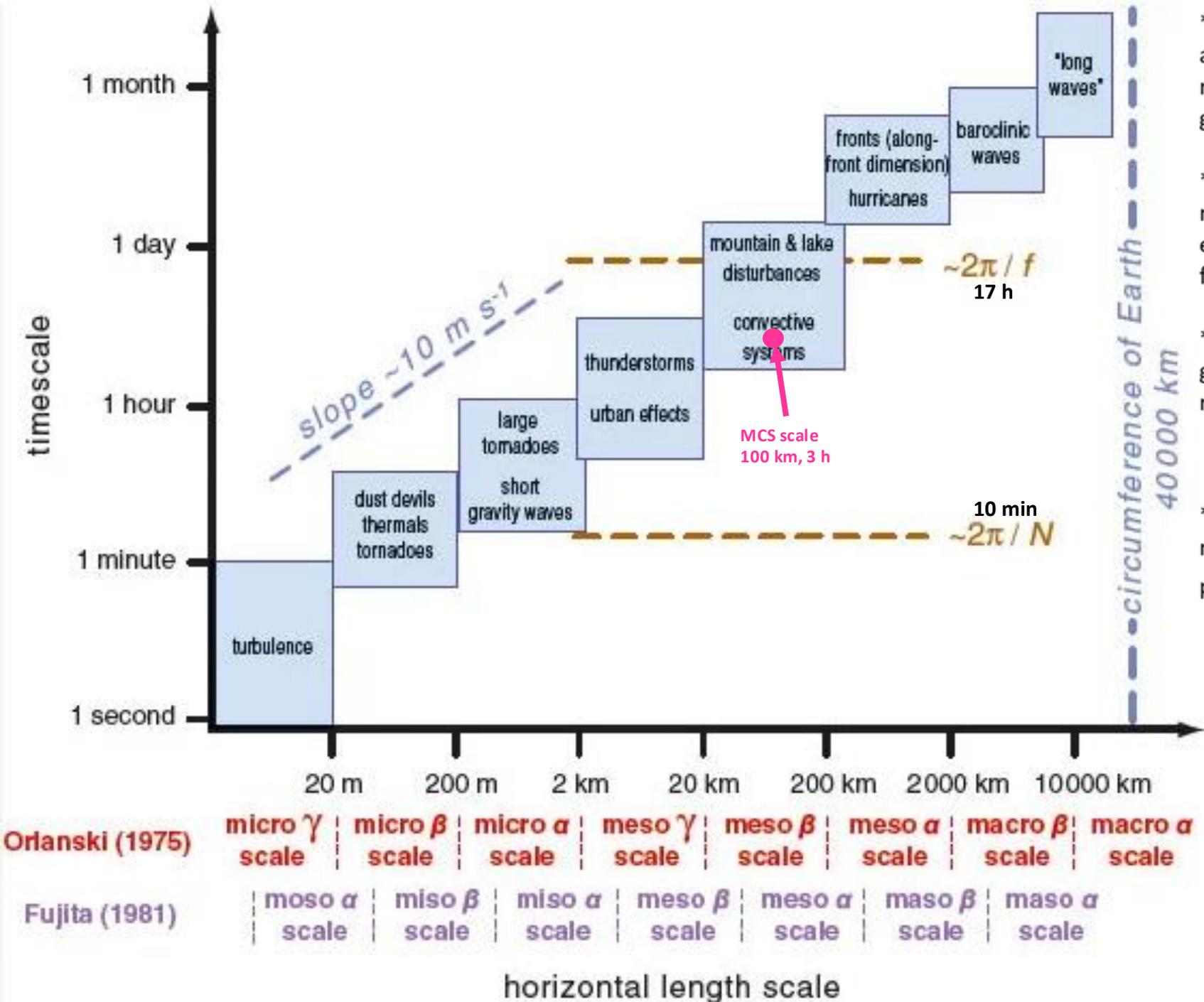
*Synoptic-scales allows for neglecting vertical accelerations and geostrophic wind advection, and micro-scale allows for Coriolis and horizontal pressure gradient forces to be neglected.

*Mesoscale doesn't allow for these forces to be neglected, and Markowski and Richardson (2010) explicitly cite long-lived MCSs as an example where all forces need to be accounted for.

*Synoptic phenomena in mid-latitudes are usually generated by baroclinic instability, which is usually realized by disturbances with 3X the wavelengths of:

$$L_R = \frac{NH}{f} \sim 1000 \text{ to } 1500 \text{ km}$$

*Markowski and Richardson: Lagrangian timescales of mesoscale fall between buoyancy oscillation and pendulum day scales, so: $\frac{2\pi}{N} < \tau < \frac{2\pi}{f}$ or $10 \text{ min} < \tau < 17 \text{ h}$



*Synoptic-scales allows for neglecting vertical accelerations and geostrophic wind advection, and micro-scale allows for Coriolis and horizontal pressure gradient forces to be neglected.

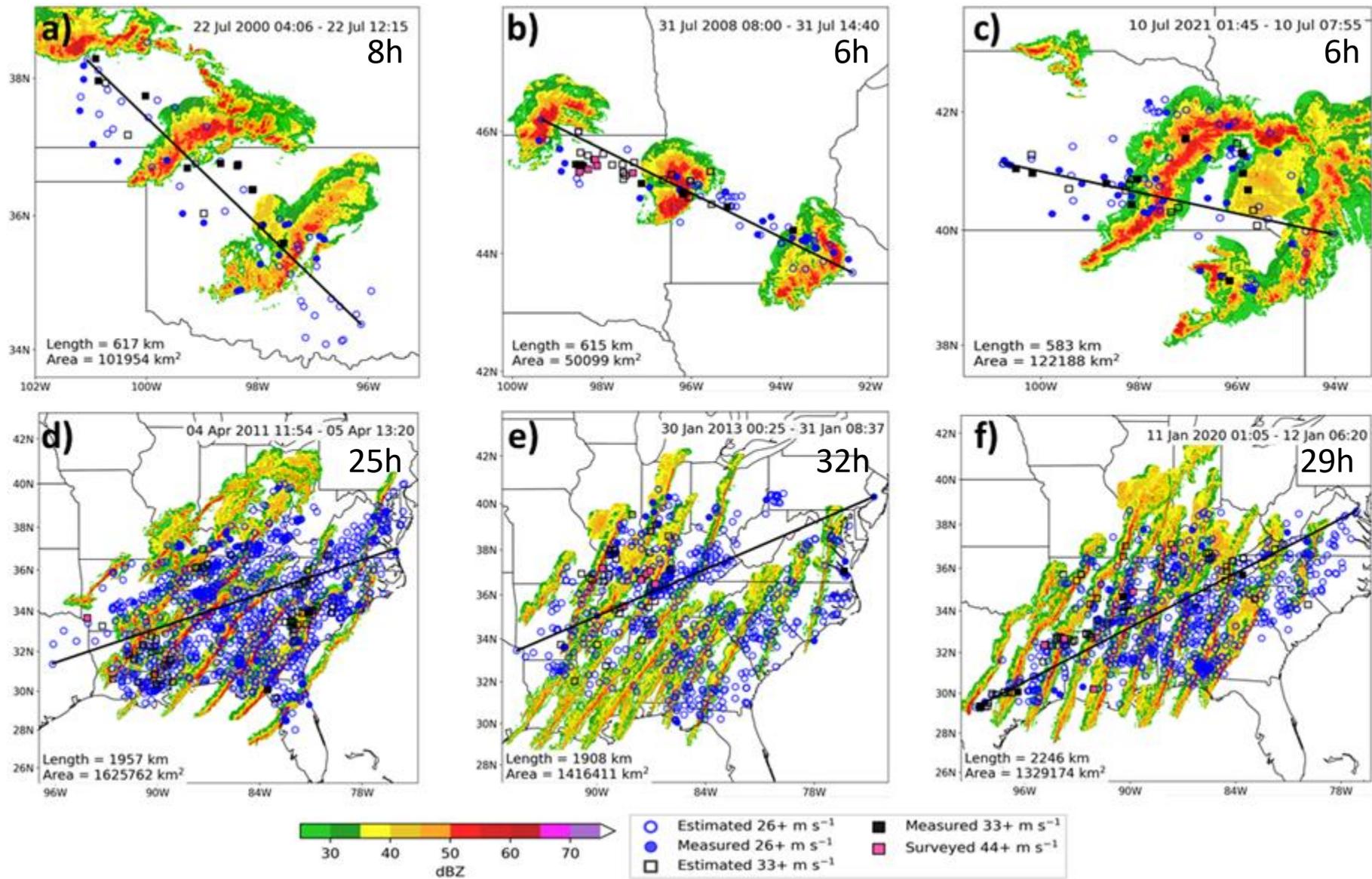
*Mesoscale doesn't allow for these forces to be neglected, and Markowski and Richardson (2010) explicitly cite long-lived MCSs as an example where all forces need to be accounted for.

*Synoptic phenomena in mid-latitudes are usually generated by baroclinic instability, which is usually realized by disturbances with 3X the wavelengths of:

$$L_R = \frac{NH}{f} \sim 1000 \text{ to } 1500 \text{ km}$$

*Markowski and Richardson: Lagrangian timescales of mesoscale fall between buoyancy oscillation and pendulum day scales, so: $\frac{2\pi}{N} < \tau < \frac{2\pi}{f}$ or $10 \text{ min} < \tau < 17 \text{ h}$

Cold-pool-driven MCSs tend to differ in structure from synoptically forced squall lines, with differences in wind swath attributes



Cross Section



Rain-cooled
Air



Radar
Depiction



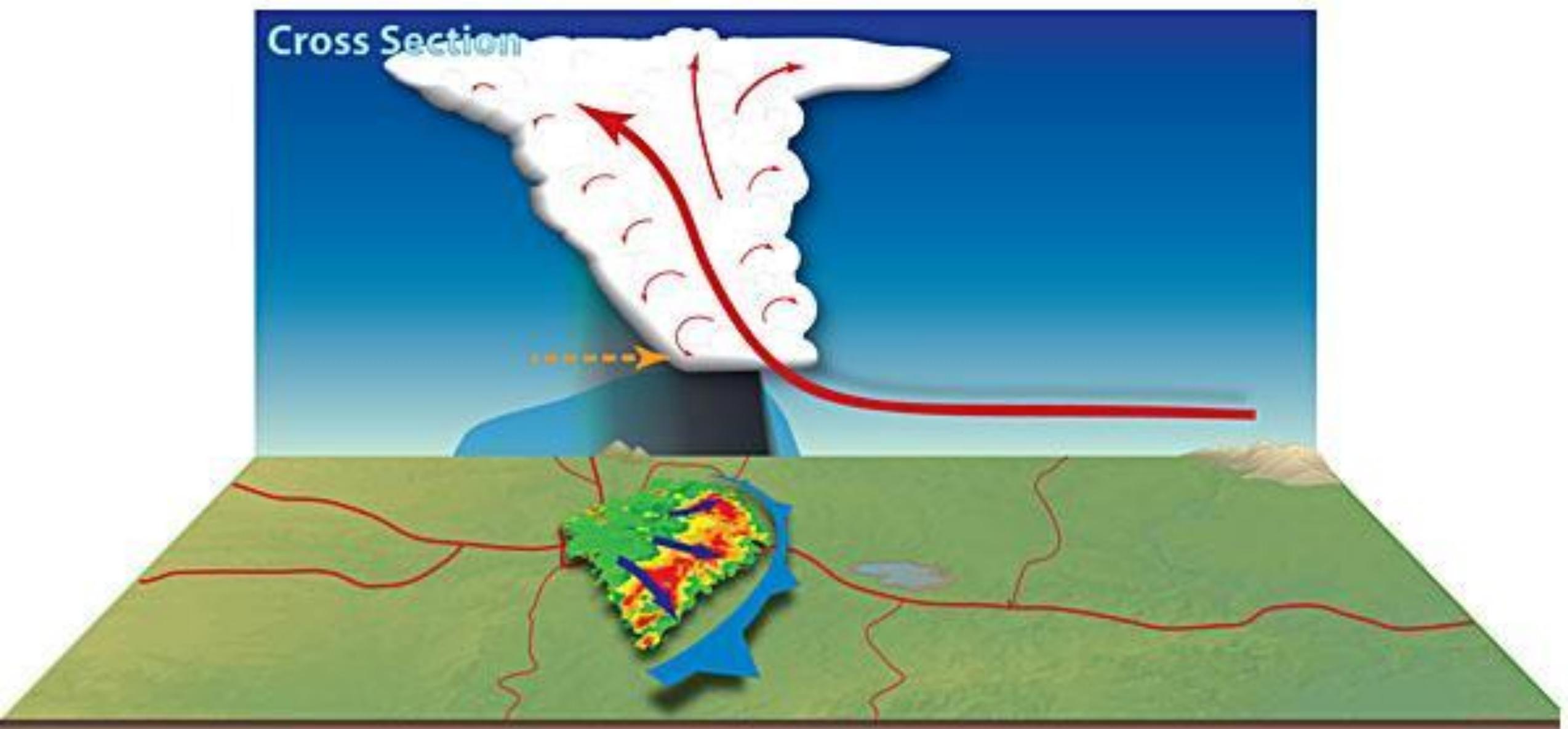
Gust
Front



Updraft



Cross Section



Rain-cooled
Air



Radar
Depiction



Gust
Front

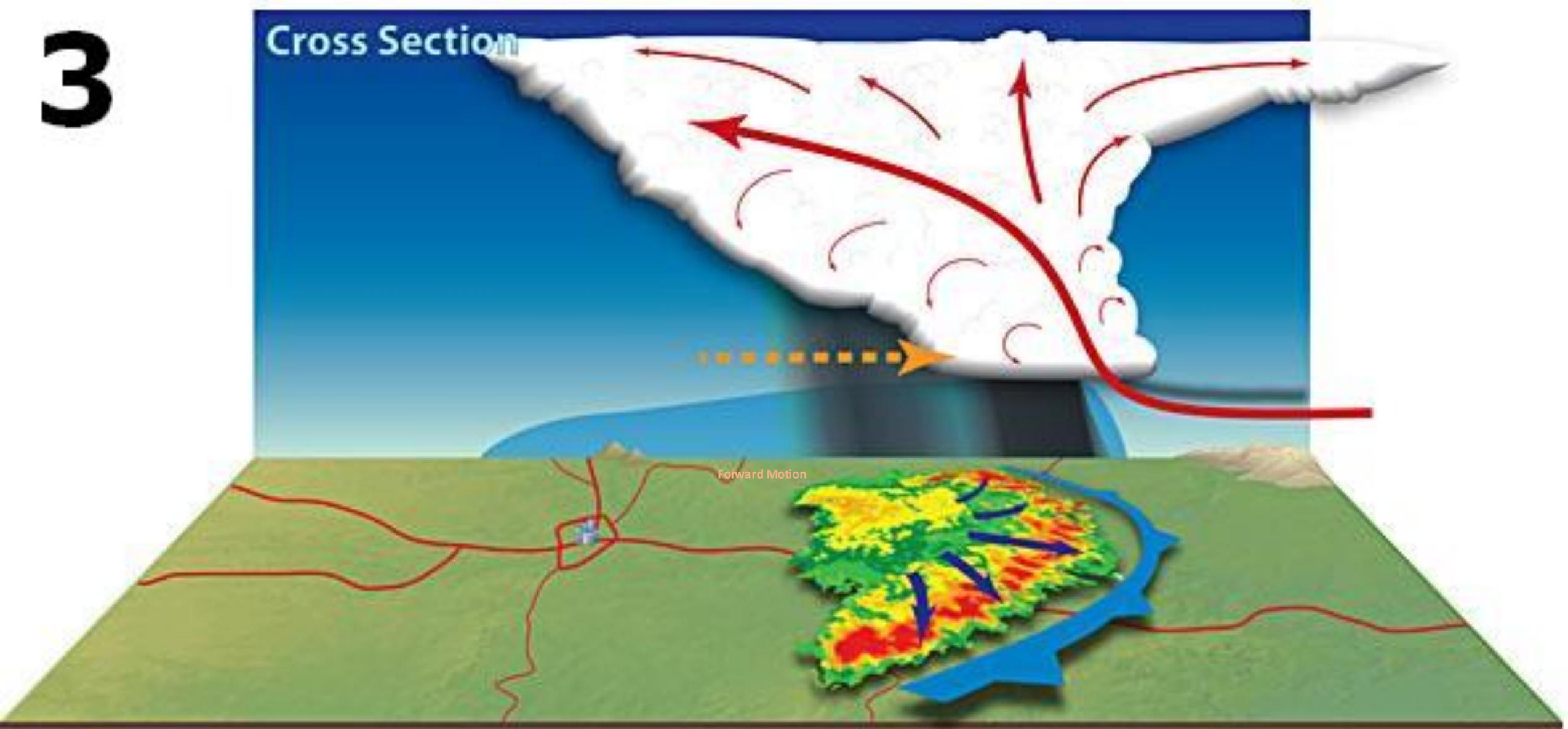


Updraft



3

Cross Section



Rain-cooled Air



Radar Depiction



Gust Front



Updraft



Cross Section

Secondary Upper
Circulation

Secondary Lower
Circulation

Forward Motion

Trailing Precipitation Region

MCS Leading Line

Rain-cooled
Air



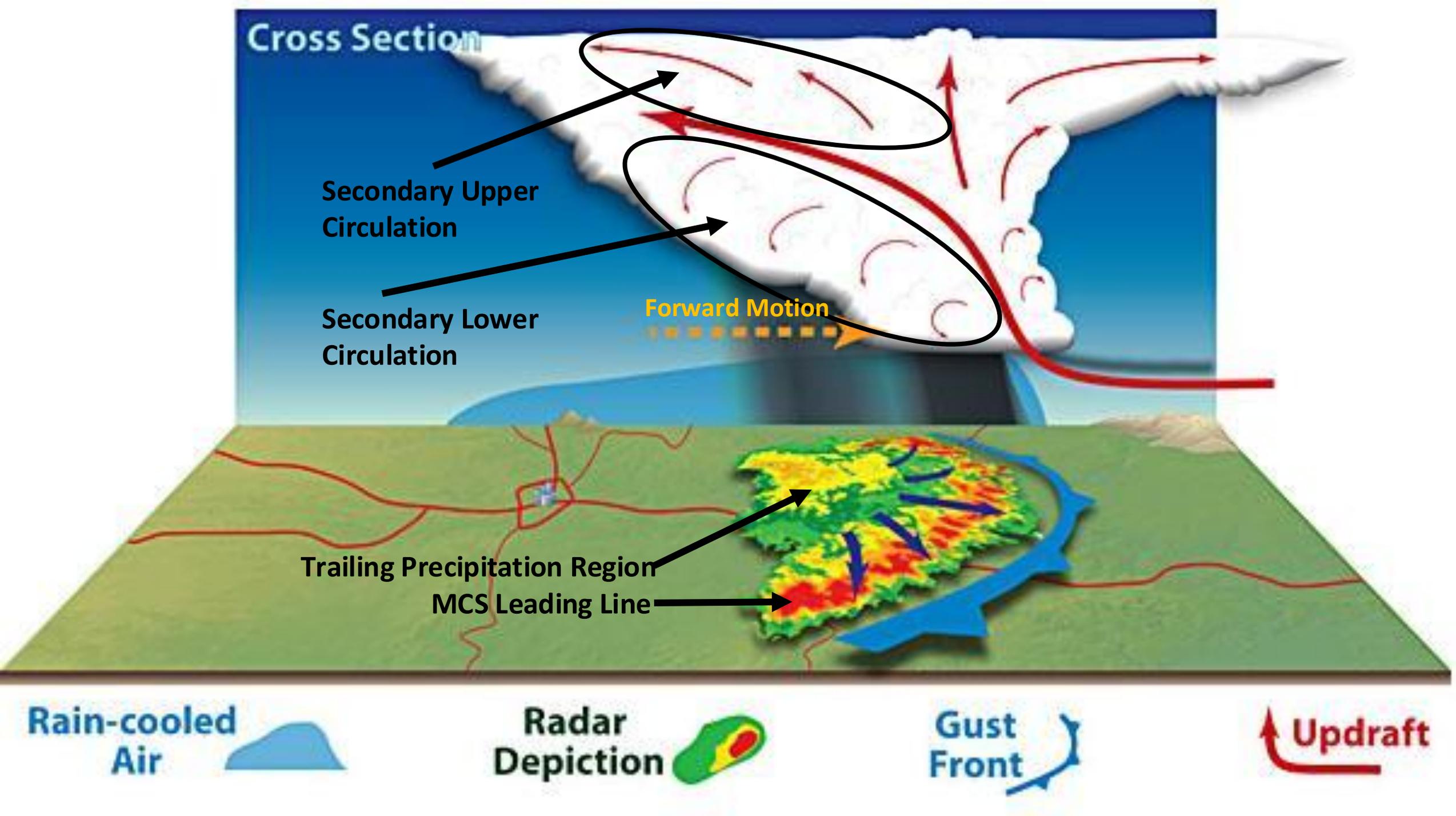
Radar
Depiction

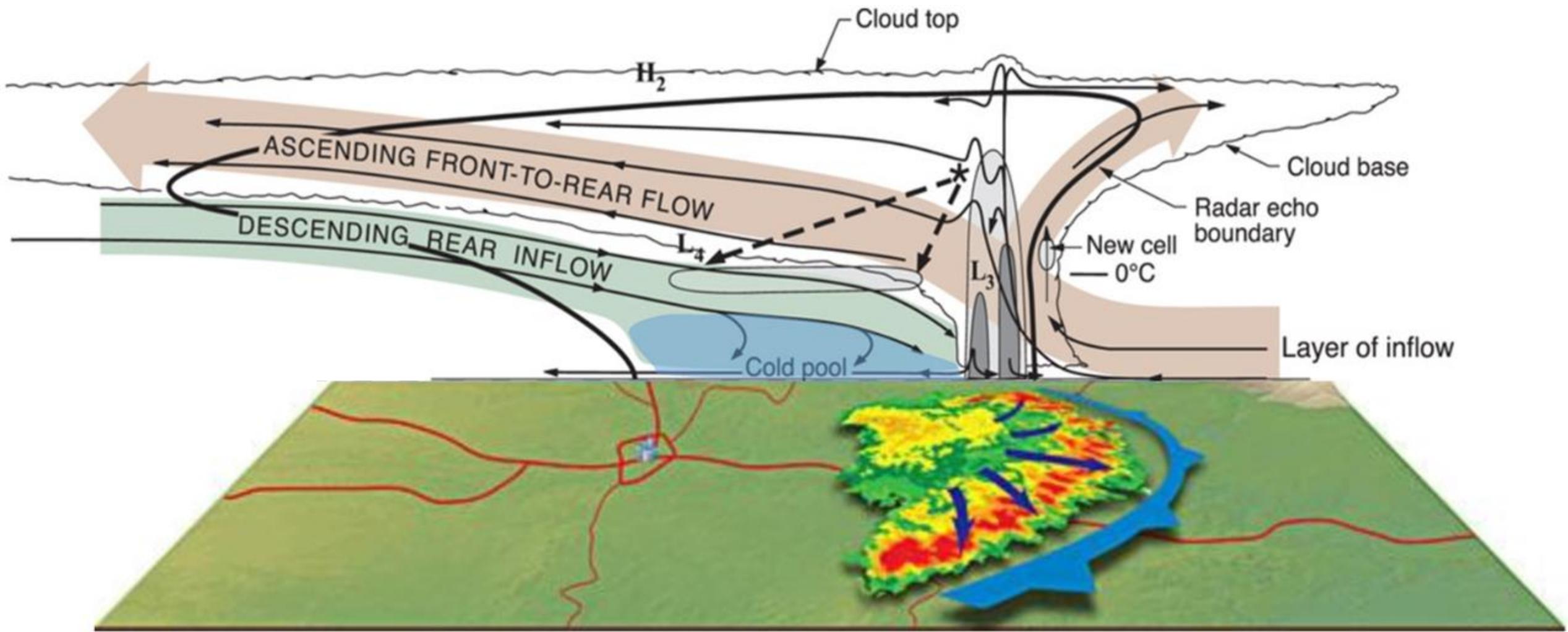


Gust
Front

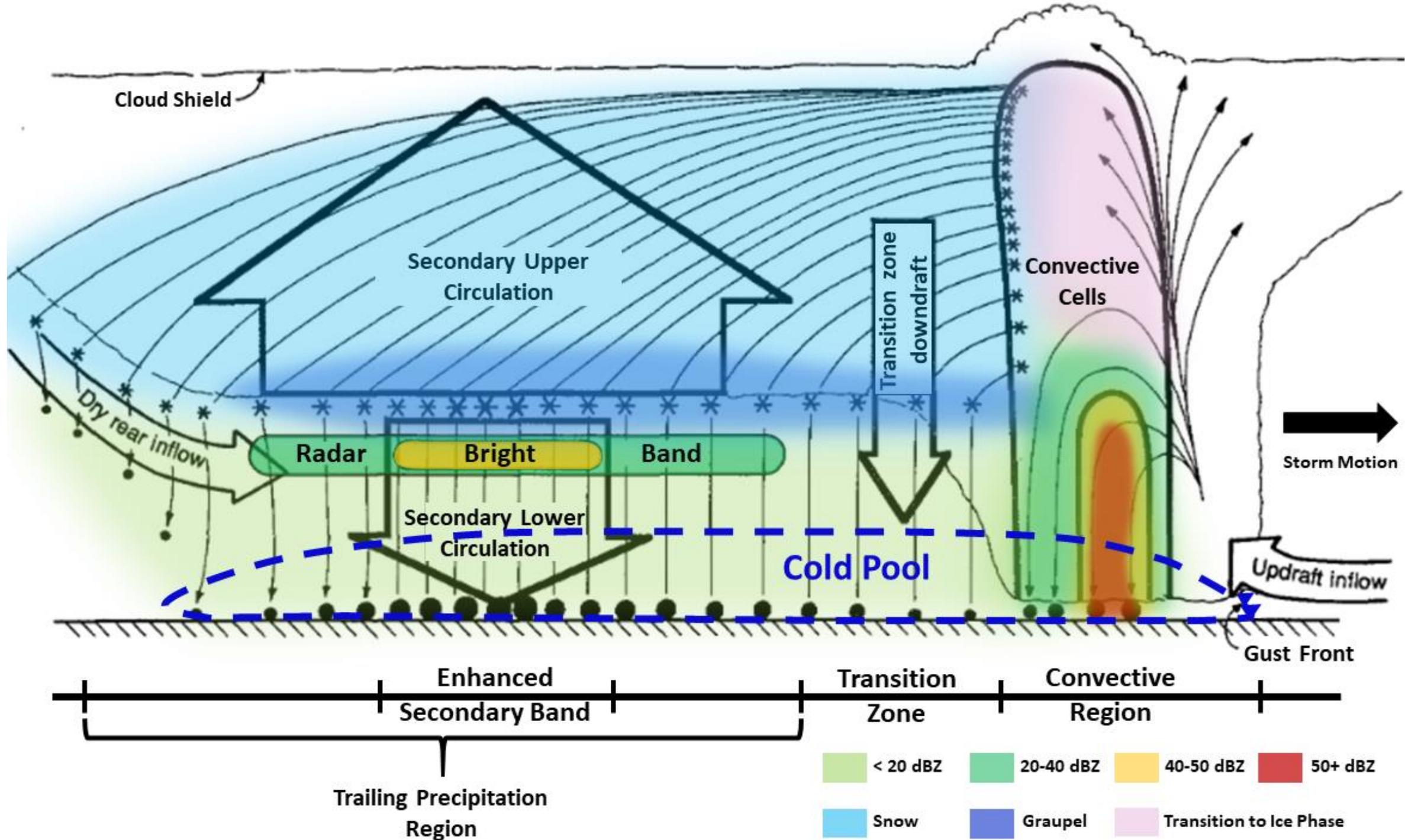


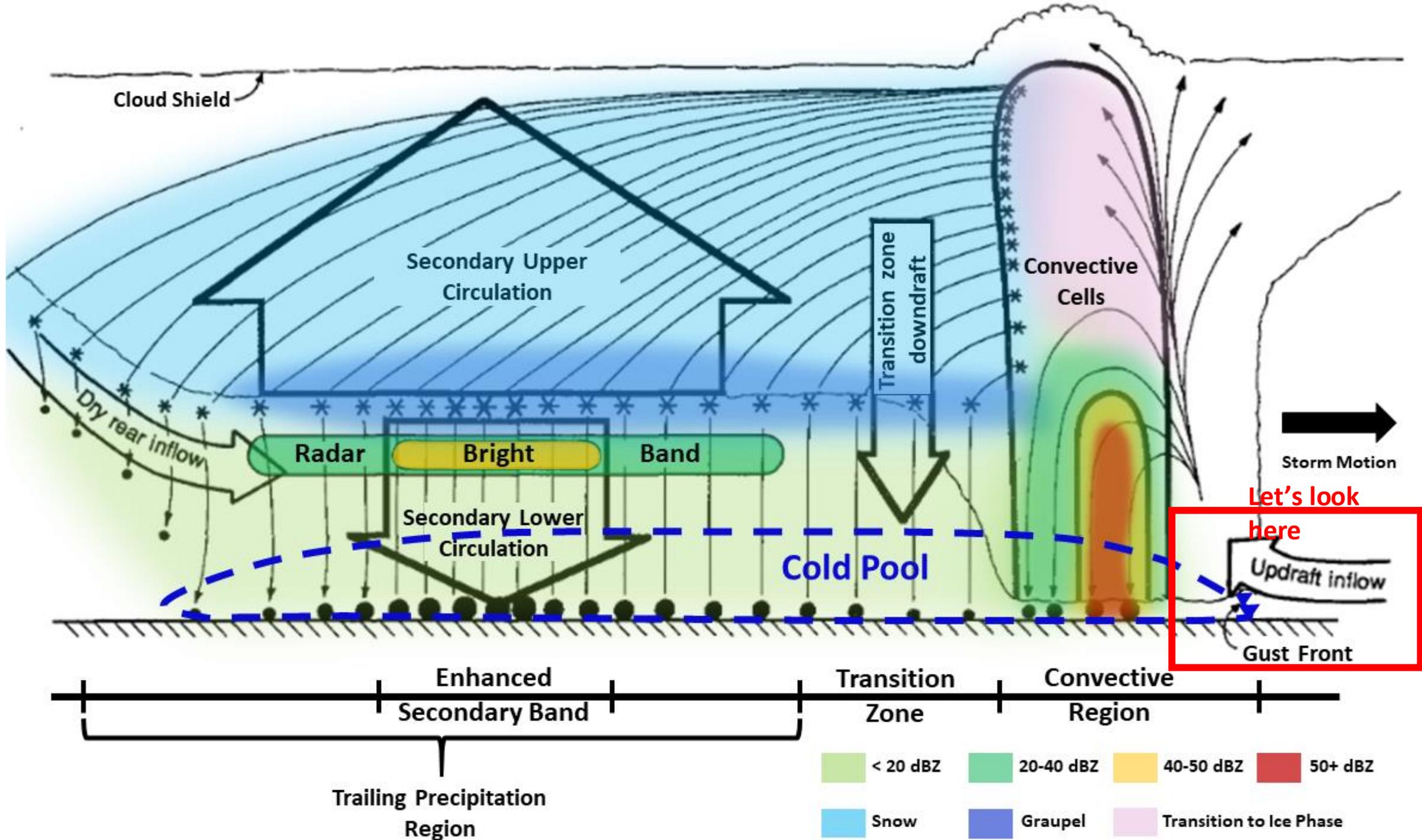
Updraft

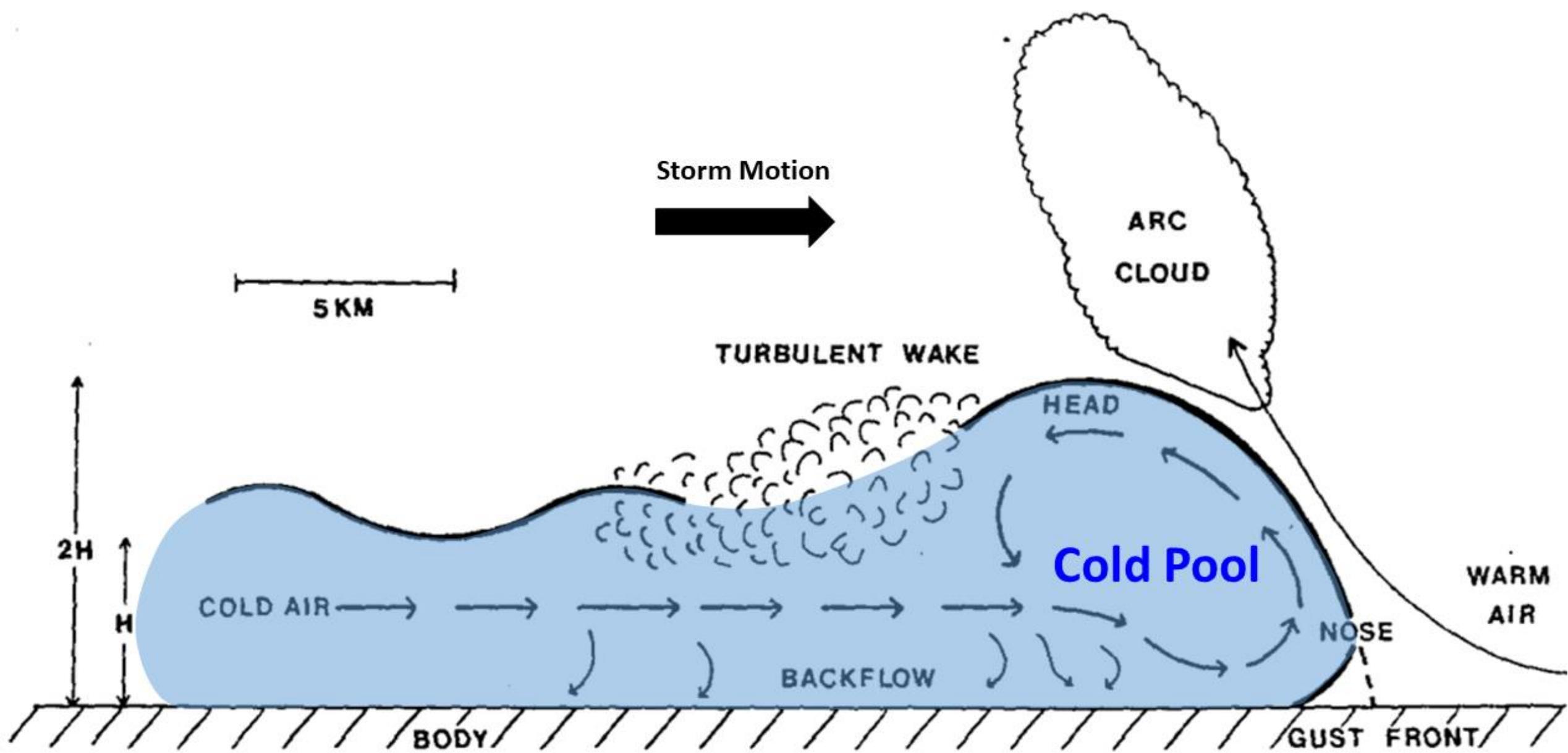




Ascending Air
 Rear Inflow Jet
 Cold Pool Air
 Radar Depiction
Gust Front







MCS Forward Motion

- Johns and Hirt (1987) and Corfidi et al. (2016) both found that derechos move faster (sometimes much faster) than the full mean wind speed.
- Derecho wind swaths are produced by thunderstorm clusters where either cold pool dynamics or other internal mechanisms dominate the processes that produce severe/destructive wind gusts.
- Corfidi et al. (1996) and Corfidi (2003) devised a routine that can determine MCS forward motion based on the interaction between the cold pool and ambient flow fields.

MCS Forward Motion

MCS Forward Motion

- MCS motion has 2 components (Advection and Propagation)

MCS Forward Motion

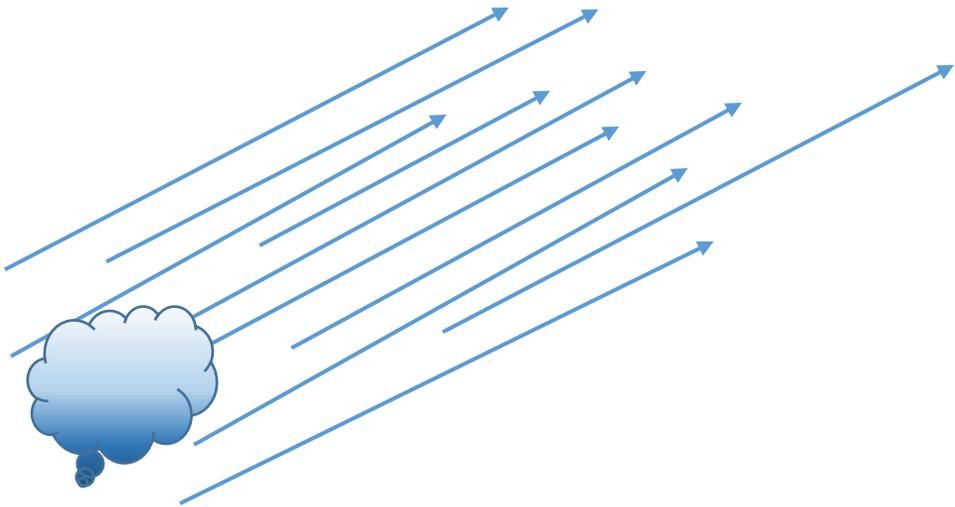
- Advection:

MCS Forward Motion

- Advection: Component of motion involving cells being carried with or by the mean wind in the cloud layer (V_{CL}).

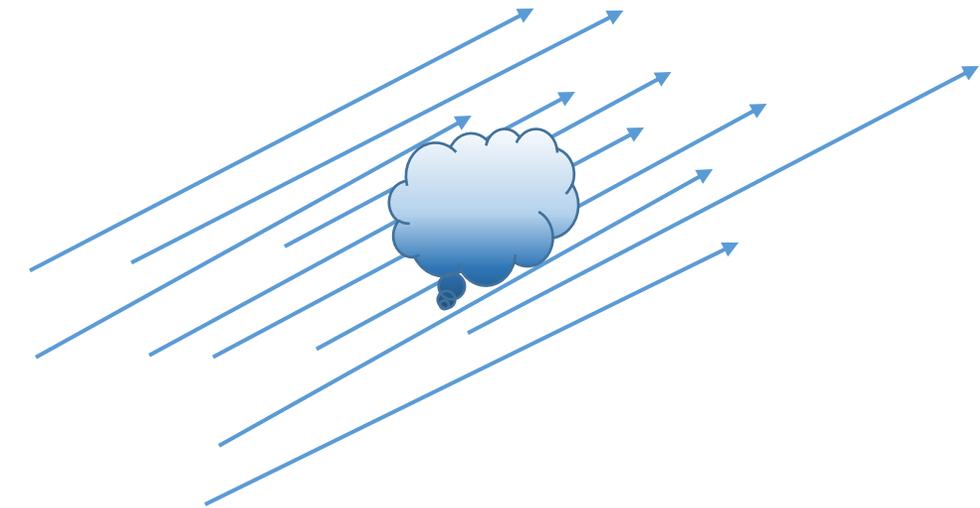
MCS Forward Motion

- **Advection**: Component of motion involving cells being carried with or by the mean wind in the cloud layer (V_{CL}).



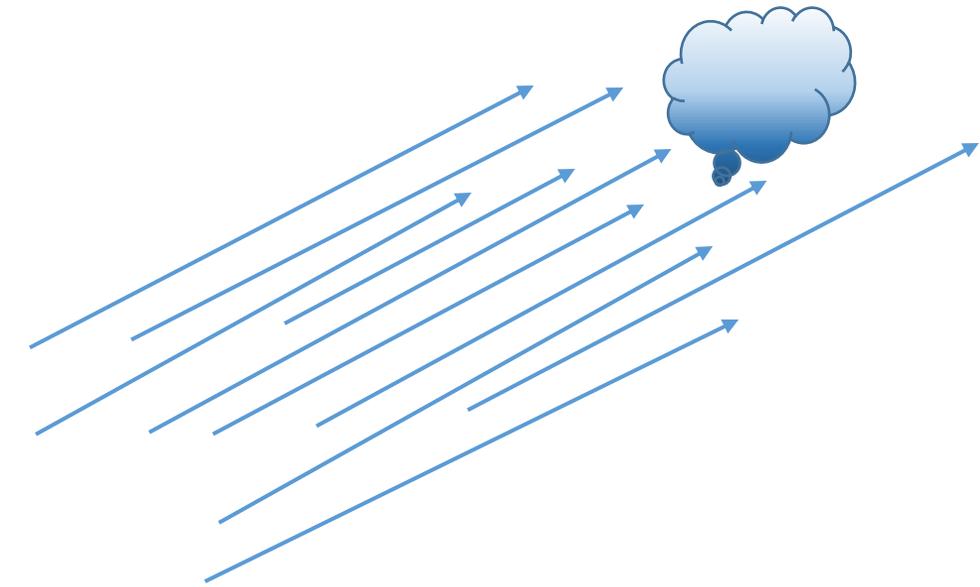
MCS Forward Motion

- **Advection**: Component of motion involving cells being carried with or by the mean wind in the cloud layer (V_{CL}).



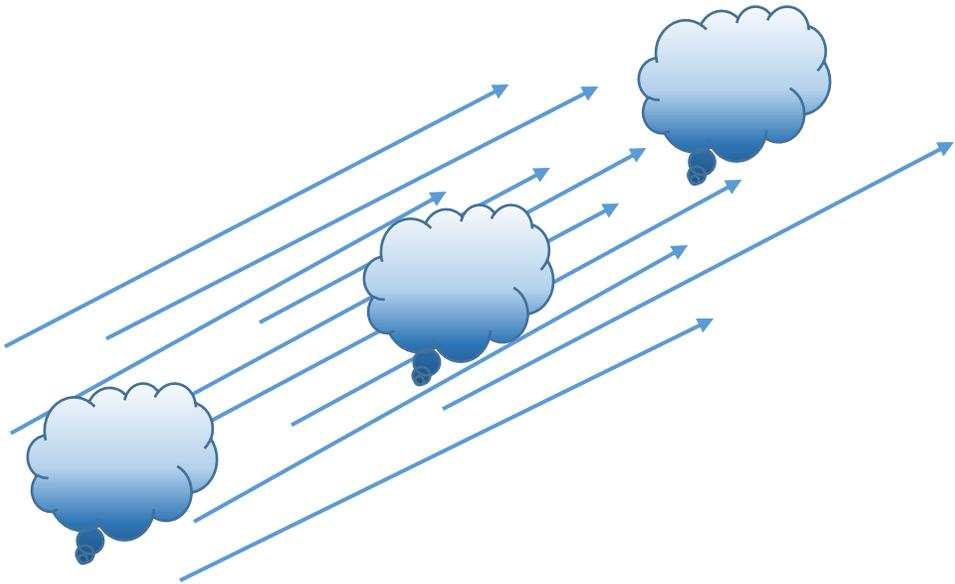
MCS Forward Motion

- **Advection**: Component of motion involving cells being carried with or by the mean wind in the cloud layer (V_{CL}).



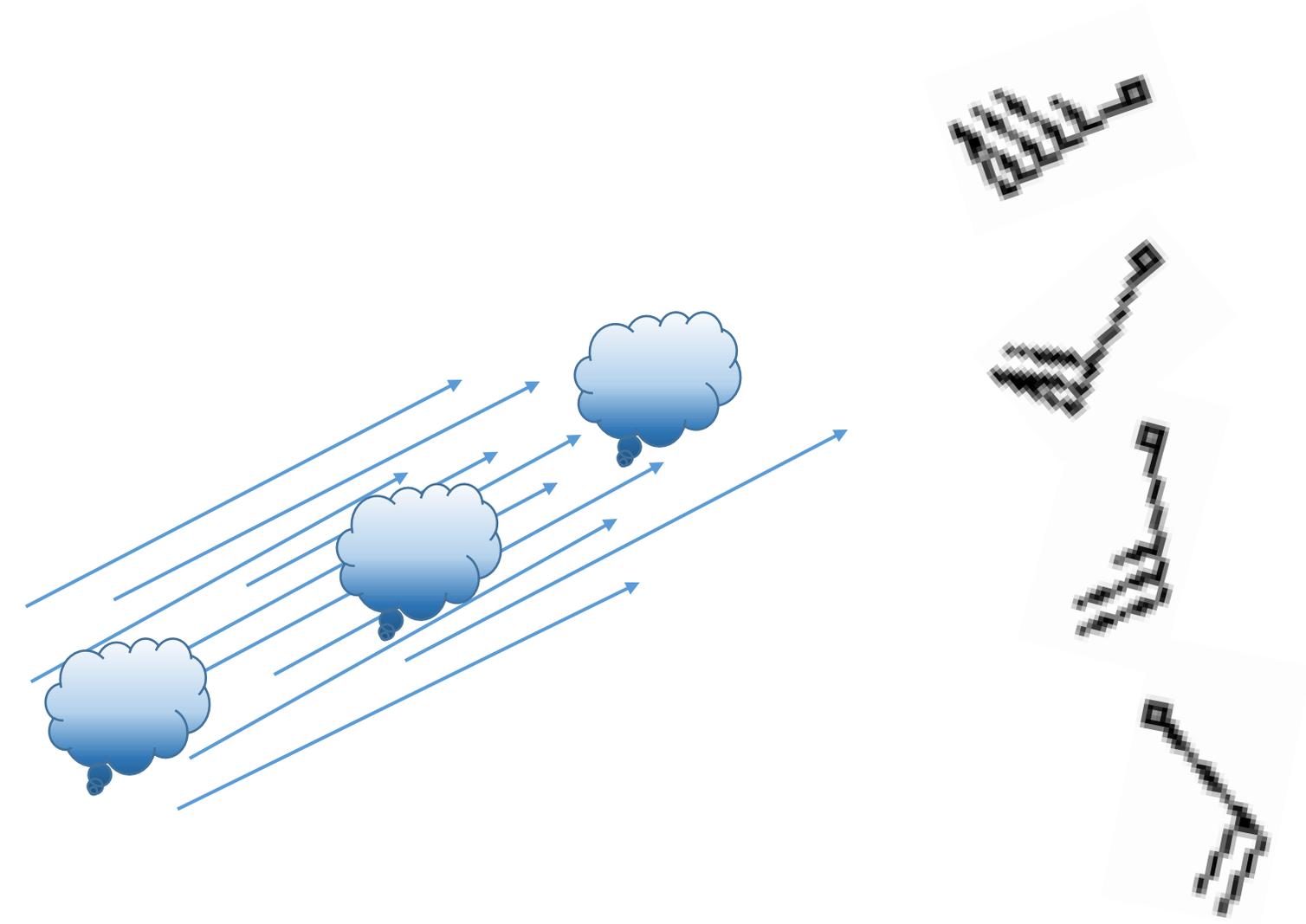
MCS Forward Motion

- **Advection**: Component of motion involving cells being carried with or by the mean wind in the cloud layer (V_{CL}).



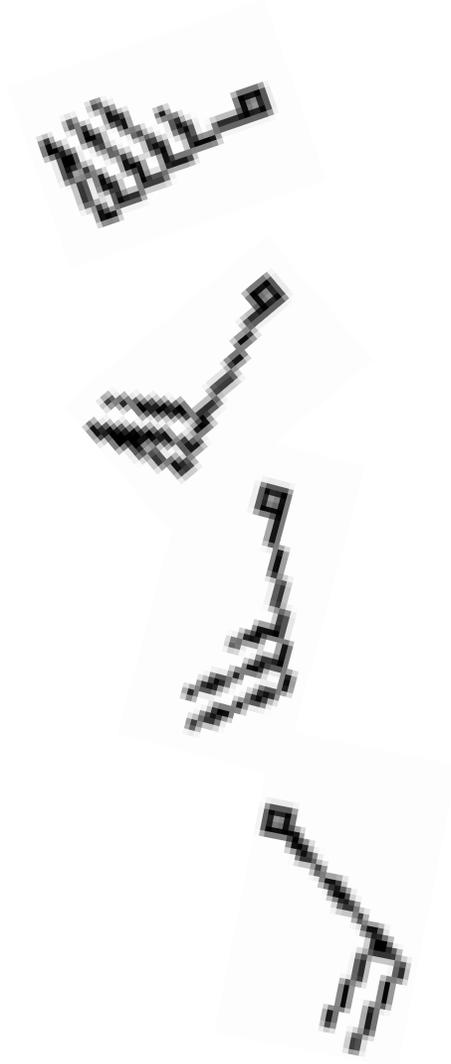
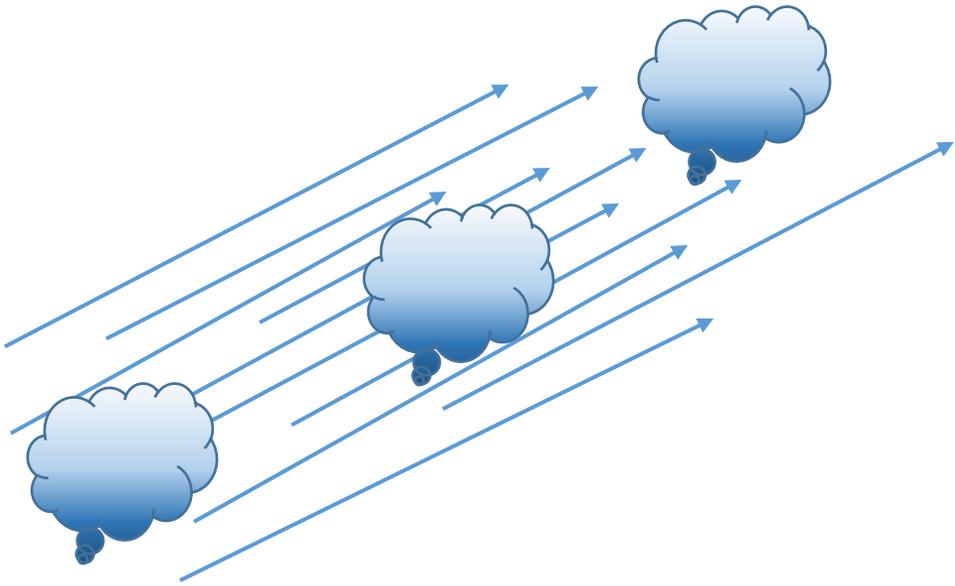
MCS Forward Motion

- **Advection**: Component of motion involving cells being carried with or by the mean wind in the cloud layer (V_{CL}).



MCS Forward Motion

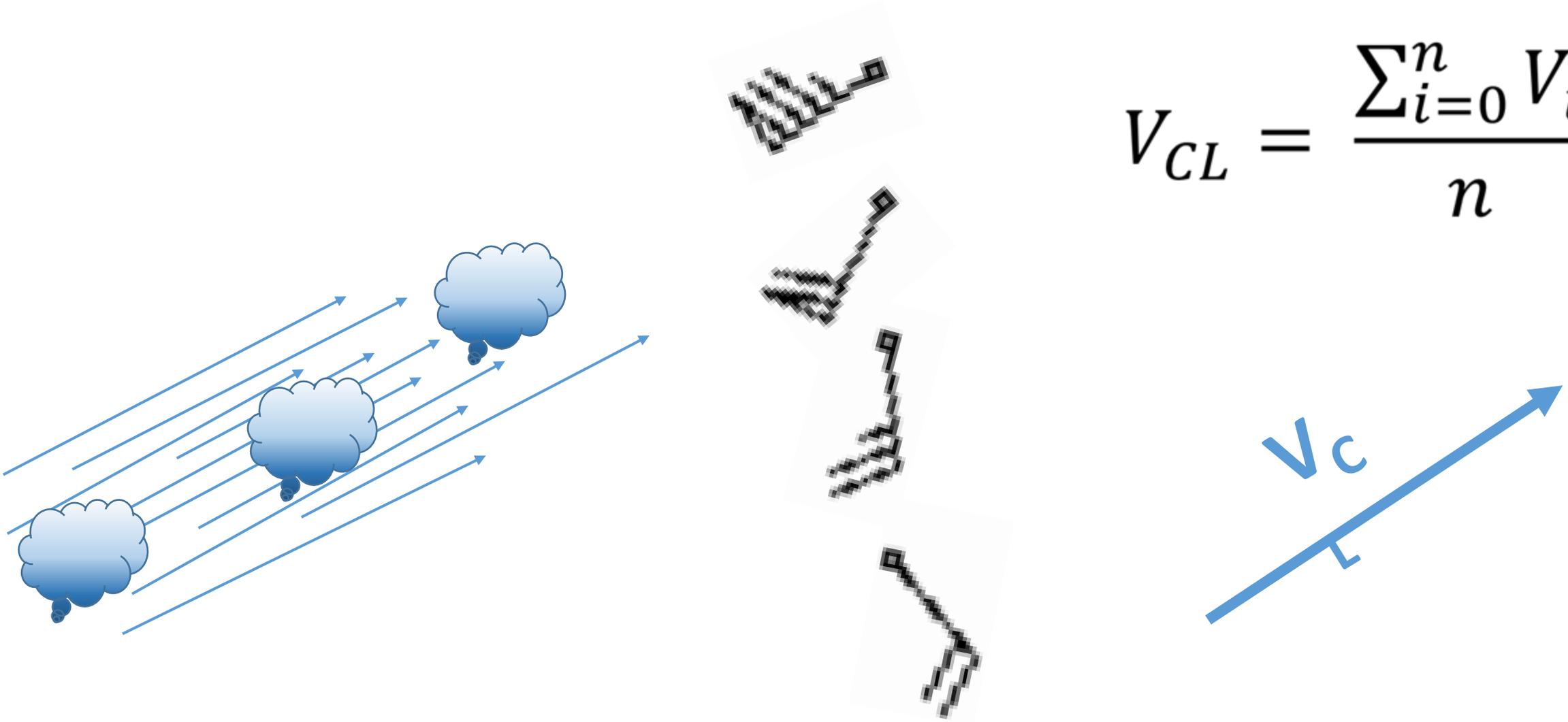
- **Advection**: Component of motion involving cells being carried with or by the mean wind in the cloud layer (V_{CL}).



$$V_{CL} = \frac{\sum_{i=0}^n V_i}{n}$$

MCS Forward Motion

- **Advection**: Component of motion involving cells being carried with or by the mean wind in the cloud layer (V_{CL}).



MCS Forward Motion

- MCS motion has 2 components (Advection and Propagation)

MCS Forward Motion

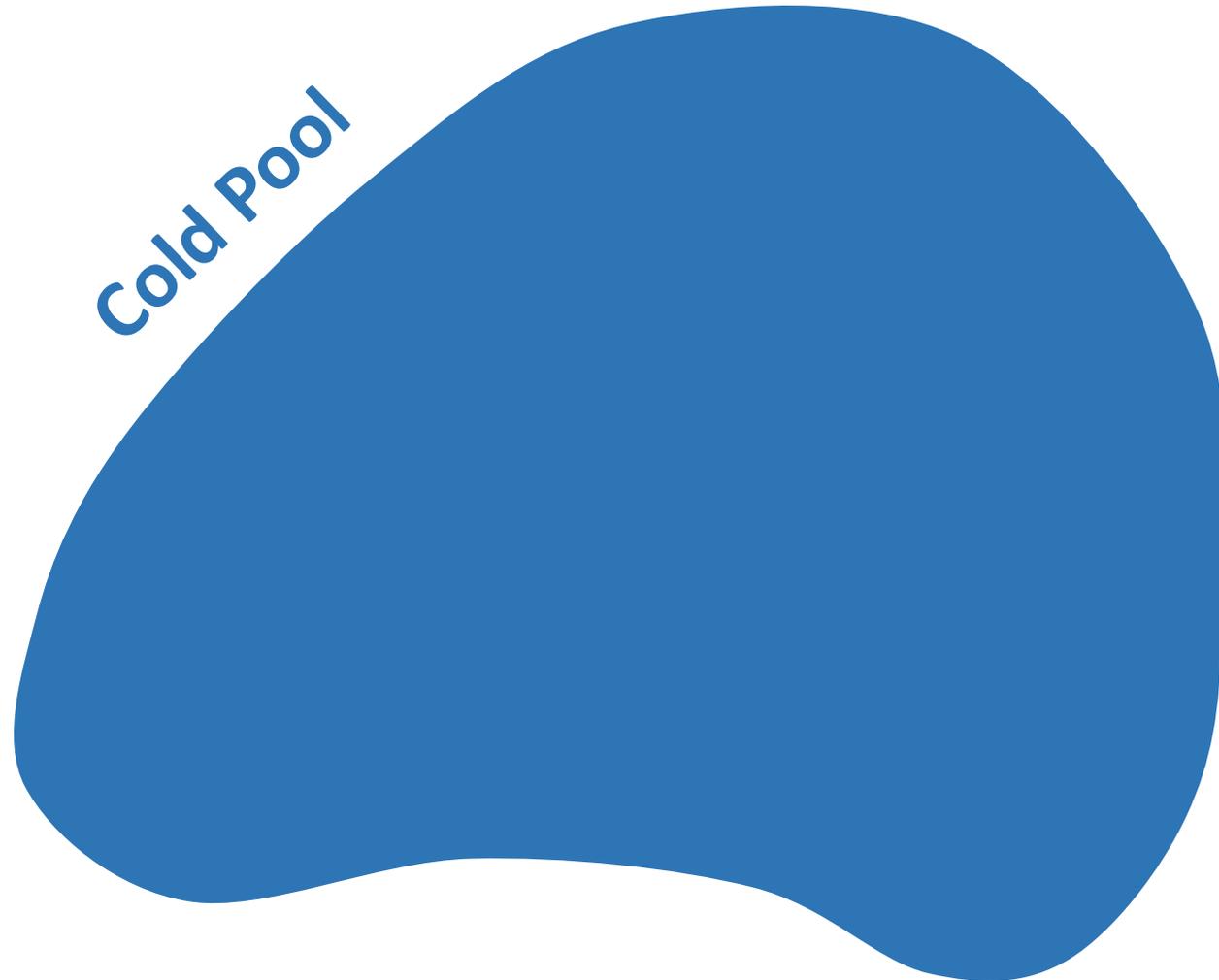
- Propagation:

MCS Forward Motion

- **Propagation**: Component of motion due to the initiation of new cells driven by the convergence of storm relative inflow (V_{Prop}).

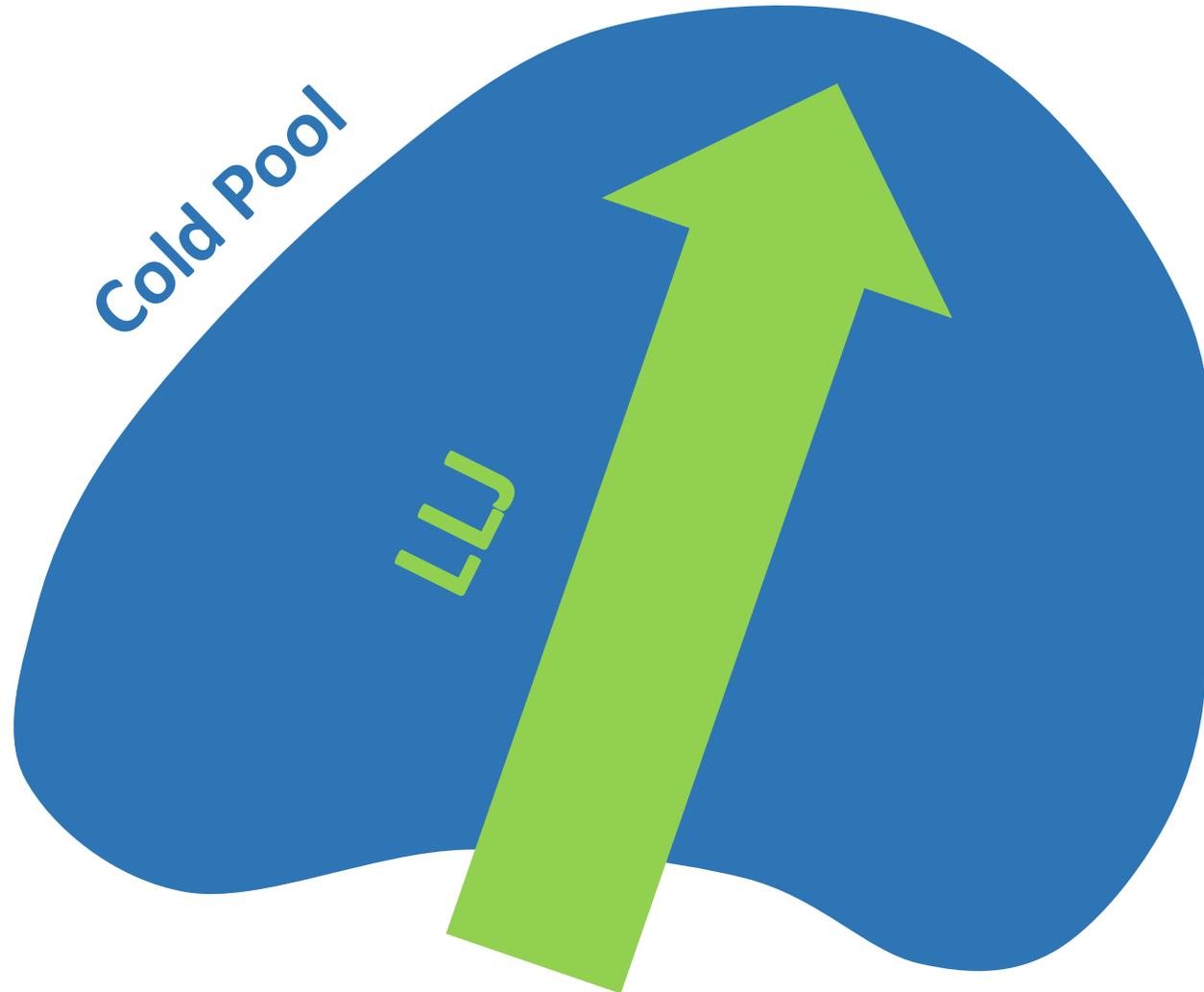
MCS Forward Motion

- **Propagation**: Component of motion due to the initiation of new cells driven by the convergence of storm relative inflow (V_{Prop}).



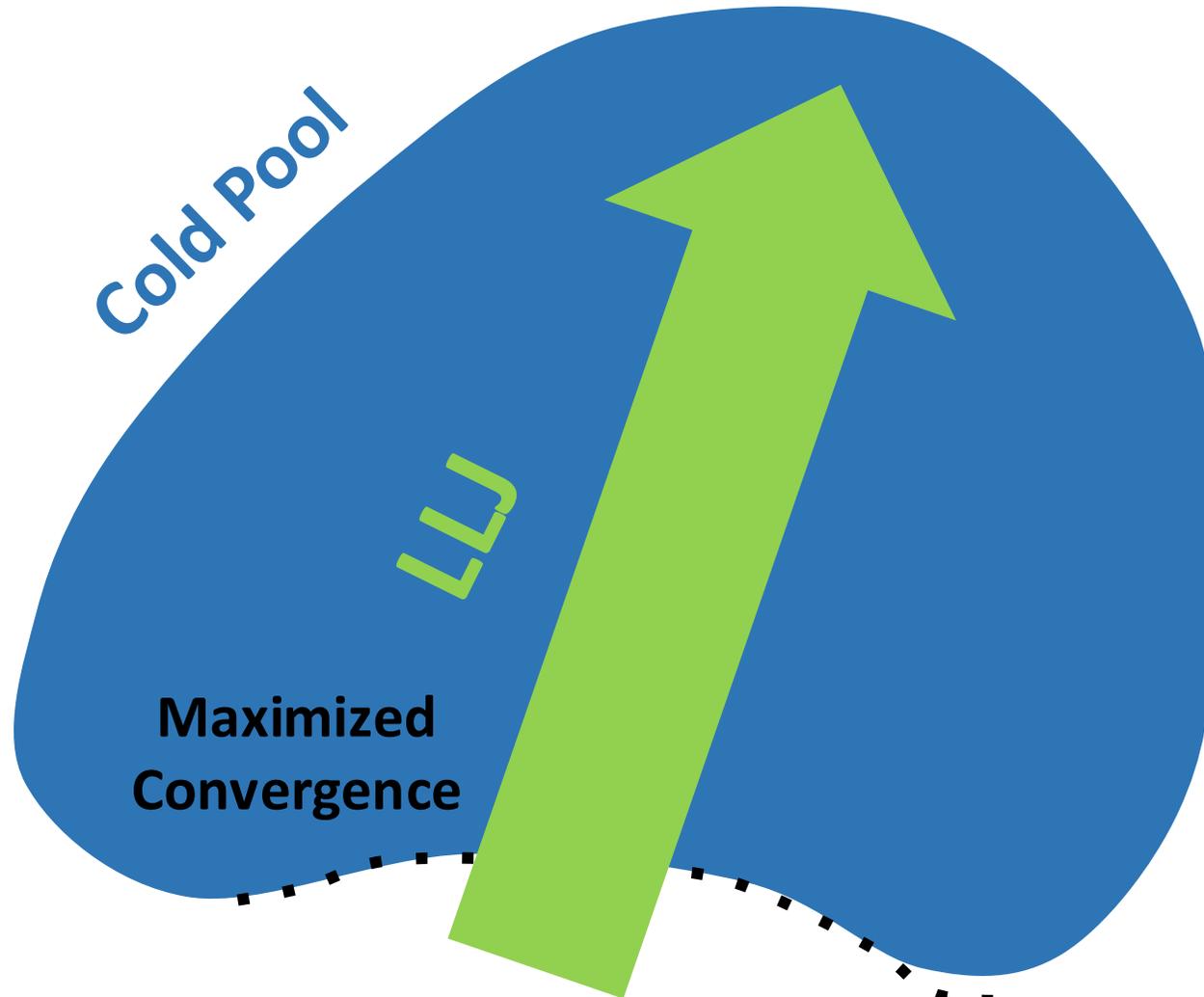
MCS Forward Motion

- **Propagation**: Component of motion due to the initiation of new cells driven by the convergence of storm relative inflow (V_{Prop}).



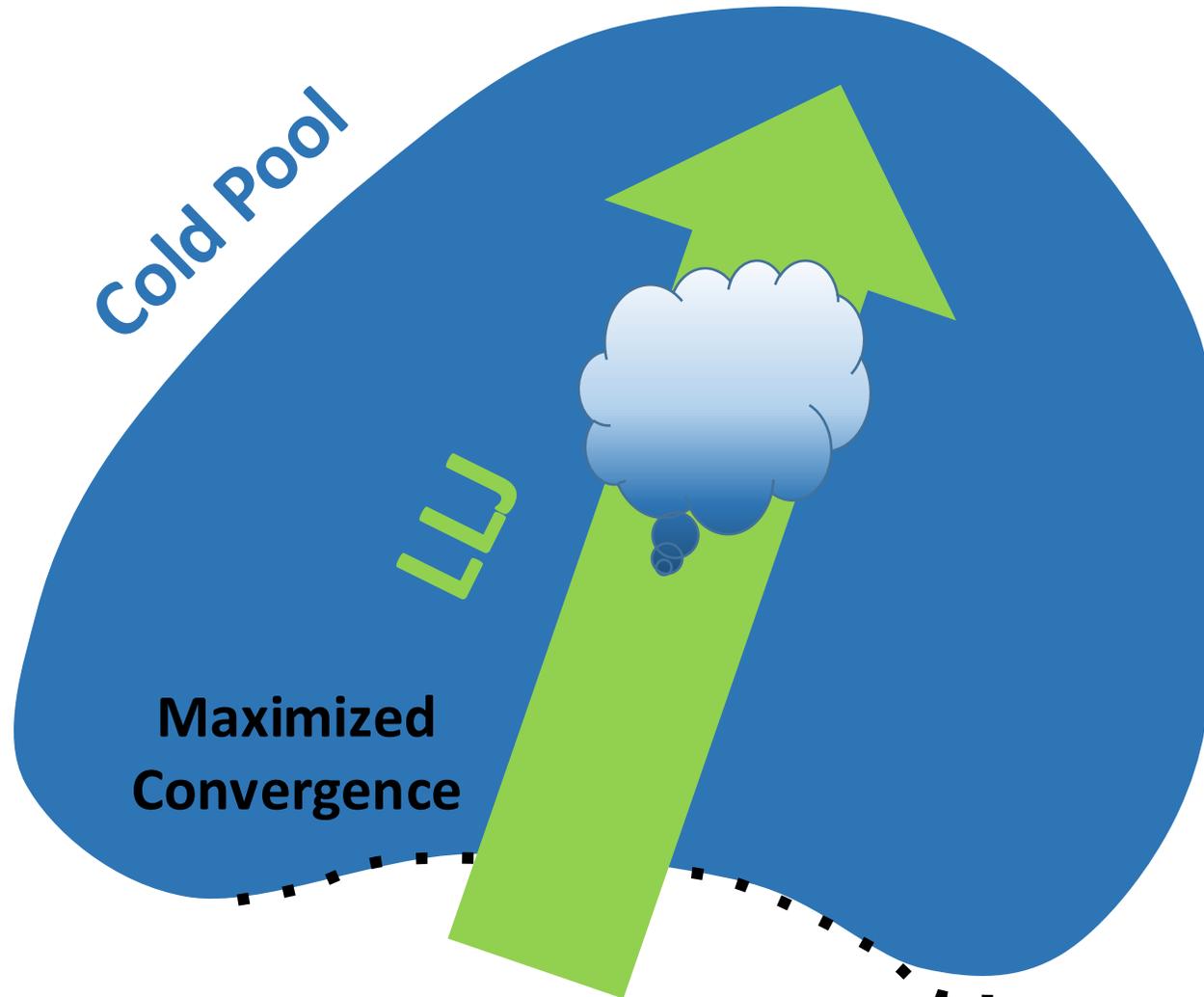
MCS Forward Motion

- **Propagation**: Component of motion due to the initiation of new cells driven by the convergence of storm relative inflow (V_{Prop}).



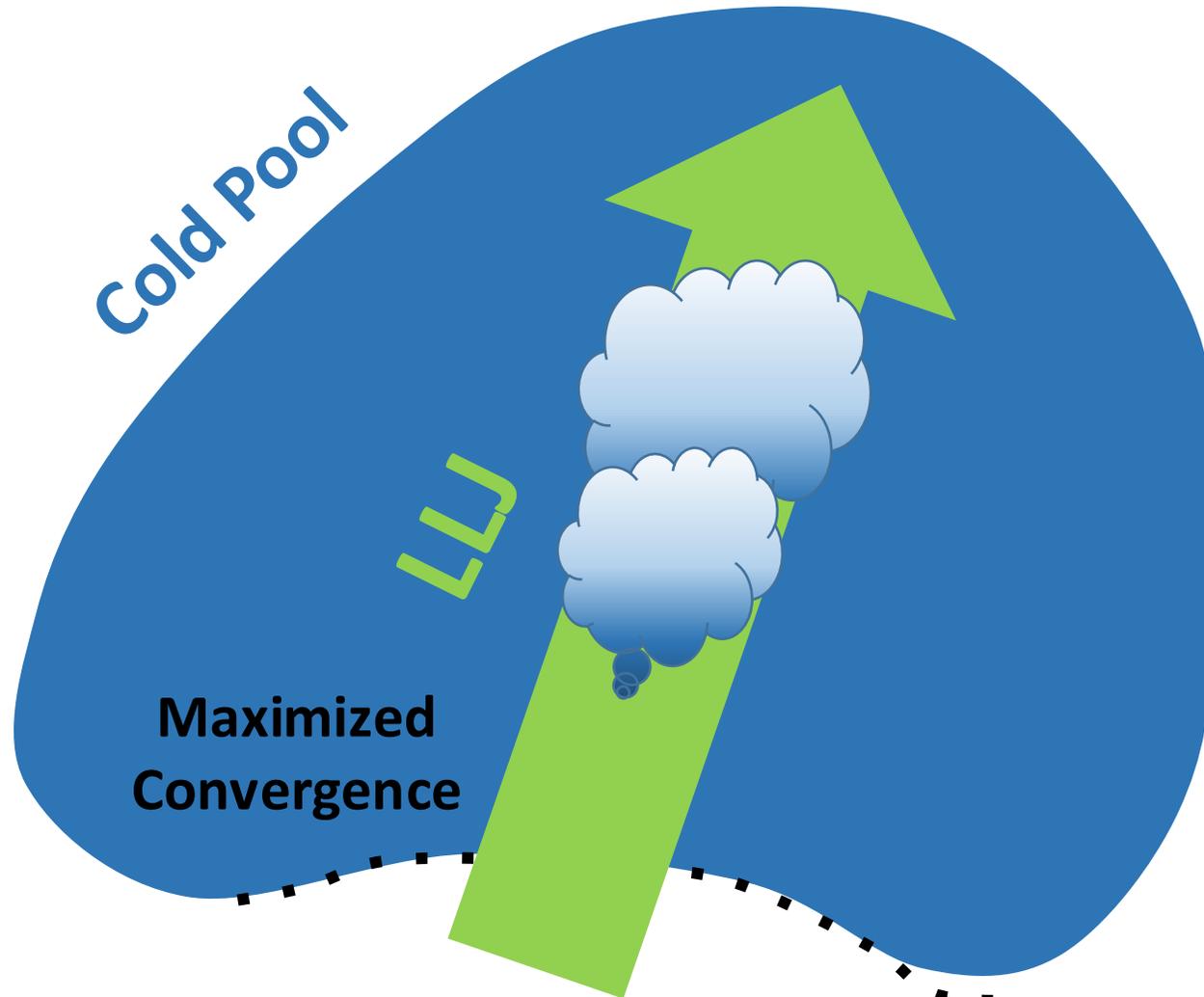
MCS Forward Motion

- **Propagation**: Component of motion due to the initiation of new cells driven by the convergence of storm relative inflow (V_{Prop}).



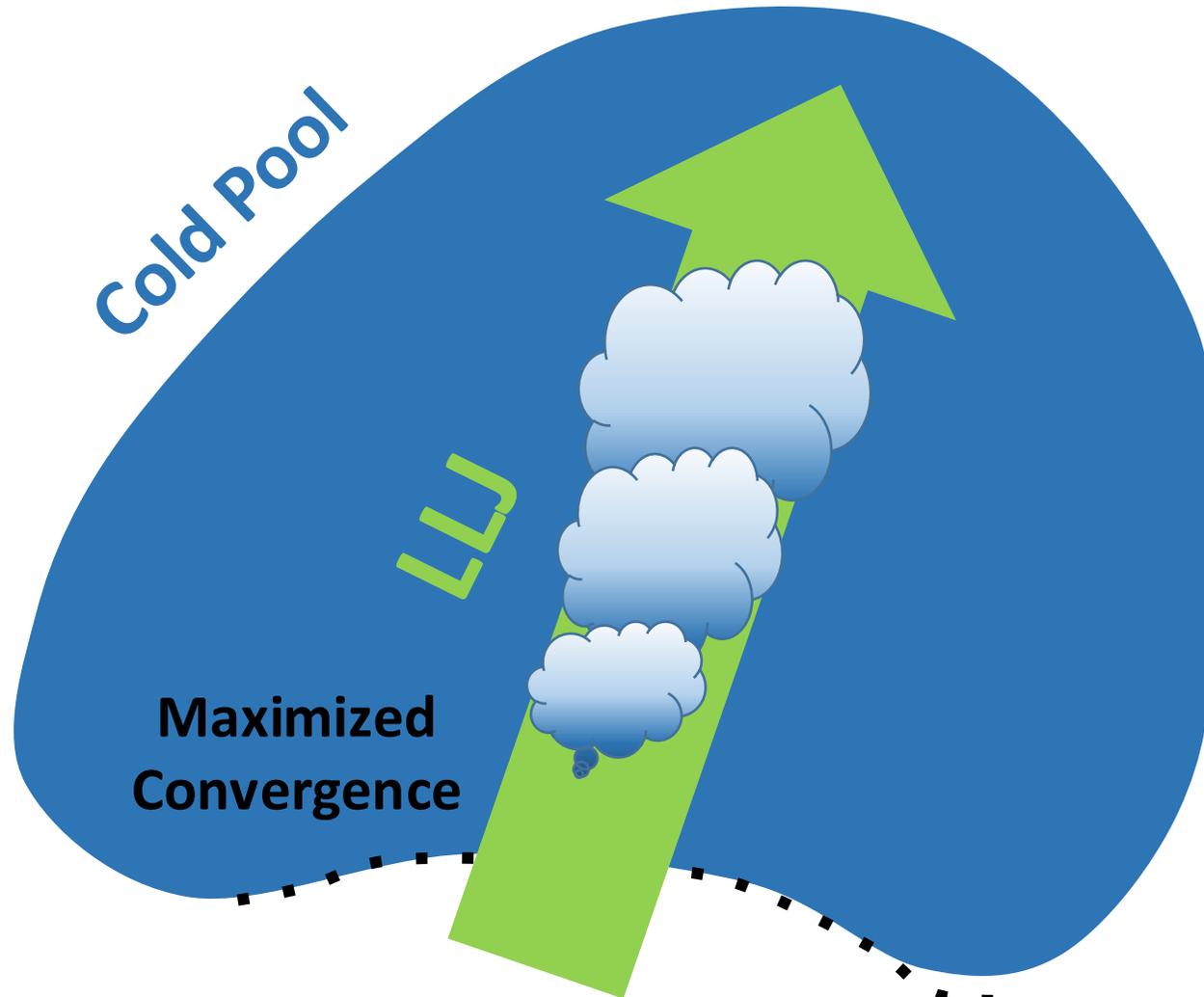
MCS Forward Motion

- **Propagation**: Component of motion due to the initiation of new cells driven by the convergence of storm relative inflow (V_{Prop}).



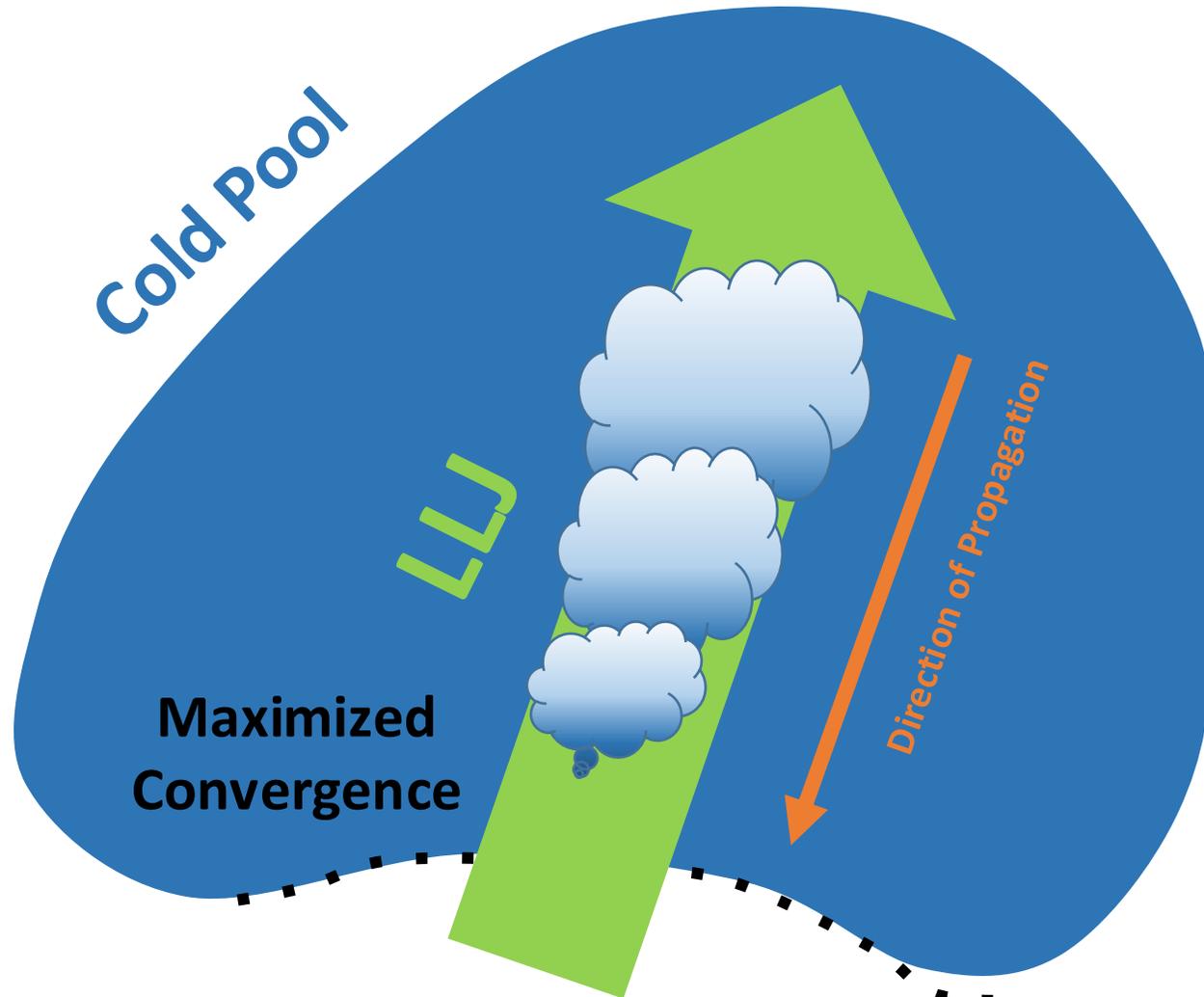
MCS Forward Motion

- **Propagation**: Component of motion due to the initiation of new cells driven by the convergence of storm relative inflow (V_{Prop}).



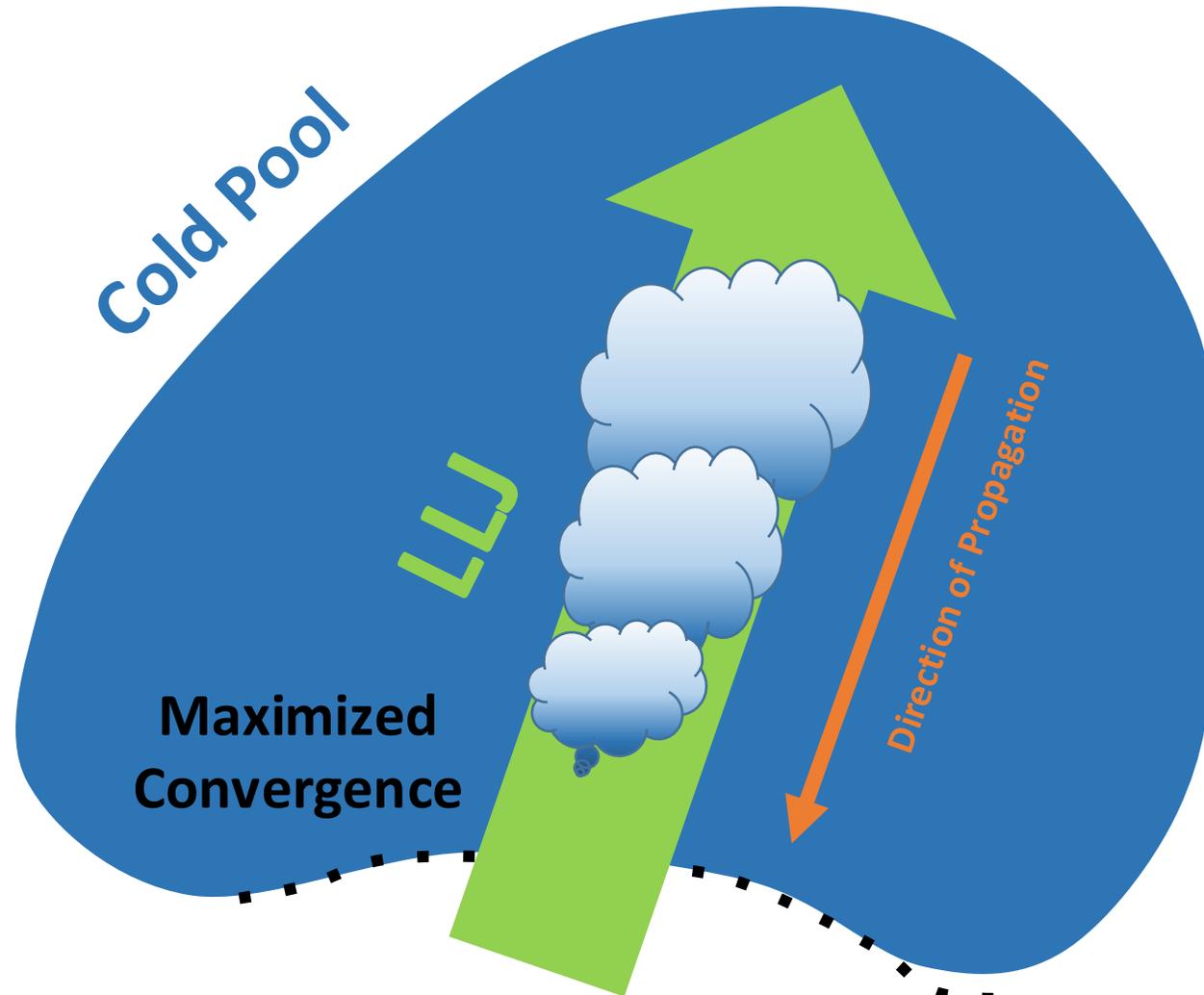
MCS Forward Motion

- **Propagation**: Component of motion due to the initiation of new cells driven by the convergence of storm relative inflow (V_{Prop}).



MCS Forward Motion

- **Propagation**: Component of motion due to the initiation of new cells driven by the convergence of storm relative inflow (V_{Prop}).



$$V_{-LLJ} = V_{Prop}$$

MCS Forward Motion

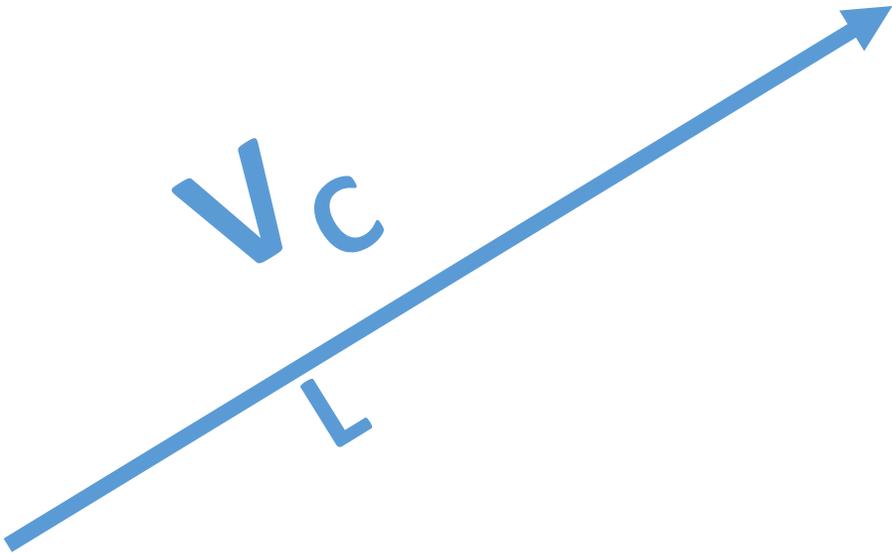
- MCS motion has 2 components (Advection and Propagation)

MCS Forward Motion

- MCS motion has 2 components (Advection and Propagation)
- The Advection and Propagation components are both additive.

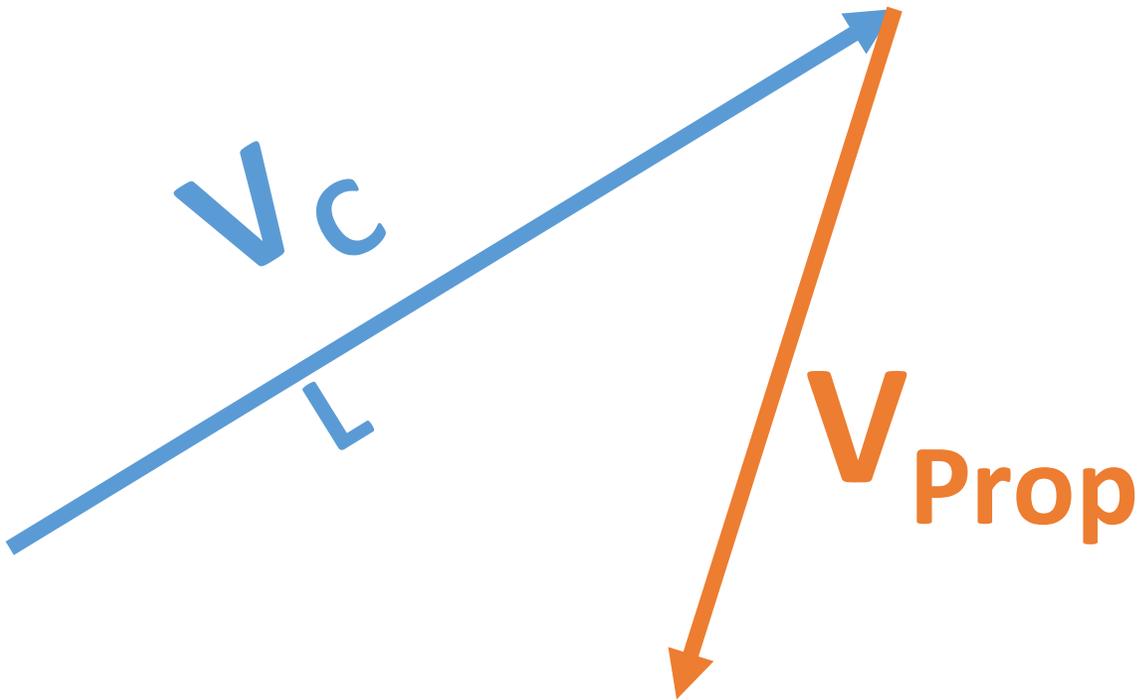
MCS Forward Motion

- MCS motion has 2 components (Advection and Propagation)
- The Advection and Propagation components are both additive.



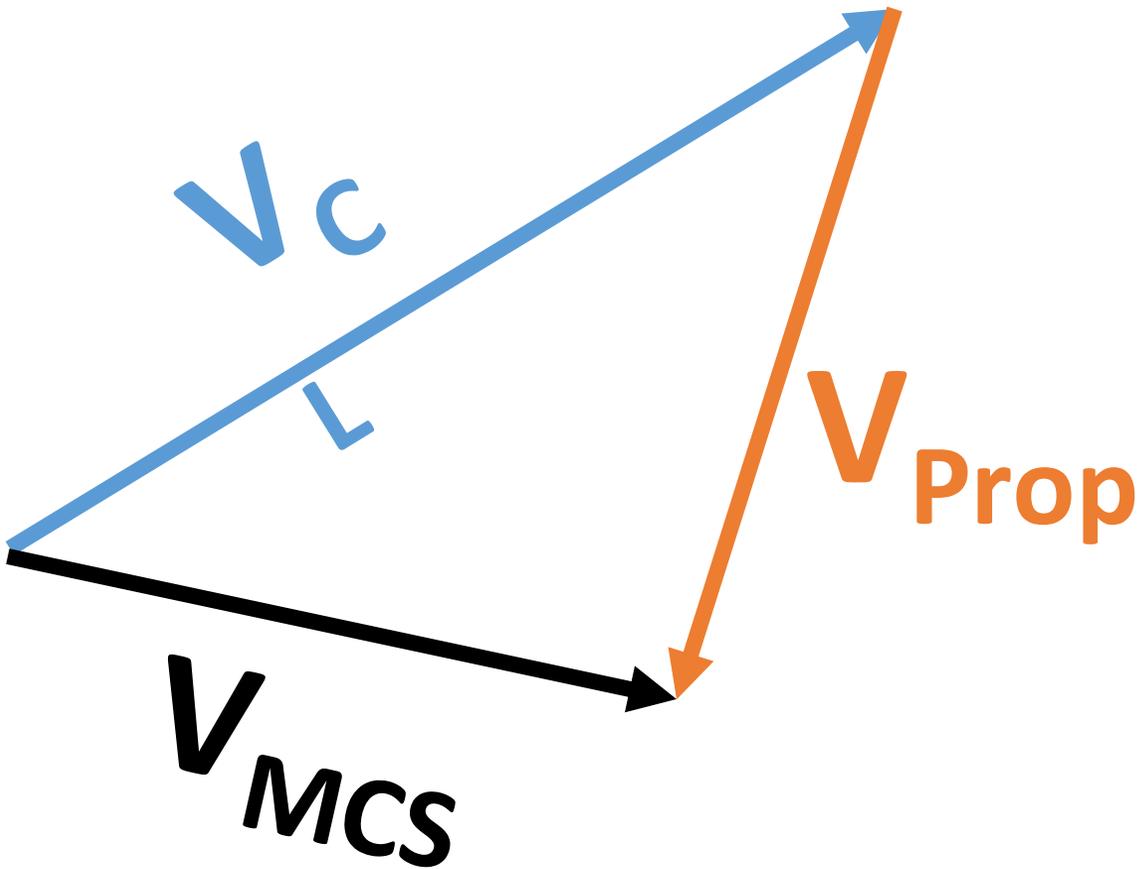
MCS Forward Motion

- MCS motion has 2 components (Advection and Propagation)
- The Advection and Propagation components are both additive.



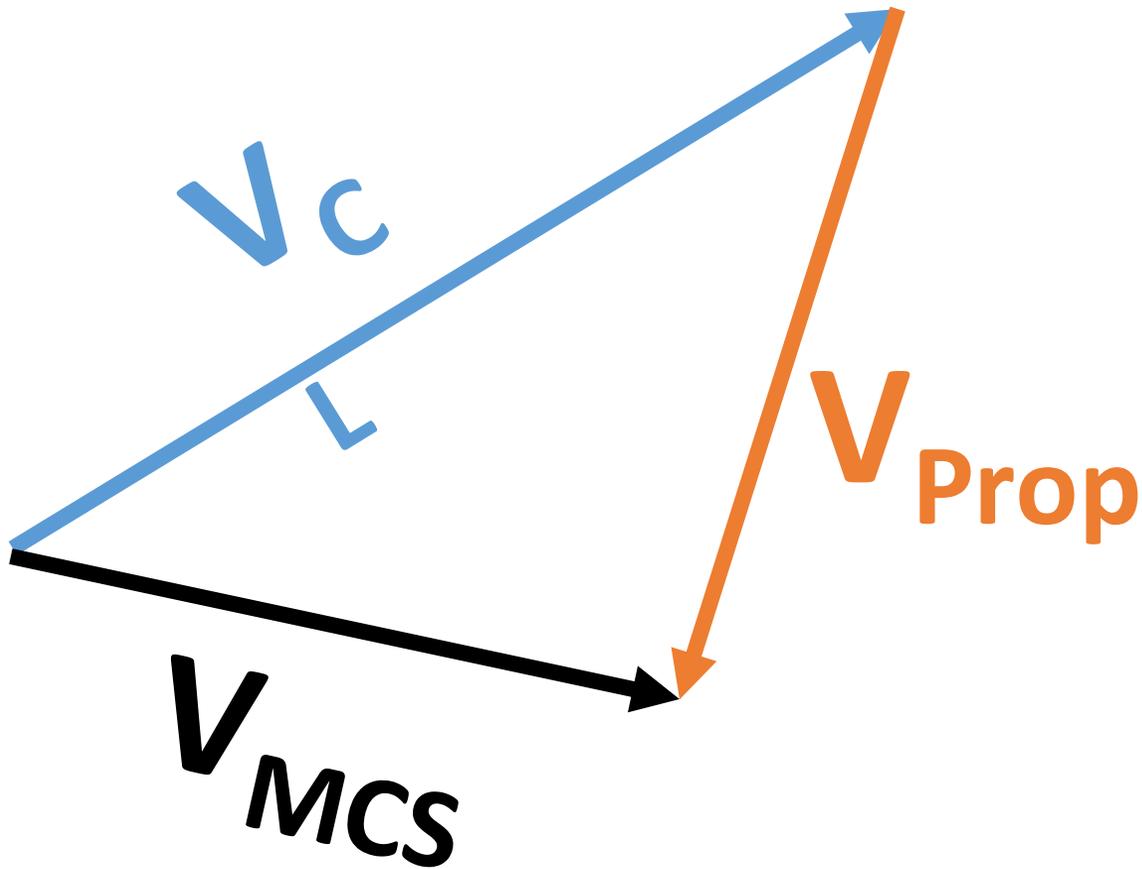
MCS Forward Motion

- MCS motion has 2 components (Advection and Propagation)
- The Advection and Propagation components are both additive.



MCS Forward Motion

- MCS motion has 2 components (Advection and Propagation)
- The Advection and Propagation components are both additive.



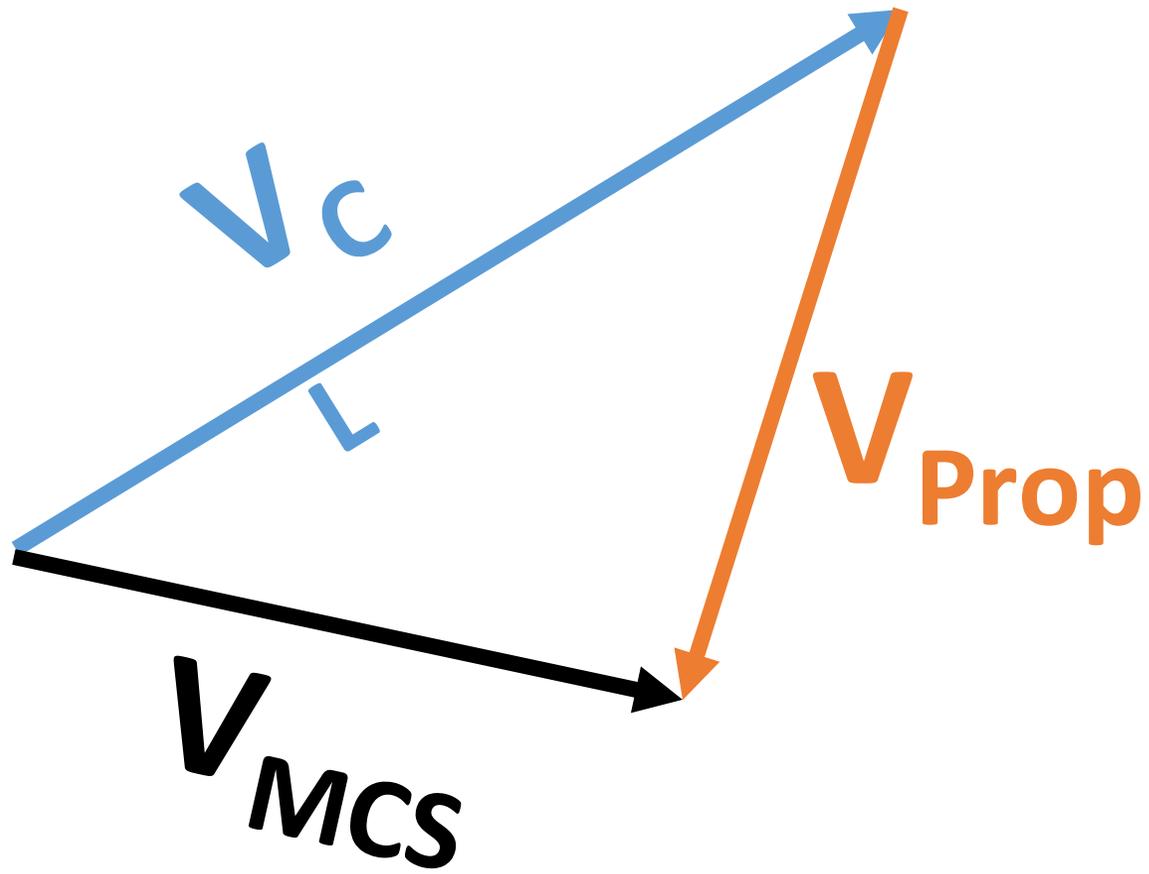
$$V_{MCS} = V_{CL} + V_{Prop}$$

OR

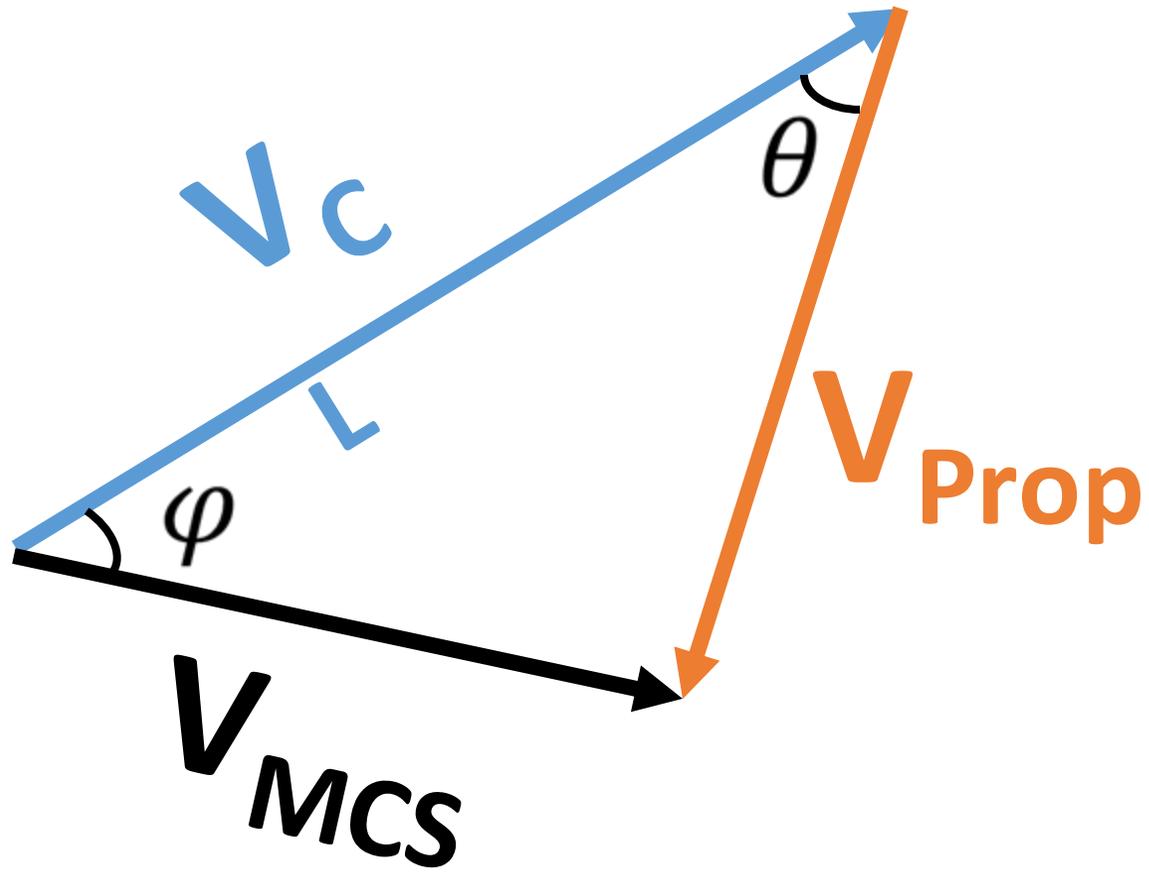
$$V_{MCS} = V_{CL} - V_{LLJ}$$

V_{MCS} Magnitude and Direction

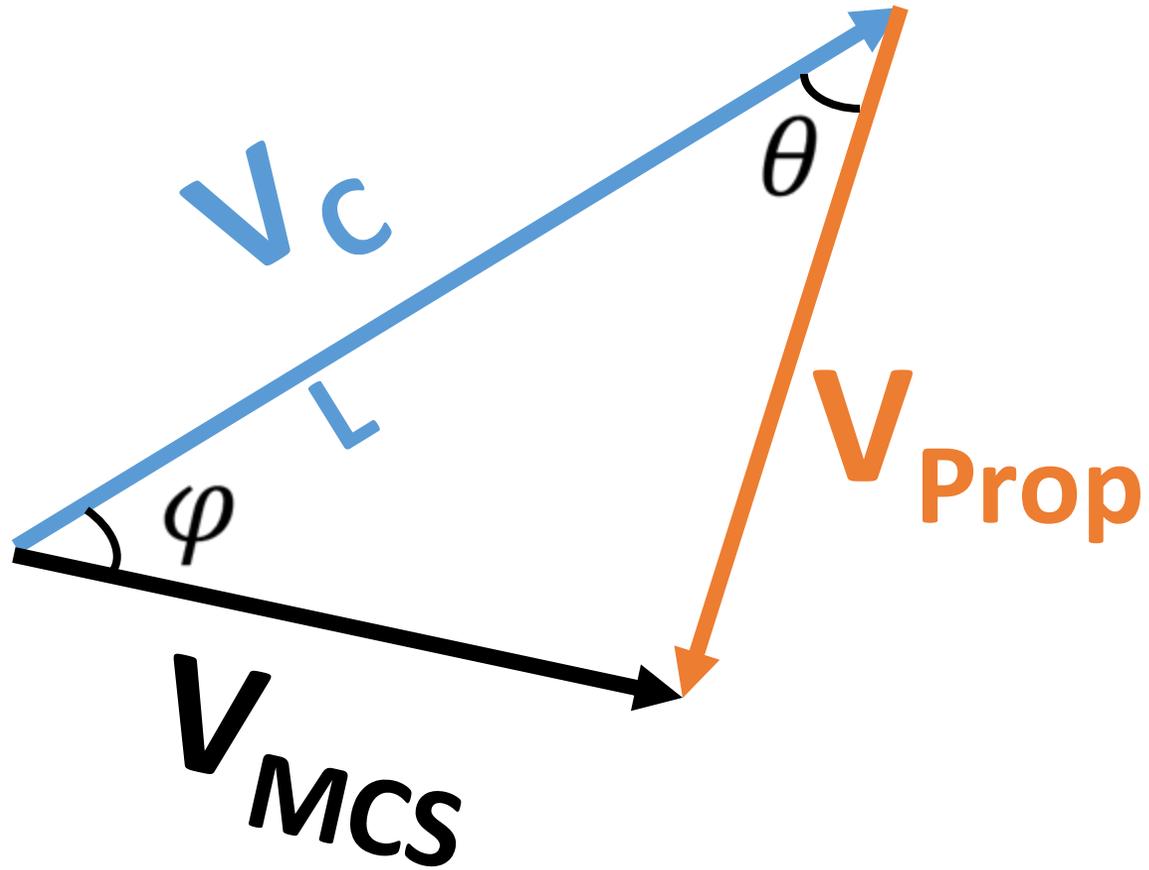
V_{MCS} Magnitude and Direction



V_{MCS} Magnitude and Direction

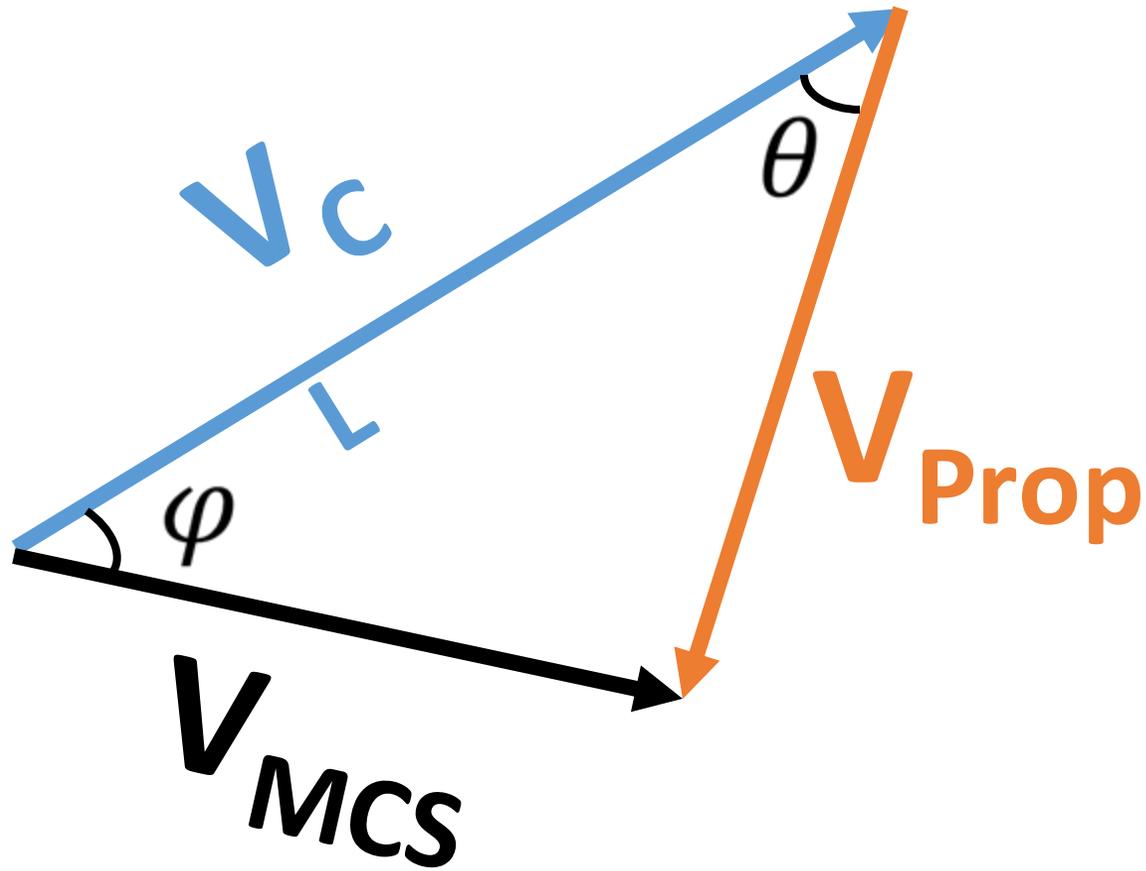


V_{MCS} Magnitude and Direction



We can find the speed and direction of the MCS motion vector using trigonometry.

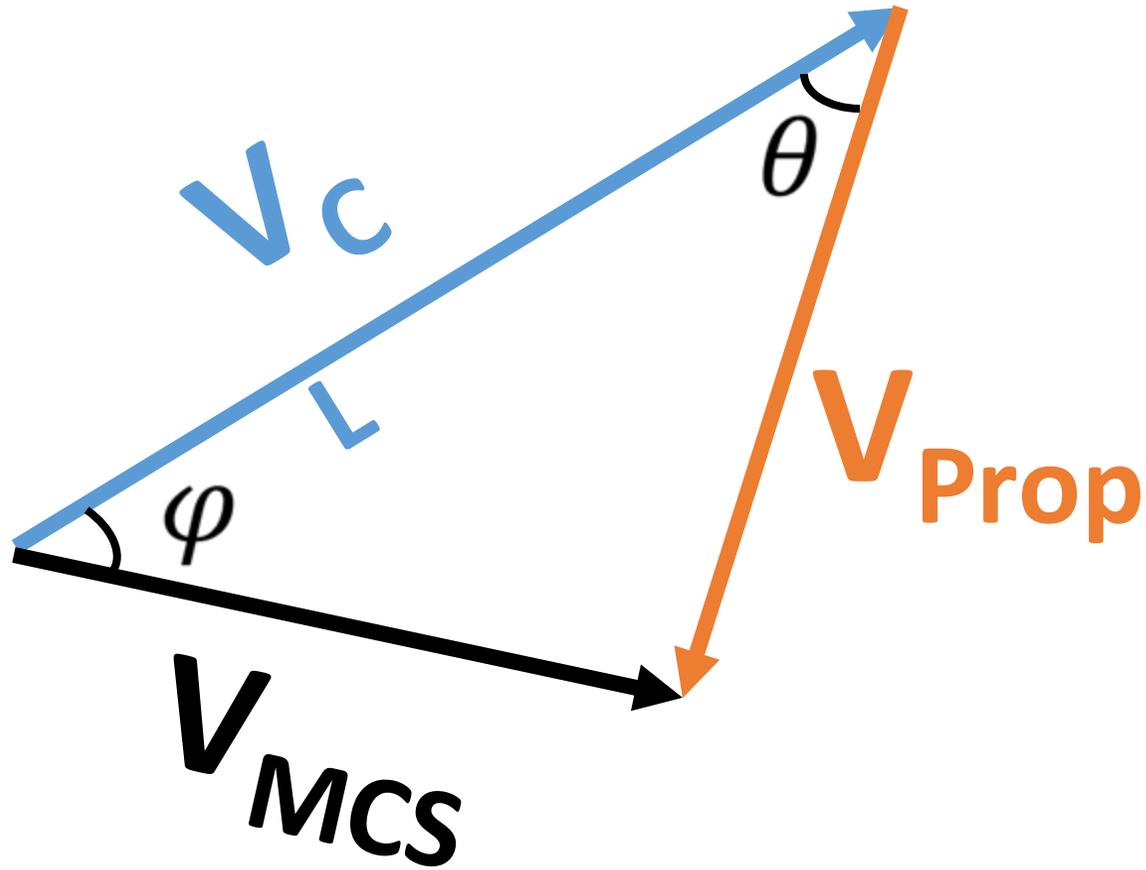
V_{MCS} Magnitude and Direction



We can find the speed and direction of the MCS motion vector using trigonometry.

However, these calculations are sensitive to the depth of the cloud layer of the mean wind, and the inflow layer available to the MCS, which can influence propagation.

V_{MCS} Magnitude and Direction

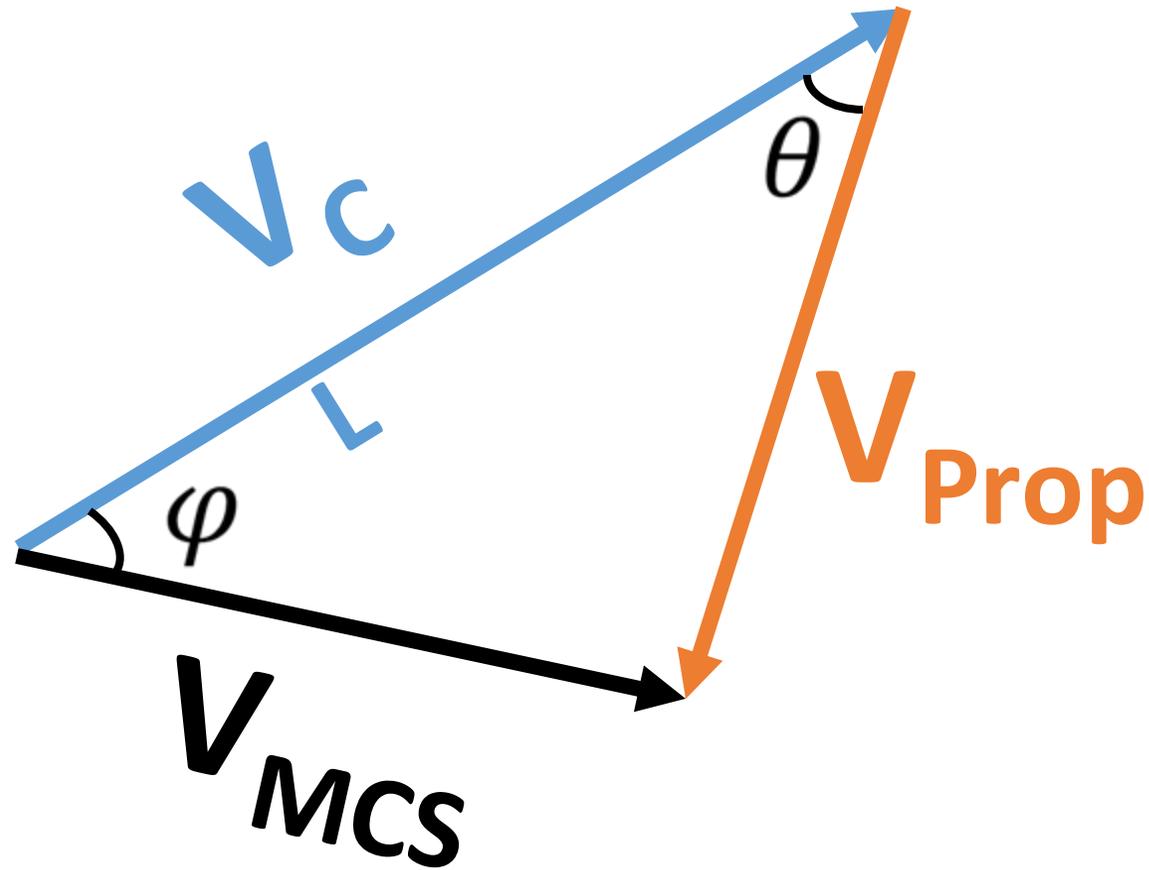


We can find the speed and direction of the MCS motion vector using trigonometry.

However, these calculations are sensitive to the depth of the cloud layer of the mean wind, and the inflow layer available to the MCS, which can influence propagation.

$$\varphi = \arcsin\left(\frac{|V_{Prop}| * \sin(\theta)}{V_{MCS}}\right)$$

V_{MCS} Magnitude and Direction



We can find the speed and direction of the MCS motion vector using trigonometry.

However, these calculations are sensitive to the depth of the cloud layer of the mean wind, and the inflow layer available to the MCS, which can influence propagation.

$$\varphi = \arcsin\left(\frac{|V_{Prop}| * \sin(\theta)}{V_{MCS}}\right)$$

$$|V_{MCS}| = \sqrt{|V_{CL}|^2 + |V_{Prop}|^2 - 2(|V_{CL}| * |V_{Prop}|)\cos\theta}$$

What happens when we increase LLJ wind speeds?

What happens when we increase LLJ wind speeds?

$$V_{CL} = 50 \text{ kts}$$

$$\theta = 50 \text{ degrees (0.87 radians)}$$

$$V_{MCS}?$$

$$\varphi?$$

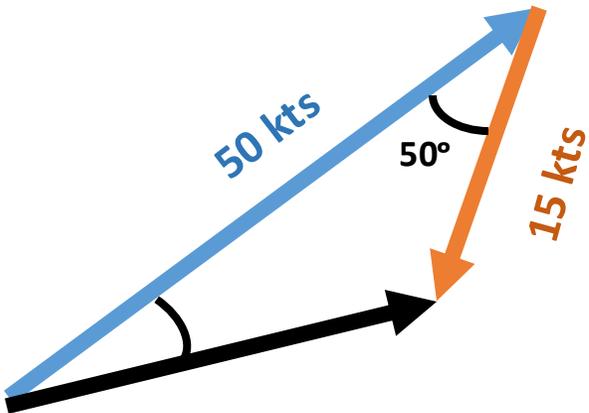
What happens when we increase LLJ wind speeds?

$$V_{CL} = 50 \text{ kts}$$

$$\theta = 50 \text{ degrees (0.87 radians)}$$

$$V_{MCS}?$$

$$\varphi?$$



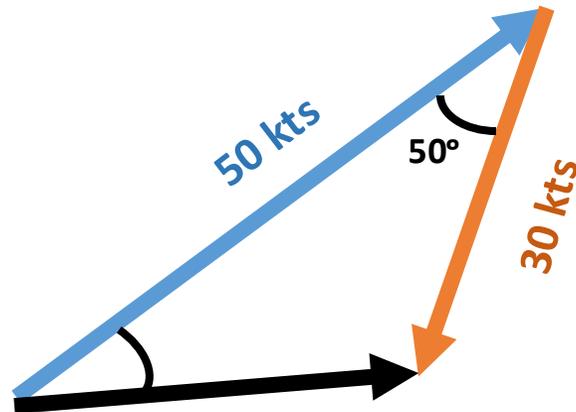
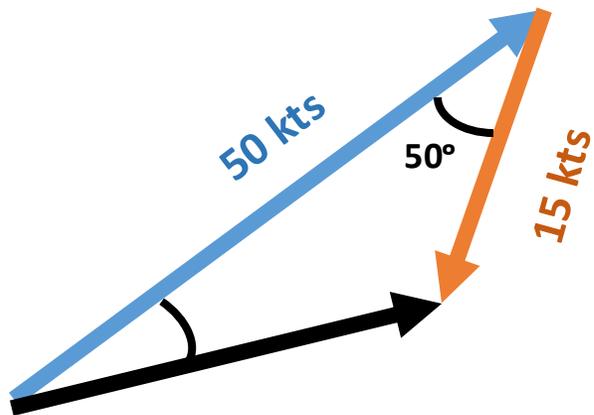
What happens when we increase LLJ wind speeds?

$V_{CL} = 50$ kts

$\theta = 50$ degrees (0.87 radians)

$V_{MCS}?$

$\varphi?$



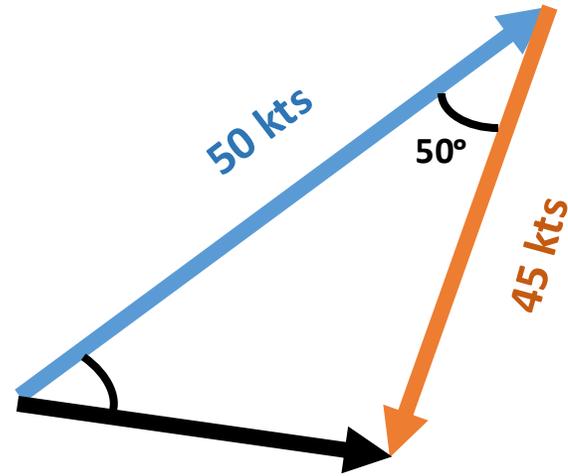
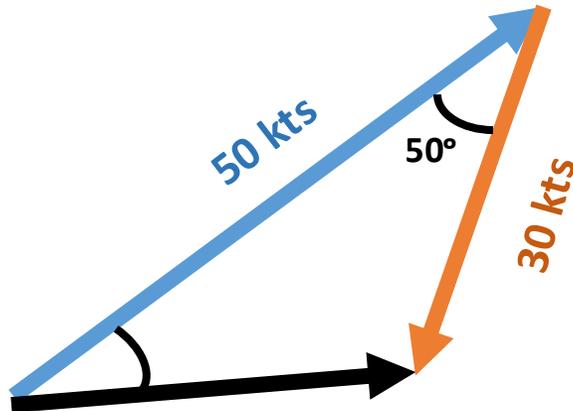
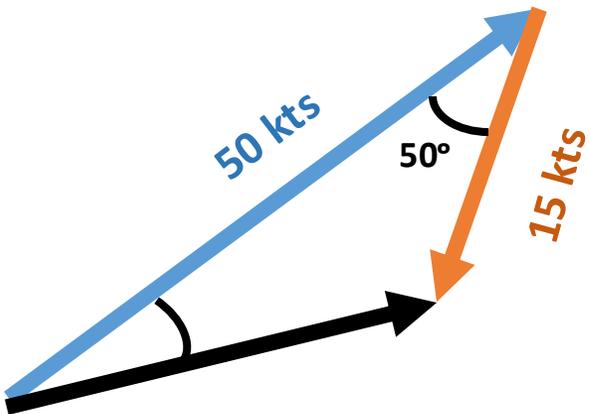
What happens when we increase LLJ wind speeds?

$V_{CL} = 50$ kts

$\theta = 50$ degrees (0.87 radians)

$V_{MCS}?$

$\varphi?$



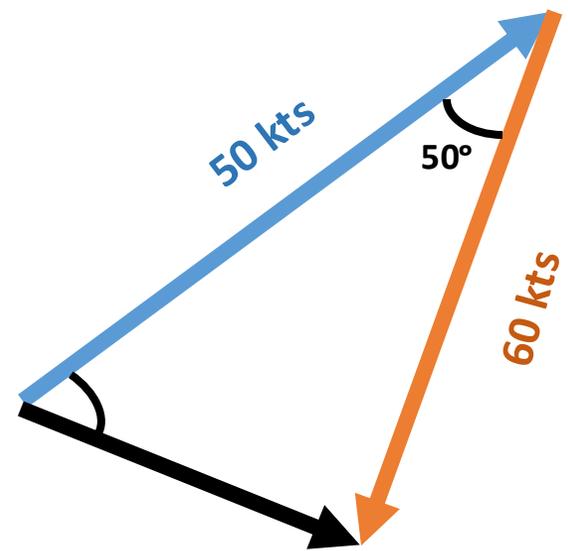
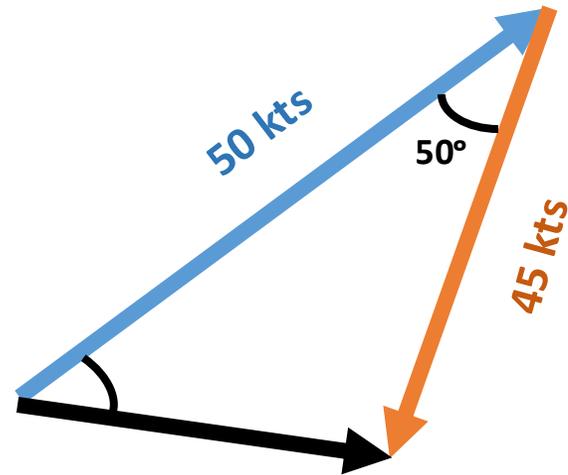
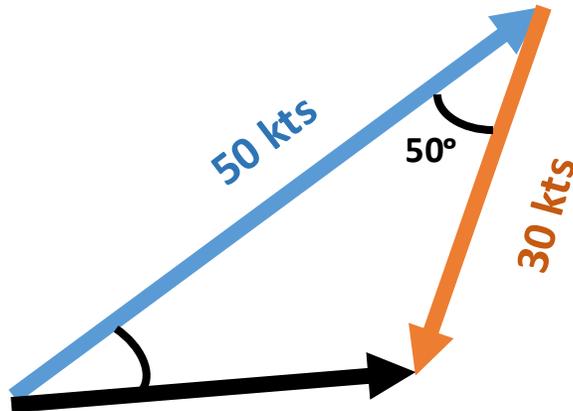
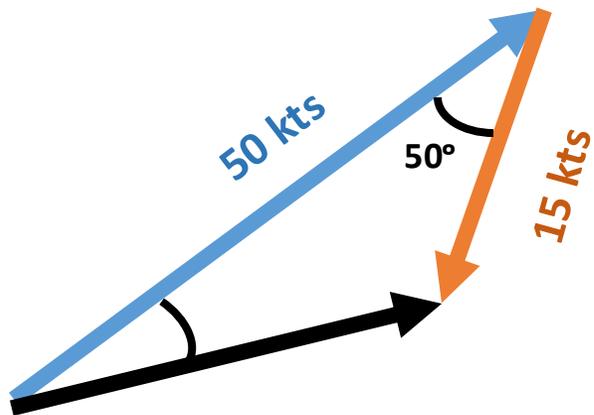
What happens when we increase LLJ wind speeds?

$V_{CL} = 50$ kts

$\theta = 50$ degrees (0.87 radians)

$V_{MCS}?$

$\varphi?$



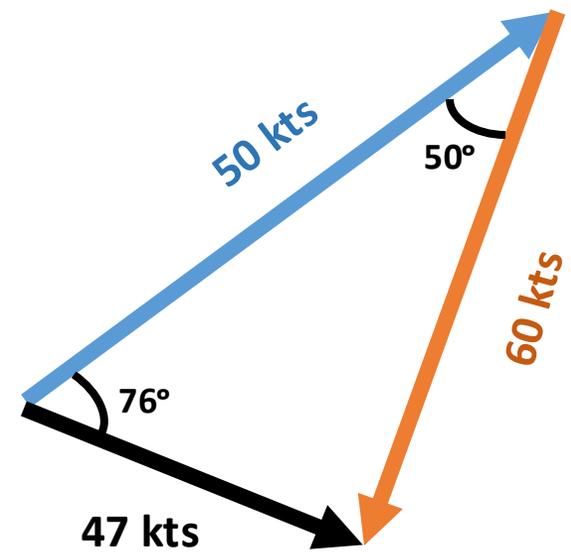
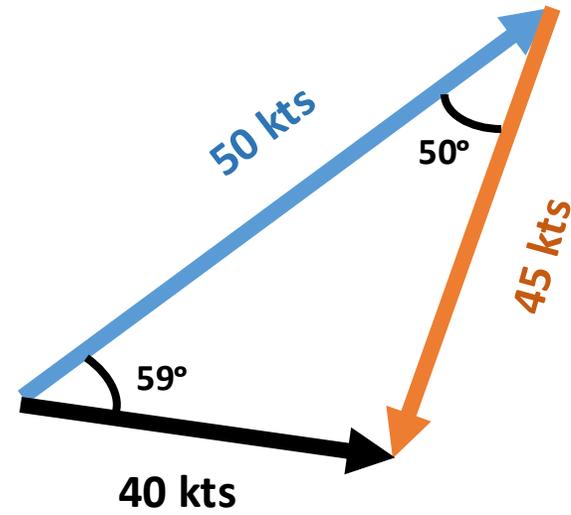
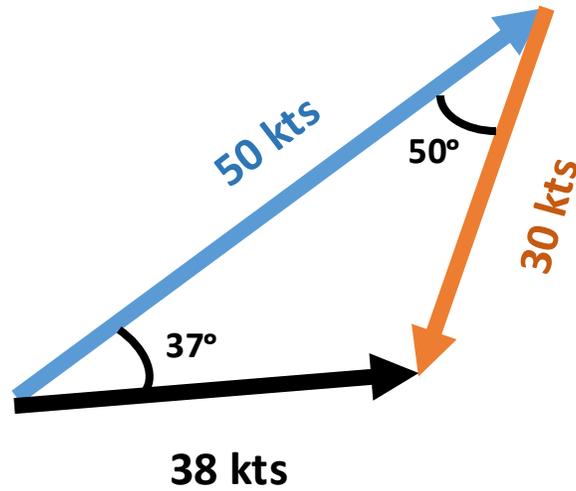
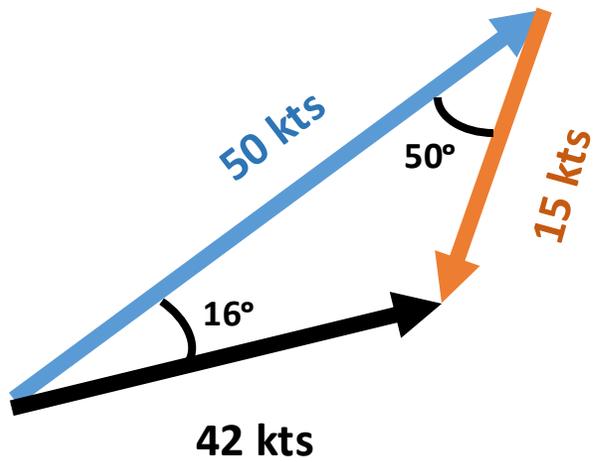
What happens when we increase LLJ wind speeds?

$V_{CL} = 50$ kts

$\theta = 50$ degrees (0.87 radians)

$V_{MCS}?$

$\varphi?$



What happens when we decrease mean wind speeds?

What happens when we decrease mean wind speeds?

$V_{\text{PROP/LLJ}} = 30 \text{ kts}$

$\theta = 50 \text{ degrees (0.87 radians)}$

$V_{\text{MCS}}?$

$\varphi?$

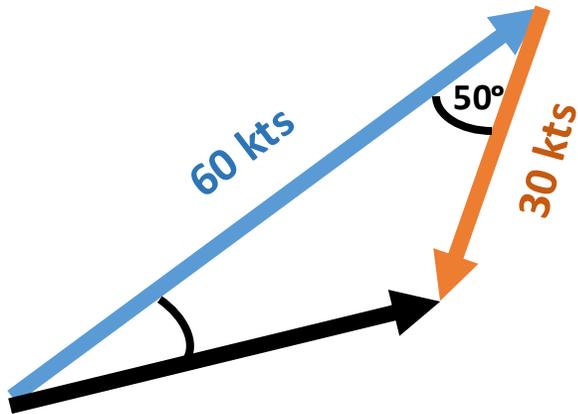
What happens when we decrease mean wind speeds?

$V_{\text{PROP/LLJ}} = 30 \text{ kts}$

$\theta = 50 \text{ degrees (0.87 radians)}$

$V_{\text{MCS}}?$

$\varphi?$



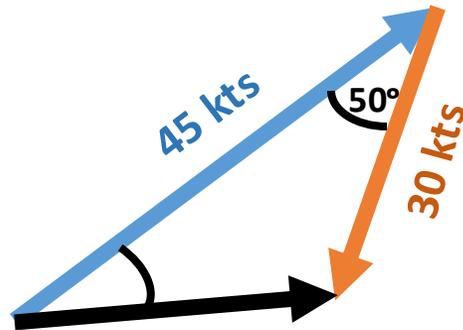
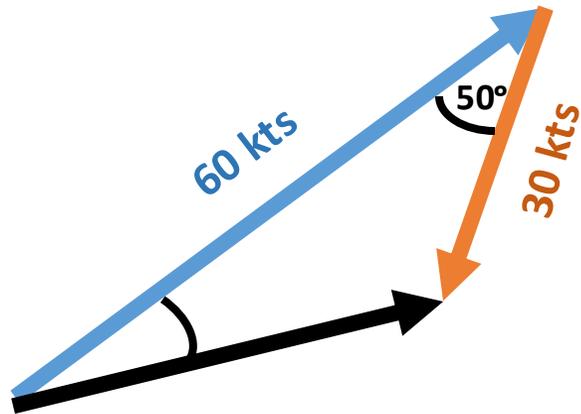
What happens when we decrease mean wind speeds?

$V_{\text{PROP/LLJ}} = 30 \text{ kts}$

$\theta = 50 \text{ degrees (0.87 radians)}$

$V_{\text{MCS}}?$

$\varphi?$



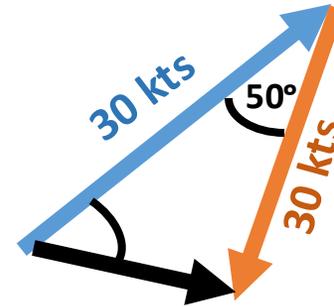
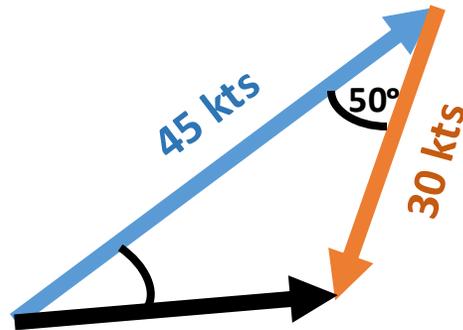
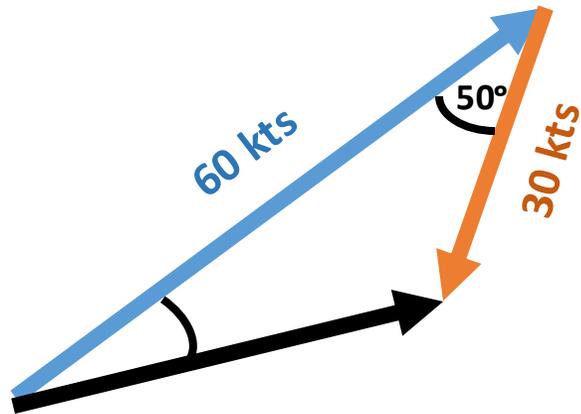
What happens when we decrease mean wind speeds?

$V_{\text{PROP/LLJ}} = 30 \text{ kts}$

$\theta = 50 \text{ degrees (0.87 radians)}$

$V_{\text{MCS}}?$

$\varphi?$



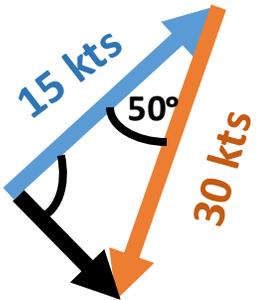
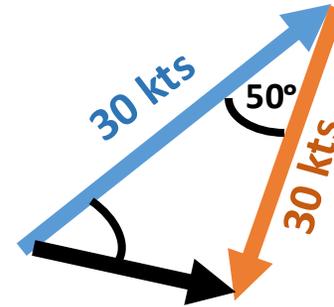
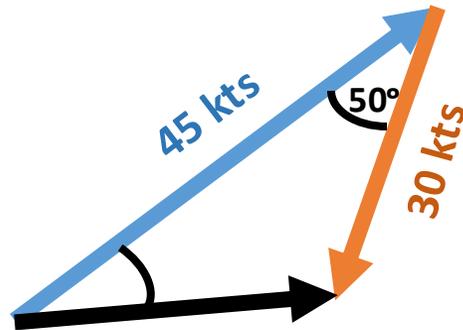
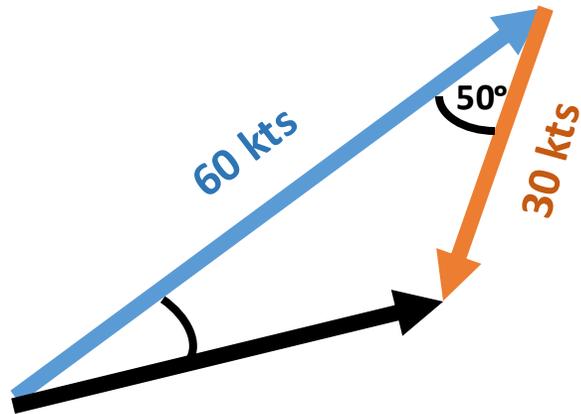
What happens when we decrease mean wind speeds?

$$V_{\text{PROP/LLJ}} = 30 \text{ kts}$$

$$\theta = 50 \text{ degrees (0.87 radians)}$$

$$V_{\text{MCS}}?$$

$$\varphi?$$



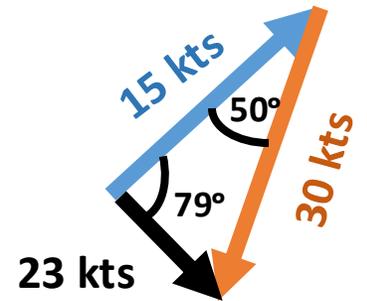
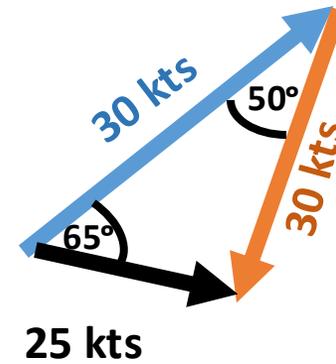
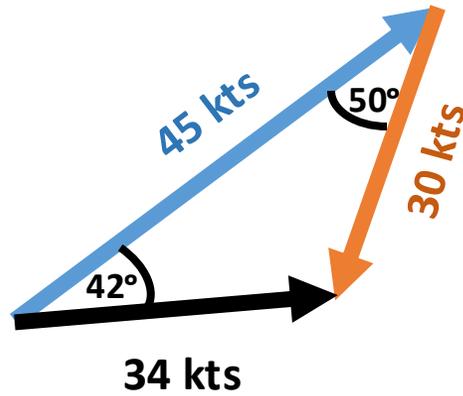
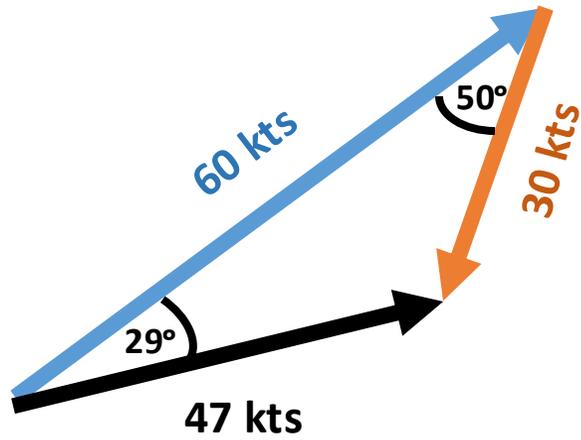
What happens when we decrease mean wind speeds?

$V_{\text{PROP/LLJ}} = 30 \text{ kts}$

$\theta = 50 \text{ degrees (0.87 radians)}$

$V_{\text{MCS}}?$

$\varphi?$



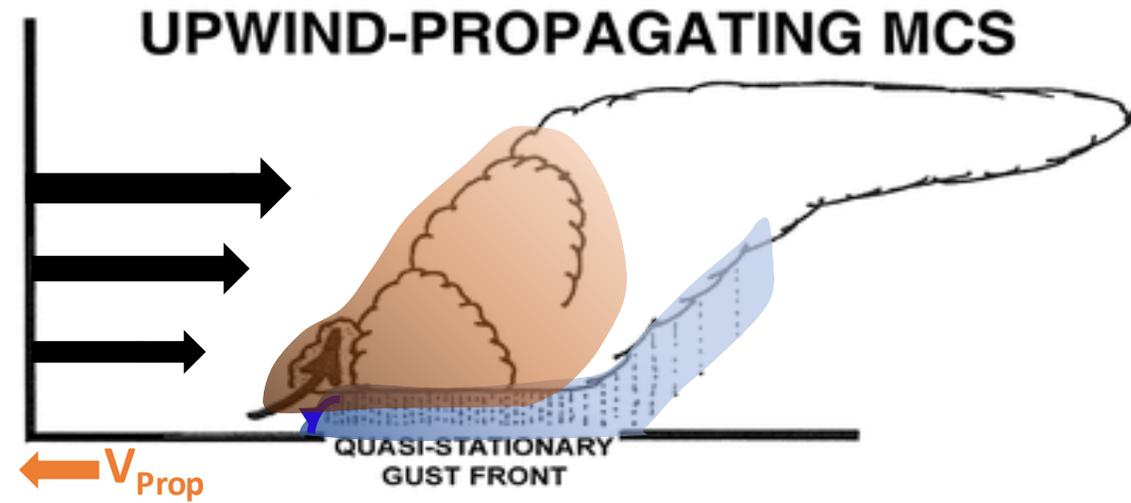
MCS Forward Motion

MCS Forward Motion

- Not a straightforward relationship. MCS can propagate both upwind and downwind of the mean vertical wind field.

MCS Forward Motion

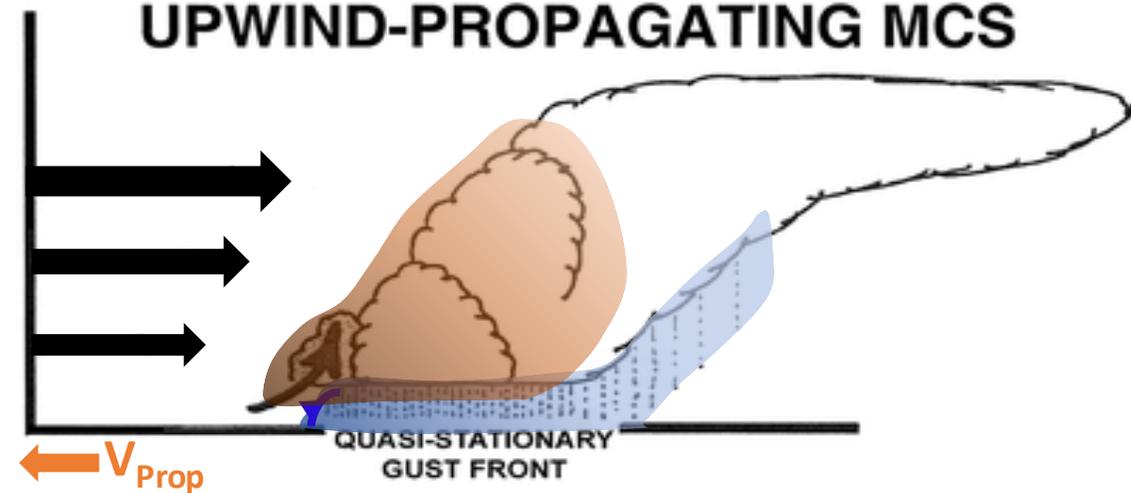
- Not a straightforward relationship. MCS can propagate both upwind and downwind of the mean vertical wind field.



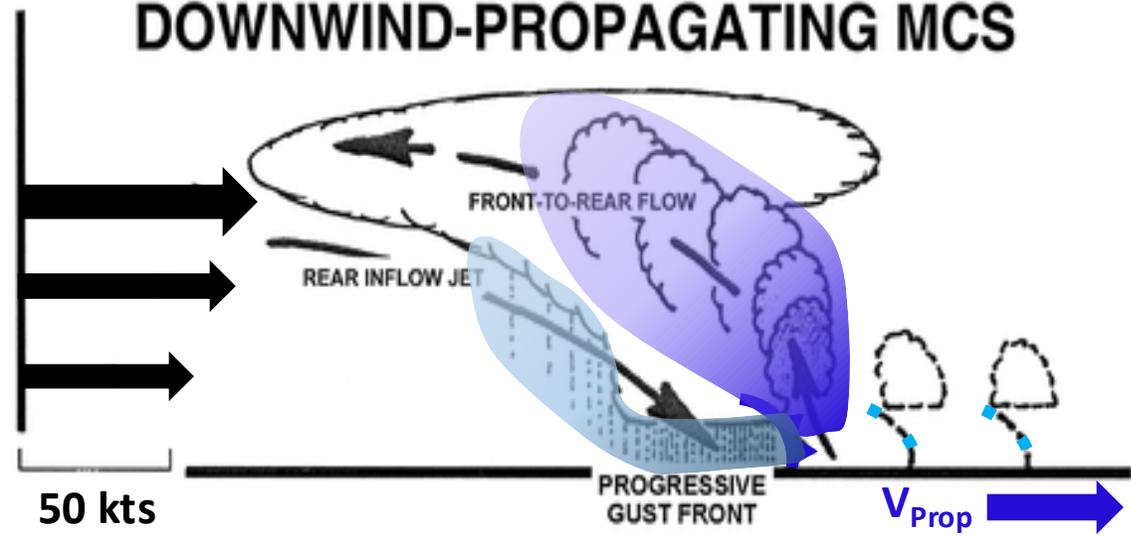
MCS Forward Motion

- Not a straightforward relationship. MCS can propagate both upwind and downwind of the mean vertical wind field.

UPWIND-PROPAGATING MCS



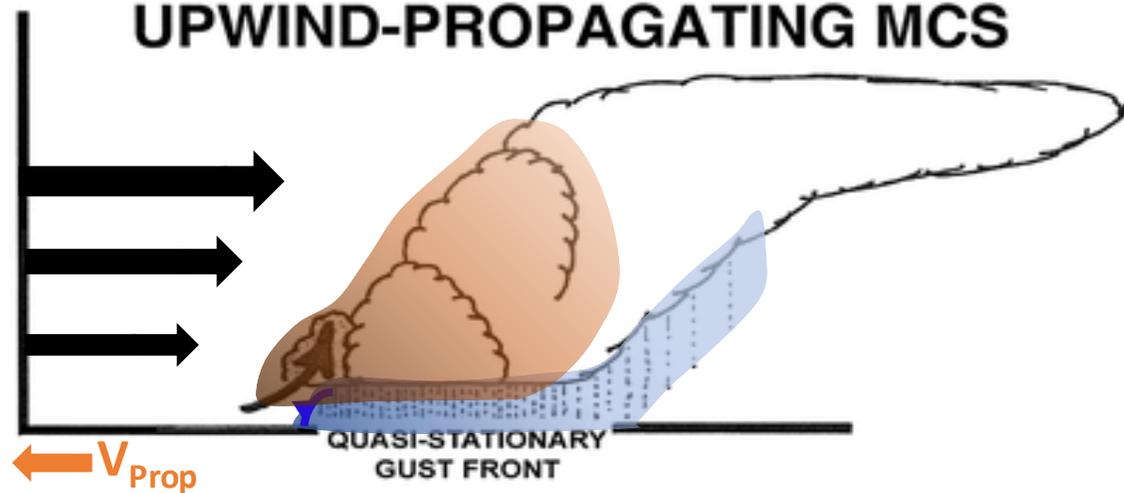
DOWNWIND-PROPAGATING MCS



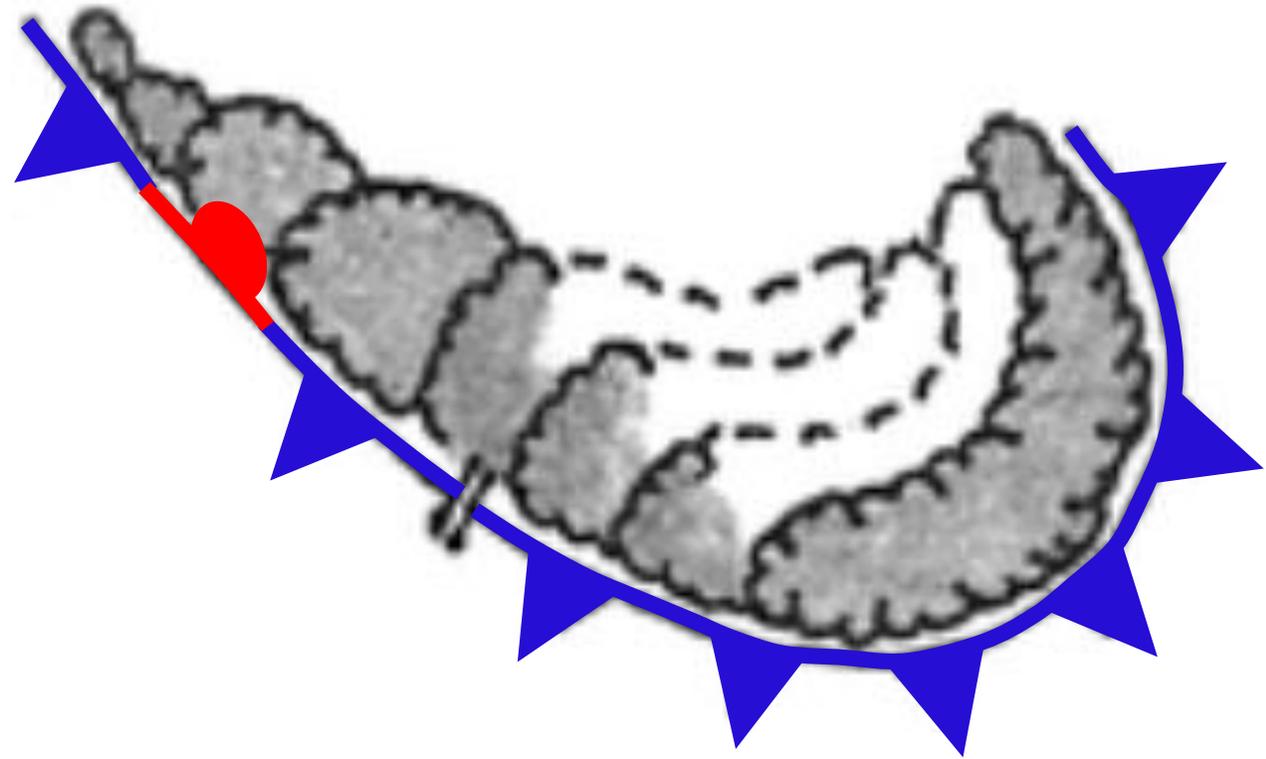
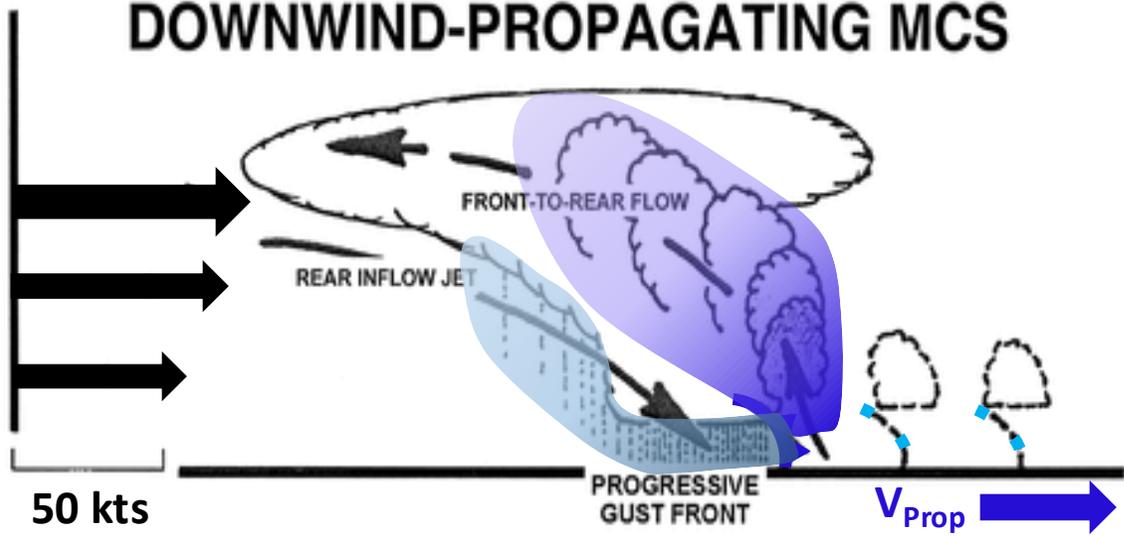
MCS Forward Motion

- Not a straightforward relationship. MCS can propagate both upwind and downwind of the mean vertical wind field.

UPWIND-PROPAGATING MCS

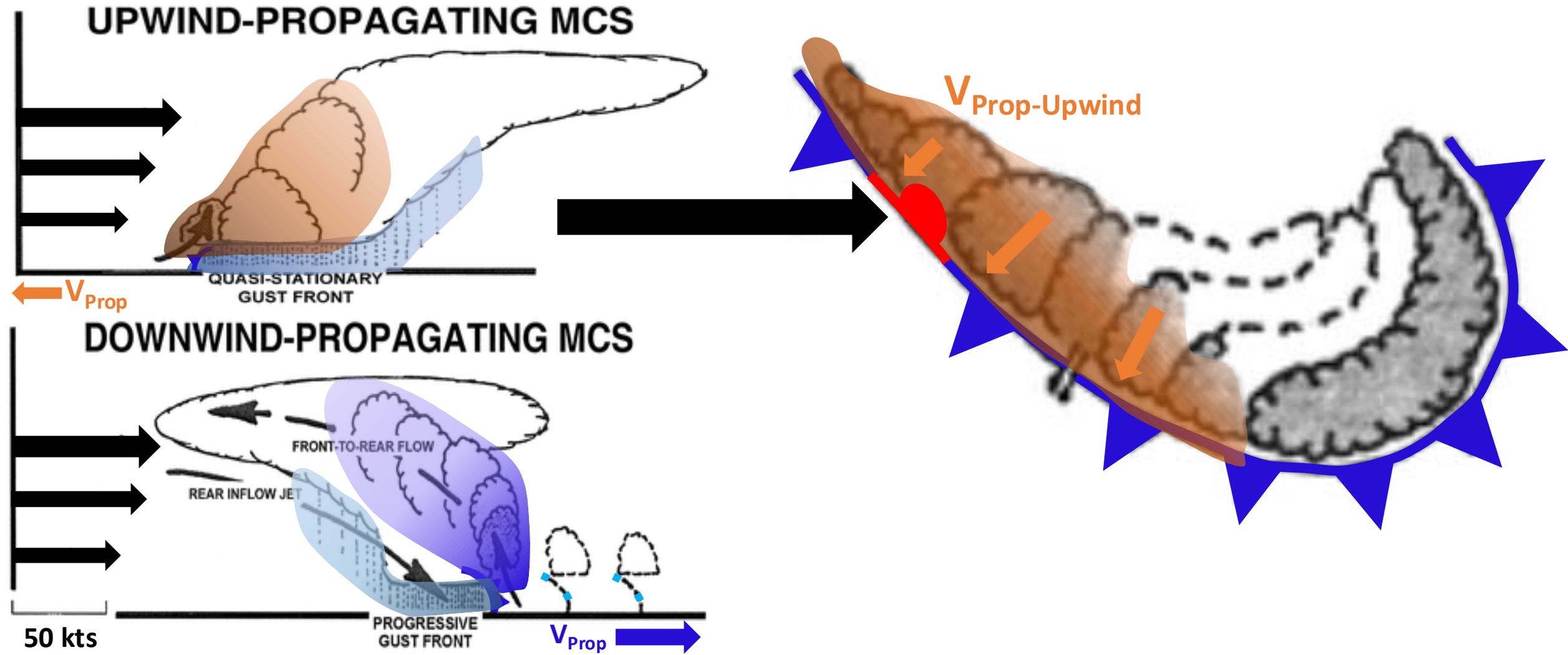


DOWNWIND-PROPAGATING MCS



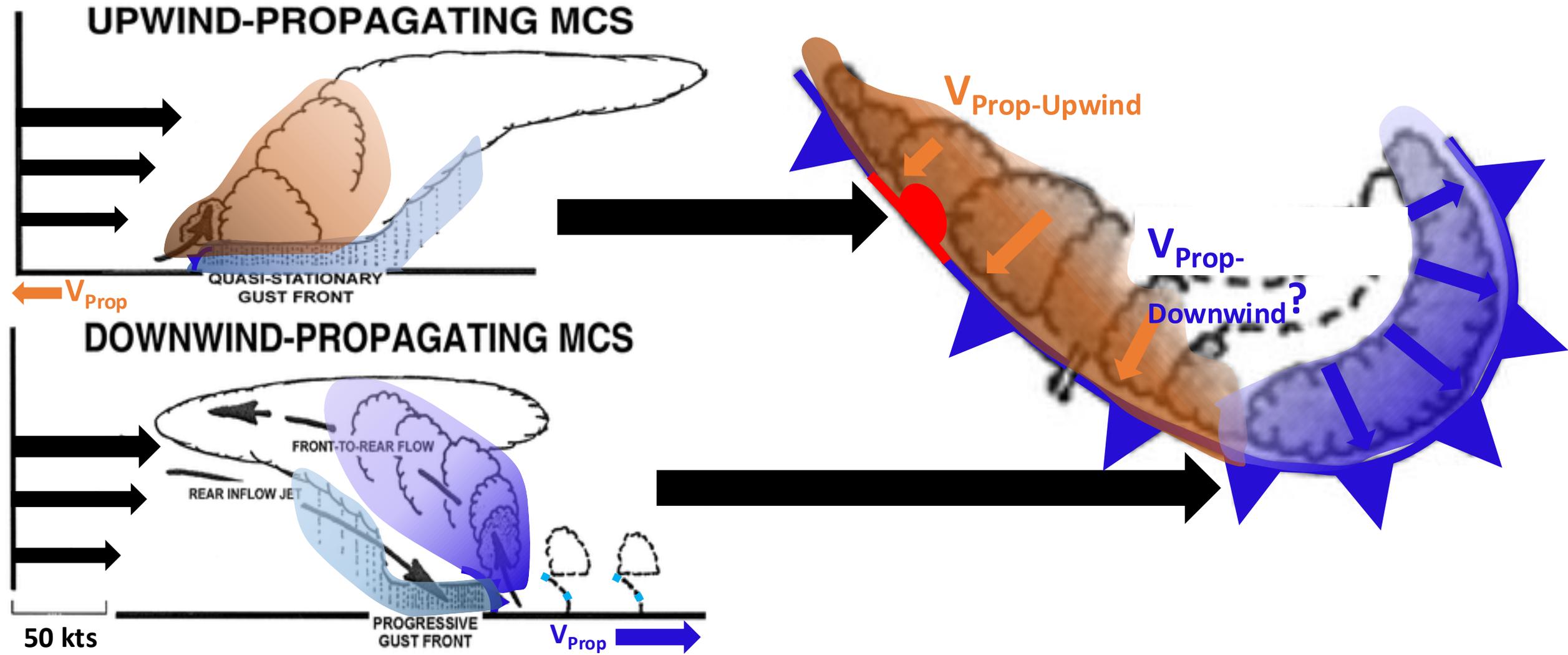
MCS Forward Motion

- Not a straightforward relationship. MCS can propagate both upwind and downwind of the mean vertical wind field.



MCS Forward Motion

- Not a straightforward relationship. MCS can propagate both upwind and downwind of the mean vertical wind field.
- Corfidi et al. (1996) demonstrated how to account for upwind propagation. What about downwind propagation?



MCS Forward Motion Factors

MCS Forward Motion Factors

- The Corfidi et al. (1996) method only factors mean-wind and low-level wind/convergence-driven propagation influences.

MCS Forward Motion Factors

- The Corfidi et al. (1996) method only factors mean-wind and low-level wind/convergence-driven propagation influences.
- What about the cold pool? Studies have shown that MCS forward speed is dependent on cold pool evolution (Charba 1974; Newton and Fankhauser 1975; Betts 1976; Miller and Betts 1977).

MCS Forward Motion Factors

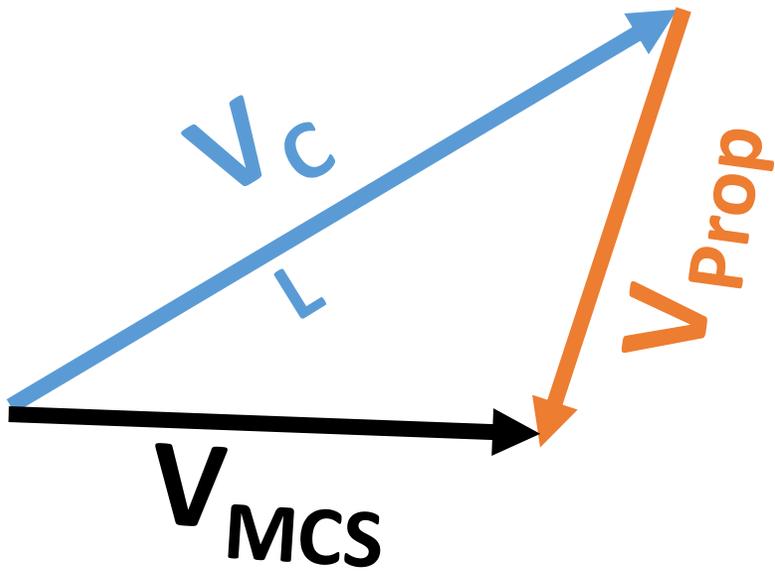
- In reality, MCS forward motion is influenced by 3 factors: Mean Wind Speed - V_{CL} , Upwind (low-level-convergence-driven) propagation - $V_{Prop-Upwind}$, and Downwind (cold-pool-driven) propagation - $V_{Prop-Downwind}$ (Corfidi 2003).
- However, the Corfidi et al. (1996) V_{MCS} vector already takes into upwind propagation, so we can substitute this vector in as a component of MCS motion. Henceforth, we will call the ' V_{MCS} ' vector ' V_{upwind} '.

MCS Forward Motion Factors

- Again, the components of MCS forward motion are additive, so we add V_{CL} and V_{upwind} to get $V_{downwind}$. As such,

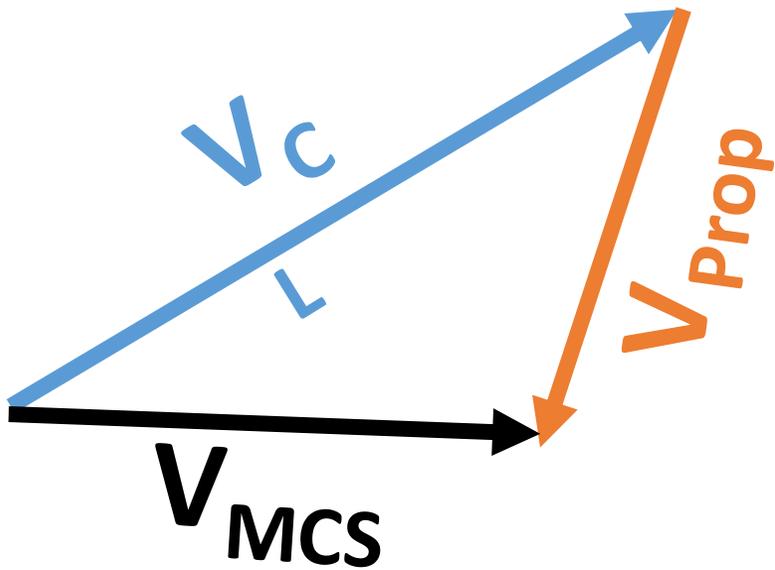
MCS Forward Motion Factors

- Again, the components of MCS forward motion are additive, so we add V_{CL} and V_{upwind} to get $V_{downwind}$. As such,



MCS Forward Motion Factors

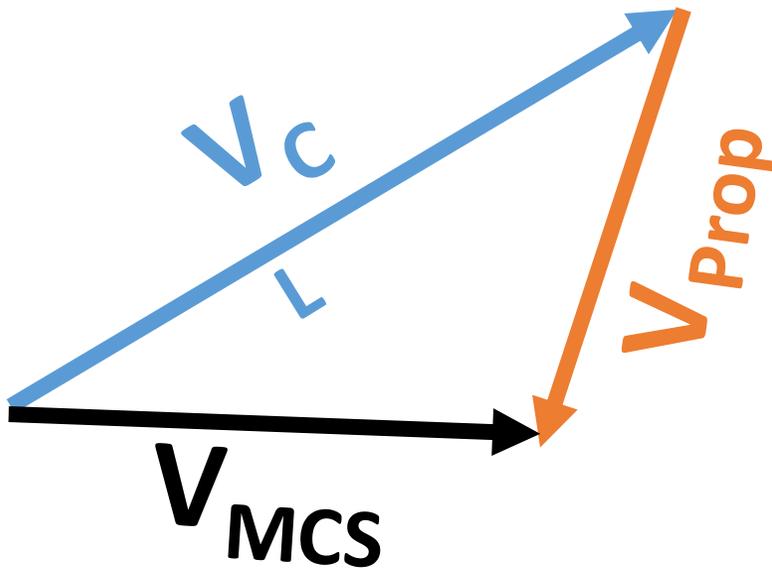
- Again, the components of MCS forward motion are additive, so we add V_{CL} and V_{upwind} to get $V_{downwind}$. As such,



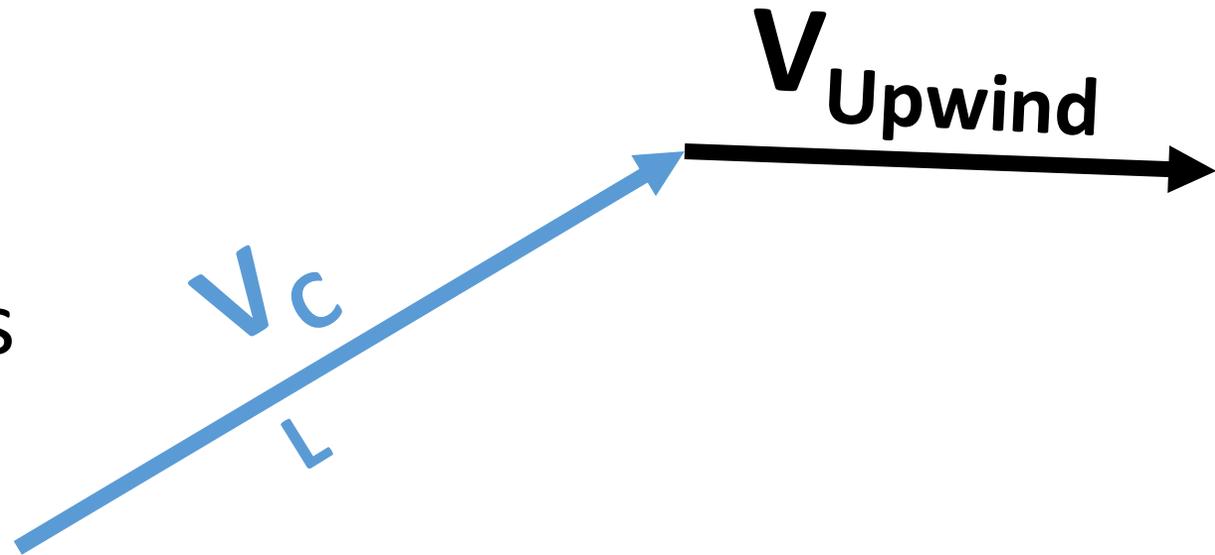
becomes

MCS Forward Motion Factors

- Again, the components of MCS forward motion are additive, so we add V_{CL} and V_{upwind} to get $V_{downwind}$. As such,



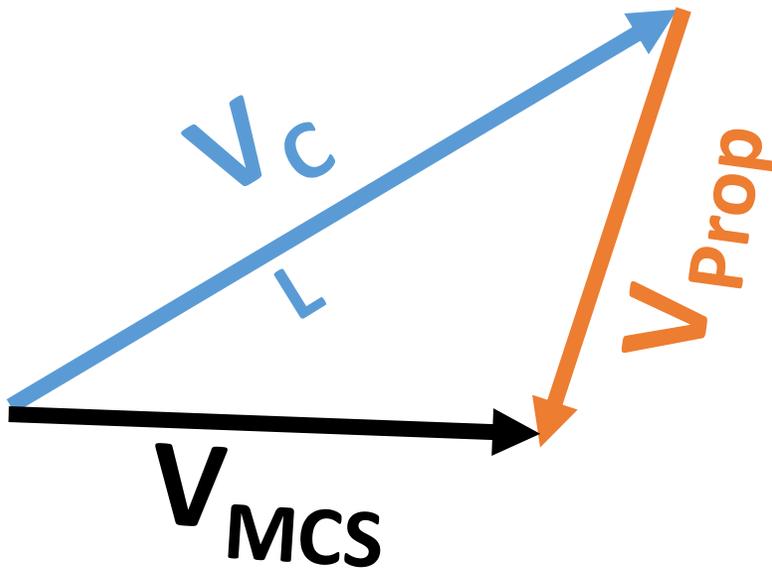
becomes



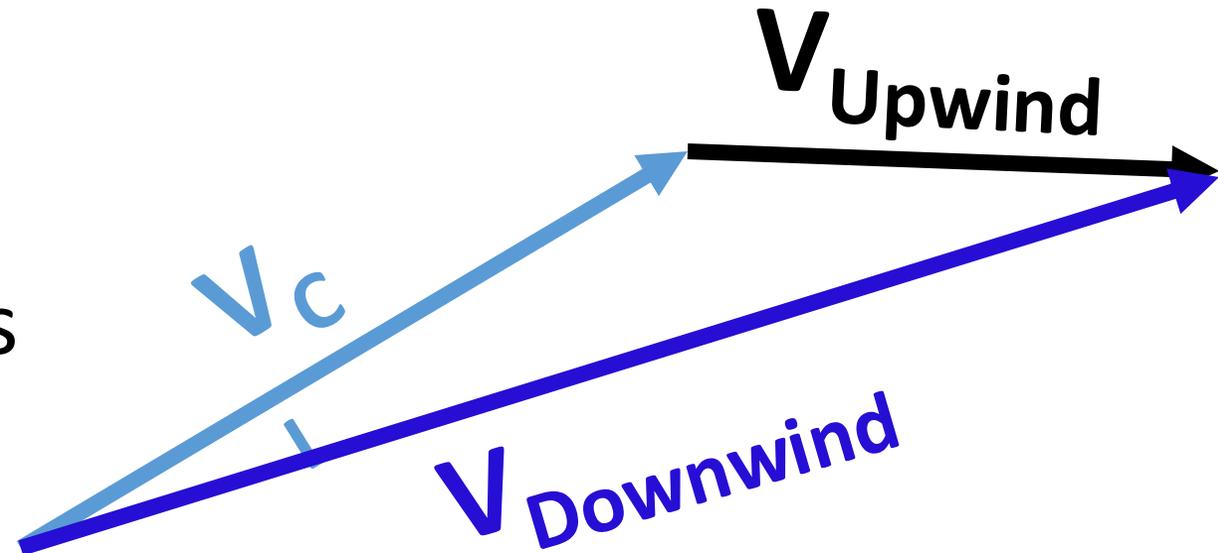
Where $V_{MCS} = V_{Upwind}$

MCS Forward Motion Factors

- Again, the components of MCS forward motion are additive, so we add V_{CL} and V_{upwind} to get $V_{downwind}$. As such,



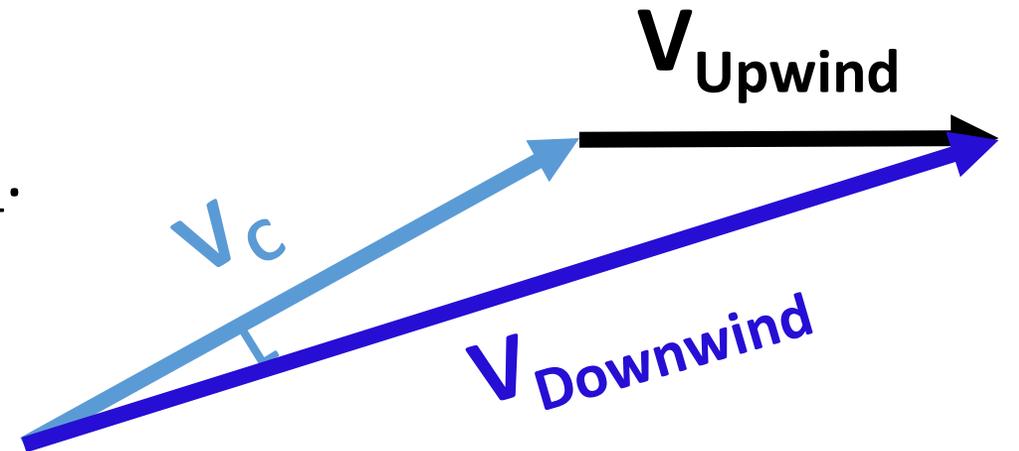
becomes



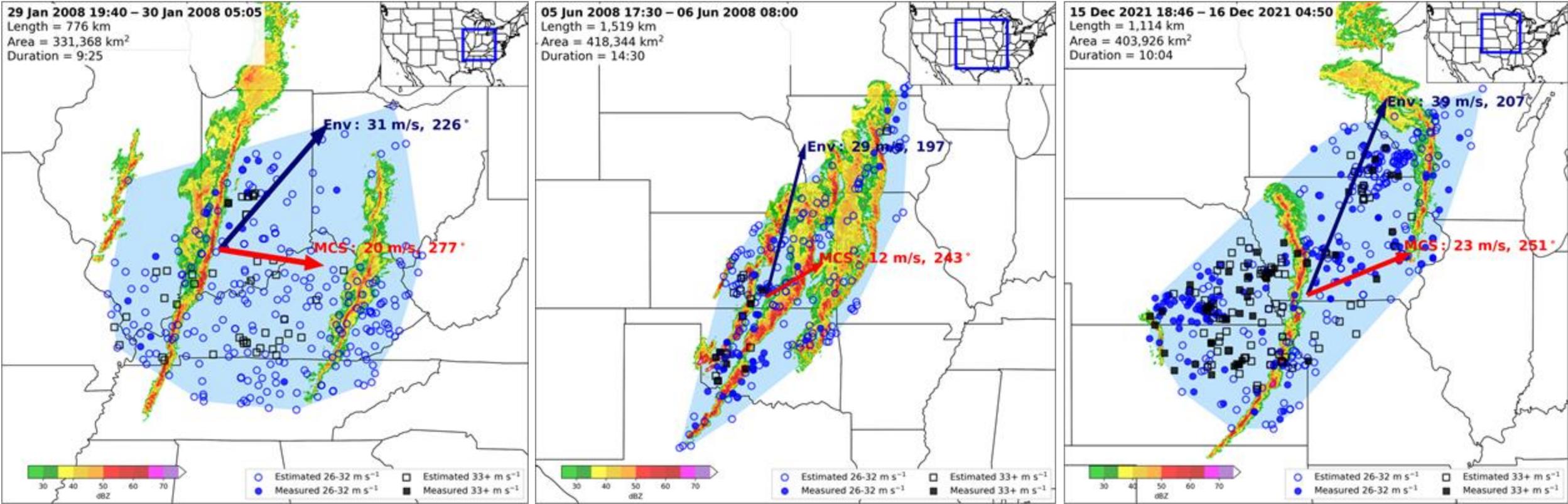
Where $V_{MCS} = V_{Upwind}$

MCS Forward Motion Factors

- Downwind-propagating MCSs are dominated by the cold pool, and derechos are also produced by cold-pool-driven MCSs, which are dominated by internal forcing mechanisms.
- As such, V_{downwind} would be a useful vector for monitoring derecho progression.
- Note that V_{downwind} is a longer vector than V_{CL} .



MCSs moving faster than the mean wind speed is an excellent discriminator between cold-pool-driven MCSs and their squall line counterparts.



Note: Cold-pool-driven MCSs and strongly forced squall lines both have degrees of internal and external forcing (i.e. a level of contribution from the cold pool)

The argument is that to define derechos as a distinct phenomena, internal forcing mechanisms must dominate, which is defined by the MCS moving faster than the full mean wind speed.

References

- Corfidi, S. F., 2003: Cold pools and MCS propagation: Forecasting the motion of downwind-developing MCSs. *Wea. Forecasting*, **18**, 997–1017, [https://doi.org/10.1175/1520-0434\(2003\)018<0997:CPAMPF>2.0.CO;2](https://doi.org/10.1175/1520-0434(2003)018<0997:CPAMPF>2.0.CO;2).
- Corfidi, S. F., Merritt J. H. , and Fritsch J. M. , 1996: Predicting the movement of mesoscale convective complexes. *Wea. Forecasting*, **11** , 41–46. [https://doi.org/10.1175/1520-0434\(1996\)011<0041:PTMOMC>2.0.CO;2](https://doi.org/10.1175/1520-0434(1996)011<0041:PTMOMC>2.0.CO;2).
- Houze, R. A., 2018: 100 years of research on mesoscale convective systems. *A Century of Progress in Atmospheric and Related Sciences: Celebrating the American Meteorological Society Centennial*, *Meteor. Monogr.*, No. 59, Amer. Meteor. Soc., <https://doi.org/10.1175/AMSMONOGRAPHS-D-18-0001.1>.
- Parker, M. D., and R. H. Johnson, 2000: Organizational modes of midlatitude mesoscale convective systems. *Mon. Wea. Rev.*, **128**, 3413–3436, [https://doi.org/10.1175/1520-0493\(2001\)129<3413:OMOMMC>2.0.CO;2](https://doi.org/10.1175/1520-0493(2001)129<3413:OMOMMC>2.0.CO;2).